

MEASUREMENT OF π^+ – ARGON ABSORPTION AND CHARGE EXCHANGE
INTERACTIONS USING PROTODUNE-SP

By

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ABSTRACT

MEASUREMENT OF π^+ – ARGON ABSORPTION AND CHARGE EXCHANGE INTERACTIONS USING PROTODUNE-SP

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ProtoDUNE-SP is a prototype detector for the upcoming Deep Underground Neutrino Experiment (DUNE). It's a Liquid Argon Time Projection Chamber (LArTPC) with a similar configuration to DUNE's detector, and is designed to provide a test-bed for the future experiment. In addition to serving as a prototype, its 0.3 – 7 GeV/ c charged particle beam line provided the ability to perform physics measurements of pions, protons, kaons, muons, and electrons. Importantly, the LArTPC allowed for the measurement of hadronic interactions on argon nuclei.

Pions are often present in the final state of neutrino interactions in the energy range of DUNE's neutrino beam. These particles can undergo various types of interactions with argon nuclei in the detector, and this can interfere with the characterization of neutrino interactions in DUNE's far detector. The rate of these so-called secondary interactions will be accounted for using Monte Carlo simulation of neutrino interactions. Measurements of secondary interaction rates provide necessary data which can be used to estimate and propagate uncertainties or provide tunes of the secondary interaction model used within DUNE's experimental simulation.

This analysis provides a simultaneous measurement of the π^+ –Ar absorption and charge exchange cross sections using 1 GeV/ c π^+ data taken by ProtoDUNE-SP during its initial run period in Fall 2018. This is one of the first hadronic interaction measurements provided by ProtoDUNE-SP. It is also the first π^+ –Ar absorption measurement in 20 years and the first ever π^+ –Ar charge exchange measurement.

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CHAPTER 1

EXECUTIVE SUMMARY

1

2 The field of study of neutrino oscillation has entered the precision era. Next-generation
3 experiments – the Deep Underground Neutrino Experiment (DUNE) and Hyper-Kamiokande
4 (Hyper-K) – will collect a large rate of accelerator-based neutrino-interaction events. This
5 will provide researchers with the ability to answer remaining key questions within oscillation
6 physics. DUNE, the physics program on which this thesis will focus, will attempt to answer
7 the following questions:

- 8 1. What are the precise values of the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix
9 mixing angles (θ_{12} , θ_{23} , θ_{13})? Specifically, is θ_{23} lower than, greater than, or even
10 equal to 45° (known as maximal mixing)?
- 11 2. Does neutrino oscillation violate Charge-Parity symmetry (is δ_{CP} of the PMNS matrix
12 non-zero)?
- 13 3. What is the ordering of the neutrino masses (what is the sign of Δm_{31}^2)?

14 Great effort must be taken to reduce systematic uncertainty to a suitable level to achieve
15 precise measurements related to the questions stated above. Necessary for this is the proto-
16 typing of DUNE’s far detector with ProtoDUNE-SP. ProtoDUNE-SP serves as a test-bed for
17 DUNE’s detector components and event reconstruction, a first-attempt at calibration which
18 will be employed at DUNE, and a source of physics measurements using its charged particle
19 beam line which will serve as necessary inputs to DUNE’s simulation. One of the particles
20 provided by the beam line – π^+ – is important to study, as it is often found in the final state
21 of neutrino interactions. As such, it has the ability to interfere with the reconstruction of
22 the incident neutrino’s energy or its flavor. For example, if a π^+ is absorbed by an argon
23 nucleus nearby the primary neutrino interaction, the pion’s energy could be missed in recon-

24 struction of the neutrino's energy. Additionally, if a π^+ is produced in a neutral current ν_μ
25 interaction it could instead undergo a charge exchange interaction, where it is converted into
26 a π^0 , nearby. The π^0 will promptly decay into two photons, which will produce showers in
27 the detector. These showers could mimic an electron shower and could cause the ν_μ neutral
28 current interaction to be misidentified as a ν_e charged current interaction. These errors will
29 be accounted for in DUNE's oscillation analyses using Monte Carlo simulation of events in
30 DUNE's far detector. However, if the rate of pion interactions are misestimated, DUNE's
31 measurements could be biased. This thesis presents a measurement of the π^+ -Ar absorption
32 and charge exchange in order to reduce these systematic effects in DUNE's analyses.

33 This thesis will be organized as follows. Chapter 2 will describe the Standard Model of
34 particle physics and how neutrinos fit (or rather *do not fit*) within this theory. It will also
35 provide an overview of interactions of both neutrinos and charged pions on nucleons and
36 nuclei. Chapter 3 will describe DUNE's physics program and detector design and provide
37 motivation for this measurement. In Chapter 4, the ProtoDUNE-SP detector – the detector
38 used for this measurement – will be discussed. Specifically, this chapter will focus on the
39 design and components of the detector, the software used for reconstructing particle trajec-
40 tories and interactions, and calibration of the detector. Chapter 5 will discuss the beam
41 line which supplies the ProtoDUNE-SP detector with its test beam particles. Chapter 6 will
42 describe the event selection used to characterize Monte Carlo events and data sets. Chapter
43 7 will discuss the strategy used to conduct this measurement including the strategy used to
44 extract the cross section and the statistical fit used in the analysis. Chapter 8 will discuss
45 the systematic uncertainties within the analysis. Chapter 9 will discuss validations of the
46 statistical fit using fake data generated from Monte Carlo simulation. Finally, Chapter 10
47 will present the results of the measurement on real data.

CHAPTER 2

48

THEORY

49 As DUNE’s physics program centers around neutrinos, this chapter will provide a description
50 of our current understanding of these particles as part of the Standard Model of Particle
51 Physics. Within this, an overview of the theory of neutrino–nucleus interactions will be
52 given. The physics of charged pions will also be discussed, as these are often produced
53 within neutrino–nucleus interactions.

54 **2.1 The Standard Model**

55 The Standard Model of Particle Physics represents the most up-to-date understanding
56 of the universe at the subatomic level. It has provided immensely accurate descriptions of
57 particle interactions (manifested as the electromagnetic and strong and weak nuclear forces)
58 and successfully predicted the presence of multiple elementary particles. The Standard Model
59 is rooted in the local symmetry group

$$\text{SU}(3) \times \text{SU}(2)_L \times \text{U}(1) \tag{2.1}$$

60 where the first term encompasses the strong interaction and the second and third terms
61 give rise to the electroweak interaction. Here, the L subscript denotes this describes a “left-
62 handed” chiral theory.

63 From the symmetry groups denoted in Equation 2.1, the interactions between matter and
64 forces arise. In the development of the theory, a Lagrangian is constructed which describes
65 a free fermion field and invariance under some local gauge transformation is enforced. If
66 the fermion field is not invariant under that gauge transformation, an interaction with some
67 vector field is introduced. Depending on the field and the gauge under consideration, these
68 vector fields may also interact amongst themselves to ensure gauge invariance. The quanta of
69 these vector fields are known as gauge bosons, and are modeled as being exchanged between

70 interacting fermions. At this point, all gauge bosons and fermions have zero mass. However,
 71 this is wrong as known from experiment: three of the bosons and all fermions have mass.
 72 With the possible exception of neutrinos, these particles all gain mass due to the presence of
 73 the so-called Higgs field (the quantum of which is the scalar Higgs Boson, famously discovered
 74 in 2012 [1][2]). The bosons gain mass as the result of spontaneous symmetry breaking, and
 75 the fermions gain mass through coupling to the Higgs fields via Yukawa interactions [3].
 76 The following sections describe the gauge bosons and the elementary fermions (quarks and
 77 leptons), as well as composite particles formed from quarks (hadrons).

78 2.1.1 Gauge Bosons

79 The most familiar of the gauge bosons is the photon, which is a massless, neutral particle
 80 that couples to electric charge and thus mediates the electromagnetic force. Figure 2.1a
 81 shows a Feynman diagram representing an elementary electromagnetic interaction vertex,
 82 where f is some charged fermion and the boson γ is the photon in the interaction.

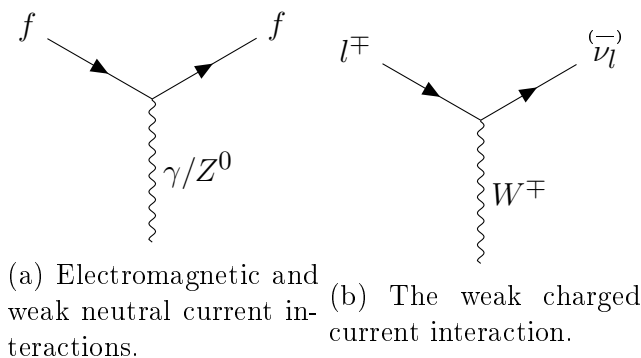


Figure 2.1: Elementary interaction vertices of the electroweak interactions.

83 The photon arises within the Standard Model as one of the vector fields necessary to
 84 achieve $SU(2) \times U(1)$ gauge invariance (which describes the electroweak interaction). The
 85 other fields introduced by requiring this invariance are the Z^0 and W^\pm bosons which mediate
 86 the neutral and charged current weak interactions respectively. Elementary weak interactions
 87 are also shown in Figure 2.1, where the neutral current interaction is represented in Figure

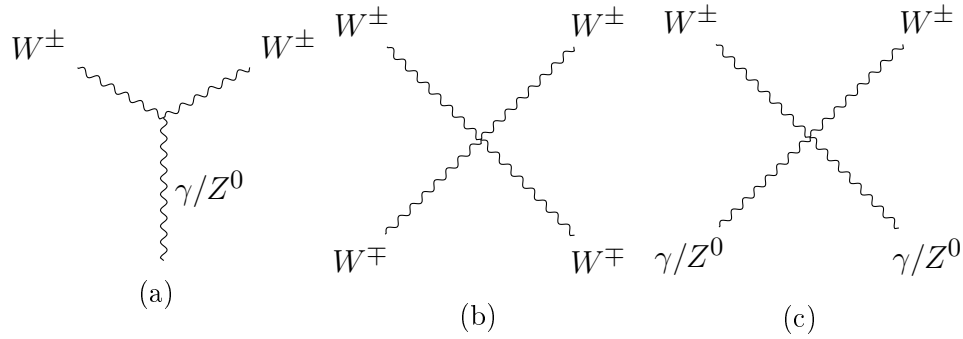


Figure 2.2: Electroweak self-interaction vertices.

88 2.1a with the Z^0 as the boson in the interaction and f is some fermion, and the charged
 89 current interaction is in Figure 2.1b. The W^\pm and Z^0 bosons are the three gauge bosons
 90 mentioned above which gain their mass through the Higgs mechanism. Within electroweak
 91 theory, interactions between the W , Z , and photon also occur. These elementary interaction
 92 vertices are shown in Figure 2.2.

93 Finally, SU(3) invariance introduces eight gluons to facilitate the strong nuclear force.
 94 This force couples to a property known as color, which, like electric charge for the elec-
 95 tromagnetic interaction, is common to all particles that experience the strong interaction
 96 (quarks and gluons themselves). Color differs from electric charge in that there are 3 colors
 97 (red, green, and blue) plus 3 anti-colors (anti-red, anti-green, and anti-blue) rather than just
 98 positive or negative electric charge. The gluons themselves carry color, and thus, due to
 99 color conservation, annihilate quarks of one color and create quarks of another color during
 100 interactions. In Figure 2.3a, the quarks entering and exiting the vertex are implied to have
 101 different colors. Similar to the electroweak bosons, interactions between the gluons arise as
 102 part of SU(3) invariance, giving rise to vertices such as Figures 2.3b and 2.3c.

103 2.1.2 Quarks

104 The elementary components of matter are a group of 12 fermions (and their antiparticle
 105 partners) separated into 6 quarks and 6 leptons. The quarks engage in all three forces

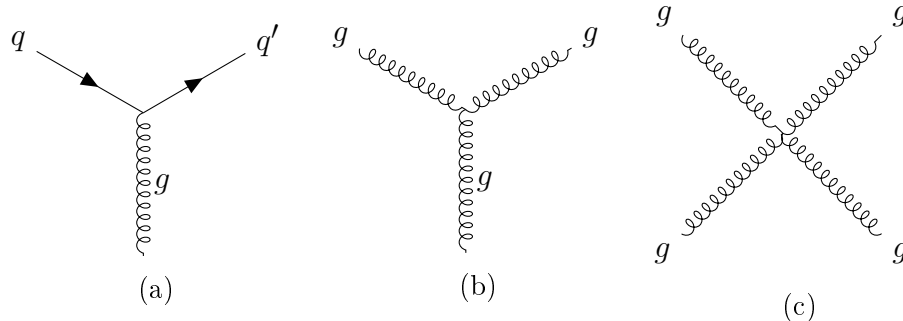


Figure 2.3: Interaction vertices of the strong force.

106 described previously, while the leptons feel only the electromagnetic and weak forces (and
 107 within the leptons, the neutrinos only engage in the weak forces as they are neutral).

108 The quarks come in six flavors: up, down, charm, strange, top, and bottom, which are
 109 separated into three generations as shown in Table 2.1. Each generation contains one quark
 110 with electric charge equal to $\frac{2}{3}e$ and the other with charge of $-\frac{1}{3}e$ (where e is the basic unit
 111 of electric charge). The first group includes u , c , and t quarks and are collectively known
 112 as “up-type” quarks, while the second group includes d , s , and b quarks and are known as
 113 the “down-type” quarks. As mentioned earlier, the quarks carry a property known as color
 114 which is similar to electric charge. In an analogy to primary colors of visible light, this
 115 comes in the form of three charges – (anti-)red, (anti-)green, or (anti-)blue. The analogy to
 116 the primary colors of light comes from the fact that the three color charges (or a color and
 117 its anticolor) add together to a net-0 or “white” color charge. A special property of color –
 118 color confinement – is that no isolated free particle can exist in a colored state. This leads
 119 to quarks forming bound states known as hadrons. Most often, these come in the form of
 120 mesons (bound states of a quark and an antiquark with one type of color and its anticolor)
 121 and baryons (bound states of three quarks each with one of the colors). Baryons and mesons
 122 will be discussed further in the next section.

123 In addition to the strong interaction, the quarks also take part in the weak interaction.
 124 In charged current weak interaction, the quarks transition between up-type and down-type
 125 flavors. Within the Standard Model Lagrangian, the quarks are described as two-component

Generation	Quark	Charge (e)	Mass (MeV/c ²)
I	u	2/3	2.4
	d	-1/3	4.8
II	c	2/3	1270
	s	-1/3	104
III	t	2/3	1.712 x 10 ⁵
	b	-1/3	4200

TABLE 2.1: The quark generations along with the charge and mass of the individual quarks.

126 states, comprised of the up-type and down-type of each generation, as shown in Equation
127 2.2, which are operated upon in the weak interaction.

$$\Psi^q = \begin{pmatrix} \psi^u \\ \psi^d \end{pmatrix}, \begin{pmatrix} \psi^c \\ \psi^s \end{pmatrix}, \begin{pmatrix} \psi^t \\ \psi^b \end{pmatrix} \quad (2.2)$$

128 If these were eigenstates of the weak interaction, one would expect this process to transition
129 between quarks only *within* generations (i.e. u would only transition to d , c to s , and t to b).
130 This is not the case, and cross-generation transitions are allowed. This can be represented by a
131 set of quark states (d' , s' , b') that are linear combinations of the normal down-type quarks.
132 The Cabibbo-Kobayashi-Maskawa (CKM) matrix mixes the normal down-type quarks into
133 the special weak-eigenstate quarks as seen in Equation 2.3.

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} U_{ud} & U_{us} & U_{ub} \\ U_{cd} & U_{cs} & U_{cb} \\ U_{td} & U_{ts} & U_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix} \quad (2.3)$$

134 2.1.3 Hadrons

135 As stated above, quarks cannot be observed in isolated states¹, and reside in composite parti-
136 cles known as hadrons. Most commonly, these are combinations of a quark and an antiquark

¹That is, at energy scales relevant to this thesis. At high enough energies, a phase transition to a state known the quark-gluon-plasma occurs wherein quarks and gluons are not confined to hadrons [4]

Generation	Lepton	Charge (e)	Mass (MeV/c ²)
I	e	-1	.511
	ν_e	0	–
II	μ	-1	105.7
	ν_μ	0	–
III	τ	-1	1777
	ν_τ	0	–

TABLE 2.2: The lepton generations along with the charge and mass of the individual leptons. Note that the neutrinos are assumed to be massless within the Standard Model and so no masses are stated for these particles here.

137 (mesons) or combinations of three quarks/antiquarks (baryons). More exotic combinations
138 such as tetraquarks (two quarks and two antiquarks) and pentaquarks (four quarks and one
139 antiquark) have recently been discovered at the Large Hadron Collider [5].

140 The baryons include familiar particles like protons and neutrons (made of uud quarks and
141 udd quarks respectively); higher energy resonances of the same sets of quarks such as Δ^+
142 and Δ^0 ; and particles including second or third generation quarks such as Λ^0 , Λ_c^+ , and Λ_b^0
143 (uds , udc , udb respectively). Many other combinations of quark flavors exist, as do similar
144 combinations of quark flavors but with differing quantum numbers. For example, the Δ^+
145 and proton have the same flavors of quarks, but have total angular momentum of 3/2 and
146 1/2 respectively.

147 The mesons are similarly characterized by properties of their constituent quarks (flavor,
148 angular momentum, etc.). The mesons include the charged and neutral pions (whose quark
149 content is $u\bar{d}$, $d\bar{u}$, and $(u\bar{u} - d\bar{d})/\sqrt{2}$ for π^+ , π^- , π^0 respectively), kaons (which include a
150 strange quark), and various other combinations of quark flavors and angular momenta.

151 The pions are of particular interest within neutrino physics. They play an important role
152 in nuclear dynamics, as they are the long-range mediator of the nuclear force according to
153 Yukawa theory [4], and are discussed further in Section 2.4.

154 **2.1.4 Leptons**

155 Similar to the quarks, the 6 leptons are separated into 3 generations, each containing 1
156 charged lepton and its neutral partner. The charged leptons are the electron (e), muon (μ),
157 and tau (τ), which each have a charge of $-1e$ and masses as seen in Table 2.2. In each
158 generation is also the neutral partner to the charged lepton: the electron neutrino (ν_e),
159 muon neutrino (ν_μ), and tau neutrino (ν_τ). The leptons are all colorless particles, and thus
160 do not feel the strong force. However, all left-handed leptons and right-handed antileptons
161 engage in the weak interaction. To represent this, the leptons are given Lepton numbers:
162 L_e , L_μ , and L_τ , which are equal to 1 (-1) for (anti-)leptons in the generation denoted by the
163 subscript. These lepton numbers are absolutely conserved in the weak interaction.

164 **2.2 Neutrinos: Not-so-standard Particles**

165 **2.2.1 Neutrino Oscillations**

166 This overview of the Standard Model particles seems tidy, but there are some subtle peculiar-
167 ities, especially with the neutrinos. This is hinted at in Table 2.2, where no masses are stated
168 for the neutrinos. Within the Standard Model, the neutrinos are predicted to have no mass.
169 However, it is now understood that at least two neutrinos in fact *do* have mass, as indicated
170 by the presence of the process known as *neutrino oscillation*. This process was first theo-
171 rized by Bruno Pontecorvo in an attempt to explain a deficit of observed electron neutrinos
172 produced from nuclear reactions in the sun. In 1968, the Davis experiment [6] measured only
173 about 1/3 of expected solar electron neutrinos. Pontecorvo suggested this deficit could be
174 explained if electron neutrinos produced in the Sun transformed into muon or tau neutrinos
175 (which the Davis experiment was unable to detect) before reaching Earth [7]. Takaaki Kajita
176 from Super Kamiokande and Arthur B. McDonald from the Sudbury Neutrino Observatory
177 were awarded the 2015 Nobel Prize in Physics for the discovery of oscillations [8][9].

178 **2.2.2 Neutrino Mixing**

179 Neutrino oscillations arise from the facts that a) neutrinos have nonzero mass and b) the
 180 flavor eigenstates are not equivalent to the mass-energy eigenstates. Similar to quark mix-
 181 ing and the CKM matrix, neutrino mixing is described by the unitary Pontecorvo-Maki-
 182 Nakagawa-Sakata (PMNS) matrix as shown in Equation 2.4, where $\nu_{1,2,3}$ are the mass-energy
 183 eigenstates.

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{PMNS} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \quad (2.4)$$

184 When neutrinos are produced via the weak interaction, they are produced as definite
 185 flavor states. However, the propagation of the neutrinos is described by the time-evolution
 186 operator (equivalently the Hamiltonian operator). As such, the neutrinos travel as eigen-
 187 states of the Hamiltonian: the mass-energy states.

188 **2.2.3 Oscillation Probability**

189 The derivation of the oscillation probability is shown in the following example. Consider
 190 a neutrino that evolves in time, $|\nu(t)\rangle$. Suppose it begins as a muon neutrino, such that
 191 $|\nu(0)\rangle = |\nu_\mu\rangle$. In a vacuum, its evolution is described by the time-evolution operator (e^{-iHt})
 192 as such:

$$|\nu(t)\rangle = e^{-iHt} |\nu(0)\rangle = e^{-iHt} |\nu_\mu\rangle \quad (2.5)$$

Because $|\nu_\mu\rangle$ is not an eigenstate of the Hamiltonian, this is expanded to the following:

$$\begin{aligned} |\nu(t)\rangle &= e^{-iHt} (U_{\mu1} |\nu_1\rangle + U_{\mu2} |\nu_2\rangle + U_{\mu3} |\nu_3\rangle) \\ &= U_{\mu1} e^{-iE_1 t} |\nu_1\rangle + U_{\mu2} e^{-iE_2 t} |\nu_2\rangle + U_{\mu3} e^{-iE_3 t} |\nu_3\rangle \end{aligned}$$

193 Suppose an experiment is attempting to measure the rate at which ν_μ oscillate to ν_e by
 194 detecting the ν_e ². The probability to detect the neutrino as ν_e at some point in time is
 195 related to the following matrix element

$$\langle \nu_e | \nu(t) \rangle = \sum_j U_{ej}^* U_{\mu j} e^{-iE_j t} \quad (2.6)$$

196 where the sum over j runs over the mass-energy states. The probability is then given by

$$P(\nu_\mu \rightarrow \nu_e) = |\langle \nu_e | \nu(t) \rangle|^2 = \sum_{j,k} U_{ej}^* U_{\mu j} U_{ek} U_{\mu k}^* e^{-i(E_j - E_k)t} \quad (2.7)$$

197 For a general pair of states a, b this probability is

$$P(\nu_a \rightarrow \nu_b) = \sum_{j,k} U_{bj}^* U_{aj} U_{bk} U_{ak}^* e^{-i(E_j - E_k)t} \quad (2.8)$$

Assuming the neutrino is ultrarelativistic, the energy can be expanded as such:

$$\begin{aligned} E_j &= \sqrt{|\vec{p}|^2 + m_j^2} = |\vec{p}| \sqrt{1 + \frac{m_j^2}{|\vec{p}|^2}} \\ &\approx |\vec{p}| \left(1 + \frac{m_j^2}{2|\vec{p}|^2} \right) = |\vec{p}| + \frac{m_j^2}{2|\vec{p}|} \\ &\approx E + \frac{m_j^2}{2E} \end{aligned}$$

198 Where the approximation $E \approx |\vec{p}|$ was used in the final step. The difference in the exponential
 199 term becomes

$$E_j - E_k \approx \frac{\Delta m_{jk}^2}{2E} \quad (2.9)$$

200 where $\Delta m_{jk}^2 = m_j^2 - m_k^2$. There are three mass splittings, Δm_{21}^2 , Δm_{32}^2 , and Δm_{31}^2 , two of
 201 which are independent³. Δm_{21}^2 is known to be positive (this is discussed later), while the
 202 sign of Δm_{32}^2 – also known as the *neutrino mass hierarchy* – is an open question in neutrino
 203 physics. This is highlighted in Figure 2.4, which shows the two possible orderings of the
 204 neutrino mass states.

²This is commonly known as an *electron neutrino appearance* analysis.

³ $\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$.

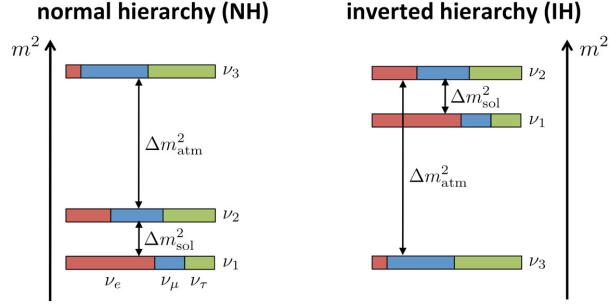


Figure 2.4: Representation of the two possible neutrino mass hierarchies. The left is the normal hierarchy with $m_3 > m_2 > m_1$. The right is the inverted hierarchy with $m_2 > m_1 > m_3$ [10].

The probability stated in 2.8 is often rewritten by splitting the real and imaginary components of the unitary PMNS matrix and approximating $t \approx L$ where L is the distance the neutrino has traveled:

$$\begin{aligned}
 P(\nu_a \rightarrow \nu_b) &= \delta_{ab} \\
 &- 4 \sum_{j>k} \Re \left[U_{bj}^* U_{aj} U_{bk} U_{ak}^* \right] \sin^2 \left(\frac{\Delta m_{jk}^2 L}{4E} \right) \\
 &\pm 2 \sum_{j>k} \Im \left[U_{bj}^* U_{aj} U_{bk} U_{ak}^* \right] \sin \left(\frac{\Delta m_{jk}^2 L}{2E} \right)
 \end{aligned} \tag{2.10}$$

205 Here, the third term is positive (negative) for (anti-)neutrinos.

206 The PMNS matrix is parameterized by 3 mixing angles – θ_{12} , θ_{13} , and θ_{23} – and a phase
 207 factor – δ_{CP} – and is commonly factored into a product of three rotation matrices as such:

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \tag{2.11}$$

208 where c_{ij} and s_{ij} are $\cos \theta_{ij}$ and $\sin \theta_{ij}$. δ_{CP} is a phase factor that determines whether
 209 neutrino oscillations violate charge-parity (CP) symmetry. This symmetry and its violation

210 are described in the next subsection, and the presence of CP symmetry violation within
211 oscillations is one of the most important unanswered questions in neutrino physics.

212 2.2.4 CP Symmetry Violation

213 The violation of parity symmetry in the weak interaction was discovered in observations of
214 cobalt 60 decays [7]. Under this symmetry, the weak interaction should behave the same
215 under the complete reversal of coordinate system used to describe the system $((\hat{x}, \hat{y}, \hat{z}) \rightarrow$
216 $(-\hat{x}, -\hat{y}, -\hat{z}))$. However, it was observed that, when the spins of cobalt atoms in a sample
217 were aligned in a particular direction, the electrons resulting from neutron decays within
218 the cobalt atoms (shown in Figure 2.5) came out opposite the cobalt spins. Under a parity
219 transformation, the spin would not flip, but the direction of emission would. The electrons
220 would then be emitted in the direction of the spins, thus violating the symmetry [7]. The
221 parity symmetry violation in weak interactions is in fact maximal [7], leading to it being
222 described by a left-handed chiral theory (hence the L in the subscript of 2.1). A result
223 of the parity violation is that there are no interacting right-handed (left-handed) neutrinos
224 (antineutrinos).

225 In addition to violating parity symmetry, the weak interaction also violates charge sym-
226 metry. For example, a left-handed neutrino would be transformed into a left-handed an-
227 tineutrino under charge conjugation and undergo weak interactions under charge symmetry.
228 This is not the case, however, and thus the weak interaction violates charge symmetry as
229 well. The weak interaction in the lepton sector *does not* violate CP symmetry⁴ under a
230 simultaneous charge conjugation and parity transformation, left-handed neutrinos turn into
231 right-handed antineutrinos which both undergo weak interactions.

232 Returning to the subject of neutrino oscillations, the CP-violating phase-factor δ_{CP} in
233 the PMNS matrix (Equation 2.11) has the ability to introduce CP violation in the neutrino
234 sector by producing an asymmetry between neutrinos and antineutrinos in the oscillation

⁴The weak interaction in the quark sector, however, does violate CP symmetry [7].

235 probability. If present (i.e. $\delta_{CP} \neq 0, \pi$), this oscillation asymmetry could be responsible
 236 the matter-antimatter asymmetry we observe in the universe [11]. The presence of a CP-
 237 violating value of δ_{CP} is one of the key unanswered questions in neutrino physics and is one
 238 focus of the upcoming oscillation experiments DUNE [12] and HyperK [13].

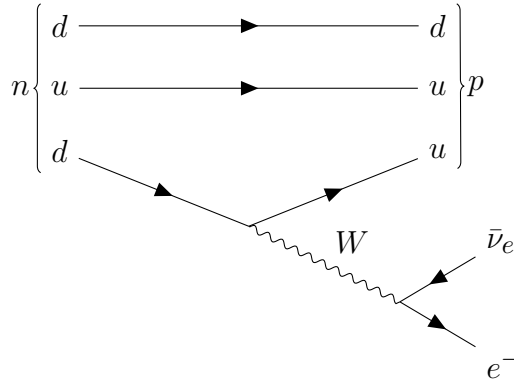


Figure 2.5: Feynman diagram of a neutron decaying to a proton.

239 2.2.5 MSW Effect

240 In matter, the oscillation probability is modified due to the presence of coherent forward
 241 elastic scattering of neutrinos by the surrounding matter. Specifically, the Charged Current
 242 process is available for ν_e , but not for $\bar{\nu}_e$, since only electrons and not positrons are present in
 243 normal matter (for example, the surrounding earth). A potential created by these processes is
 244 added to the Hamiltonian in Equation 2.5. This modifies the time evolution of the neutrino
 245 flavors, resulting in a modified oscillation probability and effective oscillation parameters
 246 in matter. This results in what is known as the MSW effect, wherein resonant behavior
 247 is exhibited in the effective mixing angles⁵. The resonance can only be present for either
 248 neutrinos or antineutrinos, depending on the ordering of the neutrino masses [14]. Thus, by
 249 measuring the asymmetry between neutrino and antineutrino oscillations, the MSW effect

⁵This is named after Mikheyev, Smirnov, and Wolfenstein. Wolfenstein first discovered that neutrinos were affected by the potential created by the surrounding matter. Mikheev and Smirnov discovered the resonant behavior.

250 can be exploited to determine the ordering of the neutrino masses. The sign of Δm_{21}^2 was
 251 determined by analyzing solar neutrinos which are subject to the matter effect as they travel
 252 through the Sun [15]. The same principle can be used to determine the mass hierarchy (the
 253 sign of Δm_{32}^2) using accelerator-based neutrino experiments. Similar to the question of CP
 254 violation, this is an important, unanswered question in neutrino physics and will be explored
 255 by future experiments such as DUNE. A discussion of DUNE and its physics potential will
 256 be given in Chapter 3. In order to give context to that chapter, the following two sections
 257 discuss interactions of both neutrinos and pions with nucleons and nuclei.

258 2.3 Neutrino Interactions

259 Neutrino oscillation experiments rely on detecting neutrinos through identifying the par-
 260 ticles produced by interactions on target nuclei. Through reconstructing these products, the
 261 flavor, sign (neutrino vs. antineutrino), and energy of the incident neutrino are inferred. This
 262 section provides a description of neutrino–nucleus interactions. In most types of neutrino–
 263 nucleus interactions⁶, the neutrino interacts primarily with a constituent nucleon. These
 264 broadly fall into two categories: Charged Current (CC) and Neutral Current (NC). The CC
 265 interaction occurs with the exchange of a W^\pm as shown by the vertex in Figure 2.1b while
 266 the NC interaction occurs with the exchange of a Z^0 shown in Figure 2.1a.

267 2.3.1 Quasielastic Scattering

268 The first major CC interaction is CC Quasielastic (CCQE). This interaction occurs in the
 269 forms given in Equation 2.12, and is represented by the Feynman diagrams shown in Figures
 270 2.6 and 2.7 for neutrinos and antineutrinos respectively.

$$\begin{aligned}
 \nu_l + n &\rightarrow l^- + p \\
 \bar{\nu}_l + p &\rightarrow l^+ + n
 \end{aligned}
 \tag{2.12}$$

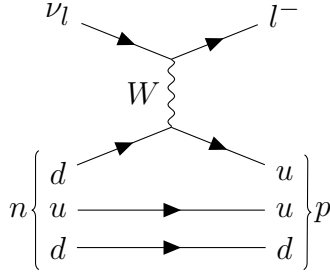


Figure 2.6: Feynman diagram for the neutrino CCQE interaction.

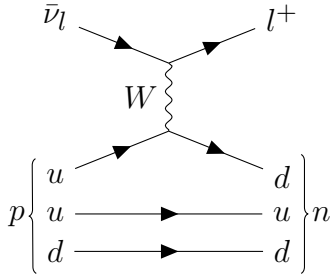


Figure 2.7: Feynman diagram for the antineutrino CCQE interaction.

271 The cross section for these processes depend on the vector form factors (F_1 and F_2) and
 272 axial form factors (F_A and F_P) of the CC interaction, which themselves depend only on the
 273 four-momentum transfer Q^2 between the neutrino and nucleon in the interaction [16]. F_1
 274 and F_2 are related to the electromagnetic form factors of the nucleons, which are extracted
 275 from electron scattering data [14]. The pseudoscalar form factor F_P can either be neglected
 276 through approximation [14] or related to F_A [16], such that F_A is the only unknown portion
 277 of the CCQE cross section. The exact shape of the axial form factor is not described by
 278 theory, and a dipole approximation is generally used as shown in Equation 2.13.

$$F_A(Q^2) = \frac{g_A}{\left(1 + \frac{Q^2}{M_A^2}\right)^2} \quad (2.13)$$

279 Here, g_A is the axial-vector coupling constant of the weak charged current, which is obtained
 280 from neutron decay data, and M_A is the axial mass, which can be obtained from fitting to
 281 neutrino scattering data.

⁶Other than coherent neutrino–nucleus scattering.

282 **2.3.2 Resonant Pion Production**

283 The next set of major neutrino–nucleon interactions are resonant pion production. These
 284 interactions occur through both CC and NC channels, and result in a pion exiting the inter-
 285 action along with the nucleon and final state lepton. In these interactions, the neutrino inter-
 286 acts inelastically with a nucleon and excites it into some resonance (i.e. a nucleon resonance
 287 or Δ resonance). The forms of the CC interactions (without specifying the intermediate res-
 288 onance) are given in 2.14, while the NC interactions are given in 2.15. Multiple resonances
 289 contribute to the amplitudes of these processes, but at lower energies, the $\Delta(1232)$ resonance
 290 dominates [17]. The most commonly used model to describe the Δ resonance interaction
 291 is the Rein-Sehgal model [16][17]. The NC interaction resulting in a π^0 is important as a
 292 background to CCQE ν_e events, where the γ showers from the π^0 decay can be mistaken
 293 as an e shower during event reconstruction, and the rate of this background is important to
 294 constrain in ν_e appearance measurements.

$$\begin{aligned}
 \nu_l + p &\rightarrow l^- + p + \pi^+, & \bar{\nu}_l + p &\rightarrow l^+ + p + \pi^- & (2.14) \\
 \nu_l + n &\rightarrow l^- + p + \pi^0, & \bar{\nu}_l + p &\rightarrow l^+ + n + \pi^0 \\
 \nu_l + n &\rightarrow l^- + n + \pi^+, & \bar{\nu}_l + n &\rightarrow l^+ + n + \pi^-
 \end{aligned}$$

$$\begin{aligned}
 \nu_l + p &\rightarrow \nu_l + p + \pi^0, & \bar{\nu}_l + p &\rightarrow \bar{\nu}_l + p + \pi^0 & (2.15) \\
 \nu_l + p &\rightarrow \nu_l + n + \pi^+, & \bar{\nu}_l + p &\rightarrow \bar{\nu}_l + n + \pi^+ \\
 \nu_l + n &\rightarrow \nu_l + n + \pi^0, & \bar{\nu}_l + n &\rightarrow \bar{\nu}_l + n + \pi^0 \\
 \nu_l + n &\rightarrow \nu_l + p + \pi^-, & \bar{\nu}_l + n &\rightarrow \bar{\nu}_l + n + \pi^-
 \end{aligned}$$

295 **2.3.3 Deep Inelastic Scattering**

Another important set of processes in oscillation experiments are the CC and NC Deep Inelastic Scattering (DIS) processes. In these, the neutrino and the intermediary gauge

boson it exchanges with the nucleon are energetic enough to resolve the individual quark constituents of the nucleon. The nucleon is broken apart and a hadron shower is produced as a result of quark confinement. The form of these interactions is given in 2.16. Here, N is either nucleon and X is a set of hadrons.

$$\begin{aligned}
 \nu_l + N &\rightarrow l^- + X, & \bar{\nu}_l + N &\rightarrow l^+ + X \\
 \nu_l + N &\rightarrow \nu_l + X, & \bar{\nu}_l + N &\rightarrow \bar{\nu}_l + X
 \end{aligned}
 \tag{2.16}$$

296 These processes dominate the total cross section at high energies ($E_\nu \gtrsim 20$ GeV) [16].
 297 The inclusive DIS cross section⁷ is described by functions representing the structure of the
 298 nucleons known as parton distribution functions (PDFs) [16][17][14].

299 There exists a transition region between the resonance and DIS regimes called the Shallow
 300 Inelastic Scattering (SIS) region [16][18]. This region is not as well understood as the DIS-
 301 dominated region [16][18], and different simulation frameworks take a variety of approach to
 302 modeling this transition [16][18].

303 2.3.4 Neutrino–Nucleus Scattering

304 Neutrino oscillation experiments use nuclear targets for their detection medium. This com-
 305 plicates the relatively simple picture of neutrino–nucleon scatter in a few key ways. Firstly,
 306 coherent scattering become possible, wherein each component of the nucleus contributes to
 307 the interaction amplitude *coherently* and the nucleus is left in its ground state. An impor-
 308 tant type of coherent scattering is coherent pion production as shown in 2.17 (top: CC,
 309 bottom: NC). The NC process is an important background for ν_e appearance channels as
 310 the γ showers from the π^0 decay can mimic an e shower.

$$\begin{aligned}
 \nu_l + A &\rightarrow l^- + \pi^+ + A, & \bar{\nu}_l + A &\rightarrow l^+ + \pi^- + A \\
 \nu_l + A &\rightarrow \nu_l + \pi^0 + A, & \bar{\nu}_l + A &\rightarrow \bar{\nu}_l + \pi^0 + A
 \end{aligned}
 \tag{2.17}$$

⁷Full expressions found in [16][17][14]

311 Additionally, the presence of the nuclear medium complicates the behavior of both the
312 initial and final state. For CCQE interactions most models assume the Impulse Approximation,
313 in which the neutrino scatters elastically off nucleons in the nuclear ground state,
314 followed by quasifree ejection of the nucleons from the nucleus. The nuclear state (i.e. the
315 kinematic distribution of nucleons within the nucleus) is commonly described by a Relativistic
316 Fermi Gas model, where the nucleons are free particles subject to Fermi motion and
317 populate states according to the Pauli exclusion principle. Despite the fact it is commonly used,
318 it poorly describes electron scattering data. Other models and approximations for the initial
319 nuclear state have been utilized in recent years to overcome this limitation. An important
320 development toward improved modeling of the initial nuclear state comes in the form of the
321 inclusion of nucleon–nucleon correlations and meson exchange currents (MEC). These contribute
322 to multinucleon excitation, and raise the cross section of events that produce no final
323 state pion. In addition to these initial state effects, DIS interactions are further complicated
324 through modifications of the nucleon PDFs by the nuclear medium [17][18].

325 Finally, the presence of Final State Interactions (FSI) can modify the observable products
326 of the primary interaction as they attempt to exit the nucleus. The hadronic products of
327 each interaction (including pions in resonance interactions and hadron showers from DIS)
328 can possibly reinteract as they travel through the nucleus. A common model for this is an
329 intranuclear cascade, wherein the interaction products step through the medium and can
330 undergo an interaction with the surrounding nucleons. These resulting particles are then
331 added into the cascade process and can then go on to interact again. This goes on until
332 all active particles exit the nucleus or are absorbed back into the nucleus. This results in a
333 modified set of observable particles (i.e. with missing or additional particles, and/or with
334 smeared kinematics). For resonance interactions, further complications arise from the fact
335 that the surrounding nuclear medium modifies the properties of the intermittent Δ resonance.
336 Processes such as those listed in 2.18 increase the width of the Δ within the nucleus [18].

$$\Delta + N \rightarrow N + N \tag{2.18}$$

$$\Delta + N + N \rightarrow N + N + N$$

$$\Delta + N \rightarrow \pi + N + N$$

337 2.4 Pions

338 The last section illustrated the complexities in neutrino scattering created by the nuclear
 339 environment. The pion is often produced in neutrino interactions, and as will be seen in
 340 Chapter 3, must be accounted for in neutrino oscillation analyses at DUNE. This section
 341 serves to describe the pion’s role within the nucleus, and its interactions with nuclei.

342 The pion is the lightest meson, and is a spin-0, isospin-1 boson. It has three charge states
 343 (as evident by its isospin). These are described in Table 2.3. Yukawa predicted that a point-
 344 particle similar to the pion mediated the force between point-like nucleons within nuclei [7][4].
 345 In fact, at ranges greater than 0.7 fm, intranuclear interactions are well described by this
 346 pion exchange picture [19]. At greater than 2 fm, one-pion exchange dominates, while two-
 347 pion exchange contributions become equal or greater than one-pion exchanges between 0.8
 348 and 2 fm [20]. Below this, the point-like approximation of the pion and nucleons breaks
 349 down, and the quark-gluon degrees of freedom become important [4][20]. In the point-like
 350 approximation, the nucleon acts as a source of the pion field, resulting in a field of the form
 351 given by Equation 2.19. Here, τ_3 and σ are Pauli isospin and spin operators and f is a
 352 coupling constant. This has a striking similarity to the potential from a magnetic dipole, as
 353 shown in Equation 2.20 [20].

$$\phi_N(\vec{x}) = -\frac{f}{m_\pi} \tau_3 \sigma \cdot \nabla_x \frac{e^{-m_\pi |\vec{x} - \vec{r}|}}{4\pi |\vec{x} - \vec{r}|} \tag{2.19}$$

$$\phi_M(\vec{x}) = -\mu \cdot \nabla_x \frac{1}{4\pi |\vec{x} - \vec{r}|} \tag{2.20}$$

Pion	Quark Content	Charge (e)	Mass (MeV/c ²)	I ₃
π^+	$u\bar{d}$	+1	139.57	+1
π^0	$\frac{u\bar{u}+d\bar{d}}{\sqrt{2}}$	0	134.98	0
π^-	$d\bar{u}$	-1	139.57	-1

TABLE 2.3: The pion along with their quark content, charge, mass, and the third component of its isospin.

354 2.4.1 Pion–Nucleon Scattering

355 It is important to consider pion–nucleon scattering as a basis for pion–nucleus scattering.
356 This interaction is purely elastic up to the threshold for the $\pi + N \rightarrow \pi + \pi + N$ process
357 at $T_\pi \approx 170$ MeV [20]. When viewed in a partial wave analysis, the s- and p-wave (angular
358 momentum $l = 0, 1$ respectively) contributions to the interaction dominate when compared
359 to the d- and f-waves ($l = 2, 3$) [20]. Furthermore, the s-wave interactions are small compared
360 to the p-wave interactions [21]. The dominant effect in the p-wave component, and thus the
361 overall interaction, is the resonance appearing around pion kinetic energy $T_\pi \approx 180$ MeV.
362 This is due to the coupling to the $\Delta(1232)$ spin 3/2, isospin 3/2 resonance [20][19][21]. This
363 resonance can be seen in Figure 2.8.

364 Two processes are of particular interest for this thesis: single charge exchange and ab-
365 sorption. For absorption, additional particles must be involved in the process in order to
366 conserve energy and momentum. As such, the absorption of pions by singular free nucleons
367 is forbidden⁸. This will be discussed in Sections 2.4.2 and 2.4.3. Single charge exchange is
368 free from this requirement and, for incident π^\pm , takes the forms in Equation 2.21. It too will
369 be discussed in Section 2.4.3.

⁸This is approximately true for bound nucleons as well, as the interaction is suppressed due to the momentum that must be supplied by the nucleon, which is much larger than the Fermi momentum [21].

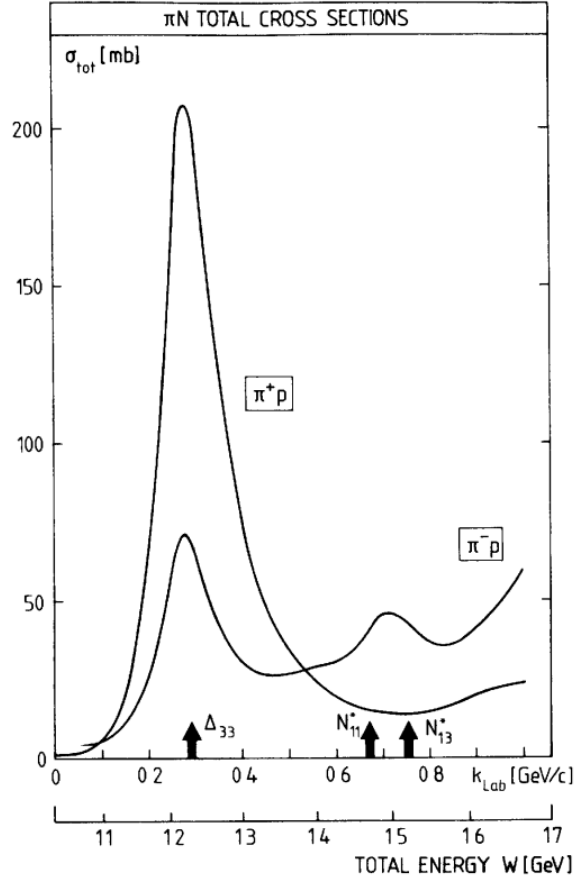
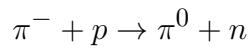


Figure 2.8: $\pi^\pm - p$ cross sections as functions of pion lab momentum k_{lab} and center-of-mass energy W [20].



370 2.4.2 Pion–Deuteron Scattering

371 The scattering of pions by deuterons (a proton–neutron bound state) is the simplest extension
 372 of pion–nucleon scattering to multiple-body systems. The total cross section is comprised
 373 of contributions from elastic scattering, inelastic scattering (wherein the deuteron is broken
 374 up), absorption, and pion production at higher energies. The interaction can be well approx-

375 imated by the sum of the $\pi - p$ and $\pi - n$ cross sections, as shown in Figure 2.9. However, the
 376 observed cross section is lower in the resonance region due to a broadening of the resonance
 377 caused by the motion of the nucleons within the deuteron as well as a shadowing effect of
 378 one nucleon by the other [20]. These effects are indicative of the complications that arise in
 379 the nuclear environment.

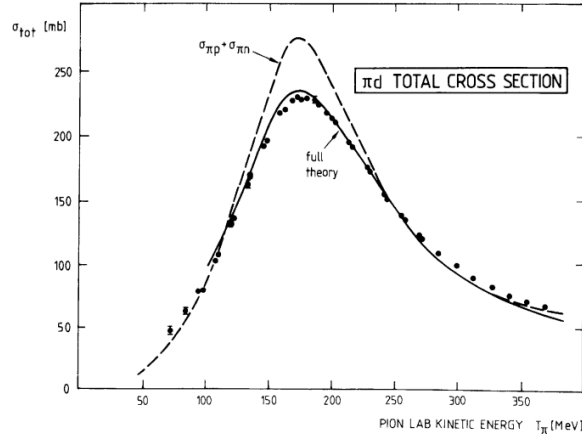


Figure 2.9: The $\pi - d$ total cross section. Black points are data, the dashed line is the sum of the $\pi - p$ and $\pi - n$ cross sections, and the solid line includes effects from nucleon motion and shadowing as described in the text [20].

380 As mentioned before, pion absorption on singular nucleons is forbidden, and multiple
 381 nucleons must contribute to the absorption process. As such, the $\pi - d$ absorption process
 382 is prototypical of this interaction in nuclei. It has been determined experimentally that
 383 two-nucleon absorption in nuclei is dominated by absorption on deuteron-like pairs in the
 384 $\Delta(1232)$ resonance region [19]. The “rescattering” model is an elementary model that gives a
 385 qualitative understanding of the physics of $\pi - d$ absorption. In this, a scatter on one nucleon
 386 is followed by absorption on the other. The leading terms in this theory are again the s- and
 387 p-wave contributions. The so-called s-wave rescattering consists of a the pion undergoing an
 388 s-wave scatter by the first nucleon, followed by p-wave absorption on the second nucleon.
 389 In p-wave rescattering, the pion strikes the first nucleon creating an intermediate Δ state.
 390 This Δ then interacts with the second nucleon creating the final dual-nucleon state [20].

391 Quantitatively, however, this description falls short, and a full three-body framework that
392 treats $\pi - d$ absorption on equal footing with other $\pi - d$ scattering processes has been more
393 successful in predicting experimental results [20].

394 **2.4.3 Pion–Nucleus Scattering**

395 Similar to the extension of pion scattering from single-nucleon targets to the deuteron, the
396 extension to the nucleus is complicated by the influence of additional nucleons on the inter-
397 action. The same basic processes (elastic and quasielastic scattering, single charge exchange,
398 absorption on more than one nucleon) are present, and the $\Delta(1232)$ resonance still plays an
399 important role. However, the dynamics are enriched by the nuclear environment. Recalling
400 the dipole-like interaction of the pion with the nucleon, the nuclear environment acts as a
401 polarizable and refractive medium for the pion, in analogy with the scattering of light by
402 electromagnetic dipoles [20]. The Δ resonance is also influenced by the medium; its peak
403 shifts lower and its width broadens as the nuclear mass increases [20][19].

404 **2.4.3.1 Elastic Scattering**

405 The analogy to light propagation is evident in elastic scattering off nuclei, whereby the
406 nucleons diffract the incoming pion wave similar to light by atoms in an optically diffractive
407 medium [20][19]. Within the resonance region, the imaginary part of the $\pi - N$ scattering
408 amplitude becomes large, producing deep minima in the angular distribution of the scatters.
409 These minima are still present, but become more shallow outside of the resonance region
410 where the real part of the scattering amplitude becomes larger [20]. This can be seen in
411 Figure 2.10.

412 **2.4.3.2 Inelastic Scattering**

413 As with elastic scattering, inelastic scattering is an extension of the $\pi - N$ interaction to the
414 nuclear environment. The difference lies in the transition of the nucleus to excited states.

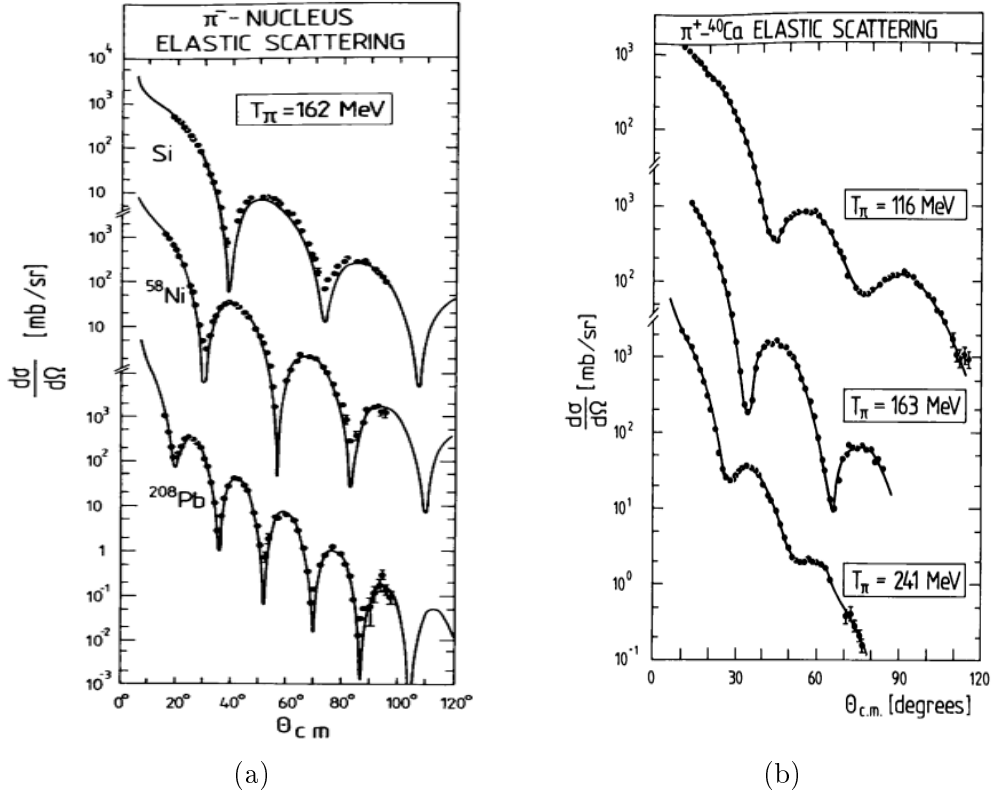


Figure 2.10: Elastic pion scattering cross sections. a) Diffractive patterns are present within the resonance region. b) Diffractive patterns are suppressed outside of the resonance region [20].

415 Two broad regimes exist for this: 1) a low energy transfer region wherein the nucleus is
 416 excited to discrete states 2) a high energy transfer region in which the quasifree $\pi - N$
 417 interaction dominates [20] and the struck nucleon is knocked out to continuum states [19].
 418 This quasifree process is the leading contribution to the inelastic cross section [20], and is
 419 subject to in-medium effects that shift the location of the quasifree peak as seen in Figure
 420 2.11.

421 2.4.3.3 Absorption

422 As previously stated, $\pi - d$ absorption is prototypical of absorption within nuclei, due to
 423 the suppression of single nucleon absorption by energy and momentum conservation. For
 424 nuclei with $A > 2$, absorption quasideuteron pairs ($I = 0$, pn pairs) remains the leading

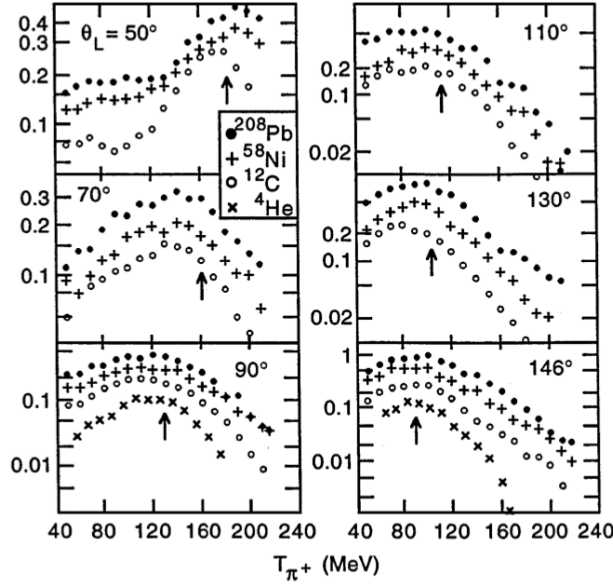


Figure 2.11: A collection of data representing spectra of pions relative to outgoing kinetic energy T_{π} and lab angles θ_L for inclusive inelastic scattering on various nuclei. The arrows represent quasifree peaks assuming no in-medium effects applied to the $\pi - N$ scattering amplitude [19].

425 contribution to 2-nucleon absorption [19]. However, the presence of more nucleons influences
 426 the interaction in a few ways. Firstly, direct interactions on multi-nucleon ($N > 2$) groups
 427 contribute to the cross section. The absorption of π by 3 nucleons becomes significant even
 428 for ${}^3\text{He}$ targets [19]. These direct absorption processes provide insight into correlations
 429 between nucleons, and experiments such as LADS have detailed measurements of π -nuclear
 430 absorption relative to outgoing nucleon multiplicities and kinematics [22]. Additionally,
 431 multiple nucleons become involved in the absorption via Initial State Interactions (ISI –
 432 wherein the pion undergoes a quasifree scatter off a single nucleon and is absorbed later)
 433 and Final State Interactions (FSI – wherein the pion is absorbed by a set of nucleons and
 434 these nucleons go on to interact with other nucleons within the nucleus). Data has shown
 435 that the average number of nucleons substantially involved in the absorption process appears
 436 to be considerable and at least somewhat A -dependent [19]. The contribution of ISI to
 437 multinucleon absorption of pions in nuclei appears constant relative to A [19], while FSI
 438 contributes more as A increases [20]. Additionally, the energy spectra of exiting nucleons is

439 similar to multistep, cascade processes [20].

440 2.4.3.4 Single and Double Charge Exchange

441 π -nuclear single charge exchange (the processes shown in Equation 2.21) is an extension of
442 the quasifree inelastic interaction discussed above, but with a $\Delta I_3 = \pm 1$ transition of the
443 nuclear isospin [20]. Charge exchange makes up roughly 10% of the π -nuclear reaction⁹
444 in the resonance region [20]. Like the other interactions described so far, the process is
445 complicated by the nuclear medium. Mainly, double charge exchange ($\pi^\pm + A \rightarrow \pi^\mp + A'$)
446 can occur. Here, two subsequent single charge exchange interactions occur. The outgoing π^0
447 from the initial interaction exchanges charge with another nucleon, resulting in a $\Delta I_3 = \pm 2$
448 isospin transition and a flip of the pion's charge. Double charge exchange is a relatively rare
449 process, with a cross section roughly 10% of the single charge exchange cross section [20].

450 2.4.4 Outlook

451 As highlighted by this section, pion-nucleus interactions contain complex dynamics. Partic-
452 ularly in heavy nuclear environments, these interactions become quite complicated. DUNE's
453 nuclear target, argon, is no exception to this, and so care must be taken to model these pro-
454 cesses within DUNE's experimental simulation. Currently, limited data exists for pion-Ar
455 interactions, especially the exclusive interactions like absorption (a single measurement [22])
456 and charge exchange (no measurements). This thesis provides data for these interactions
457 that can be used to validate and improve the pion interaction model used by DUNE.

⁹Absorption + Inelastic + Single Charge Exchange + Double Charge Exchange interactions

CHAPTER 3

458

THE DEEP UNDERGROUND NEUTRINO EXPERIMENT

459 As discussed in the previous chapter, there are a few questions in neutrino physics that
 460 remain unanswered. The Deep Underground Neutrino Experiment (DUNE) seeks to an-
 461 swer these questions once it begins to take data later this decade. Many experiments have
 462 already made enormous progress in getting us to the point where the answers to these ques-
 463 tions are in reach. These experiments focused on neutrinos produced from several sources:
 464 neutrinos produced in nuclear reactors, neutrinos produced by cosmic ray interactions in
 465 the atmosphere, neutrinos produced within the Sun, and neutrinos produced from particle
 466 accelerators. A global fit to the data from these experiments has been performed to provide
 467 current estimates of the oscillation parameters [23]. These are presented in Table 3.1. Note
 468 that only the normal mass ordering is given here, as the analysis found the inverted ordering
 469 was disfavored at $\Delta\chi^2 = 4.7$ [23].

470 Though systematic uncertainties in previous-generation experiments have required great
 471 effort to overcome, the experiments were limited primarily by statistical uncertainty. DUNE

Parameter	Best-Fit $\pm 1\sigma$	3σ Range
$\sin^2 \theta_{12}$	$0.310^{+0.013}_{-0.012}$	0.275 – 0.350
$\theta_{12} [^\circ]$	$33.82^{+0.78}_{-0.76}$	31.61 – 36.27
$\sin^2 \theta_{23}$	$0.582^{+0.015}_{-0.019}$	0.428 – 0.624
$\theta_{23} [^\circ]$	$49.7^{+0.9}_{-1.1}$	40.9 – 52.2
$\sin^2 \theta_{13}$	$0.022240^{+6.5 \times 10^{-4}}_{-6.6 \times 10^{-4}}$	0.02044 – 0.02437
$\theta_{13} [^\circ]$	$8.61^{+0.12}_{-0.13}$	8.22 – 8.98
$\delta_{CP} [^\circ]$	217^{+40}_{-28}	135 – 366
$\Delta m_{21}^2 [10^{-5} \text{eV}^2]$	$7.39^{+0.21}_{-0.20}$	6.79 – 8.01
$\Delta m_{31}^2 [10^{-3} \text{eV}^2]$	$2.525^{+0.033}_{-0.031}$	2.431 – 2.622

TABLE 3.1: Oscillation parameters as determined by the fit to global data in Reference [23]. Only the normal ordering of the mass hierarchy is shown here.

472 is a next-generation long baseline accelerator-based neutrino oscillation experiment, and
473 will collect enough neutrino events to become limited primarily by systematic uncertainties.
474 This chapter provides the motivation for the results of this thesis which will be used to meet
475 DUNE’s stringent systematic uncertainty requirement.

476 **3.1 DUNE’s Physics Program**

477 The goals of DUNE’s accelerator-based oscillation analyses will be to determine whether
478 neutrino oscillations violate CP-symmetry, determine the neutrino mass hierarchy, and to
479 determine precise values of the oscillation parameters. The DUNE Far Detector Technical
480 Design Report [24] presents sensitivity studies which show DUNE’s ability to achieve these
481 goals. In these studies, simultaneous fits to $\nu_\mu \rightarrow \nu_\mu$, $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$, $\nu_\mu \rightarrow \nu_e$, and $\bar{\nu}_\mu \rightarrow$
482 $\bar{\nu}_e$ far detector samples were performed, with near detector samples included in order to
483 introduce flux and cross section constraints. $\sin^2 2\theta_{13}$, θ_{12} , and Δm_{12}^2 were all constrained
484 with uncertainties derived from those shown in Table 3.1, while $\sin^2 \theta_{23}$, Δm_{32}^2 , and δ_{CP}
485 were freely varied. More details on the fits can be found in [24].

486 The sensitivity studies show promise in DUNE’s physics program. For 50% of true
487 δ_{CP} values, DUNE can determine the presence of CP violation at the 5σ level after 10
488 years of its nominal run plan. If $\delta_{CP} = -\pi/2$ (which provides a maximal CP-violating
489 effect), CP violation can be discovered after only seven years. For any value of δ_{CP} , the
490 mass hierarchy can be determined after only two to three years. This reduces to only
491 about one year if $\delta_{CP} = -\pi/2$. After about fifteen years of the nominal run plan, the
492 resolution on the measurement of δ_{CP} approaches 5° for CP-conserving values and 15°
493 for CP-violating values. After high exposure, the measurement of $\sin^2 2\theta_{13}$ approaches the
494 precision of reactor experiments (which currently provide the main constraints on that angle),
495 and the simultaneous measurement of all oscillation parameters without external constraint
496 becomes possible.

497 3.2 The DUNE Detectors

498 DUNE seeks to achieve these goals as a long-baseline oscillation experiment, and, as such,
499 is comprised of two sets of detectors: its far detector (FD) and near detector (ND) complexes.
500 This is shown in Figure 3.1, which gives an overview of DUNE’s facilities including the
501 neutrino beam facility and the near detector complex located at Fermilab (Batavia, IL) and
502 the far detector complex at Sanford Underground Research Facility (Lead, SD).

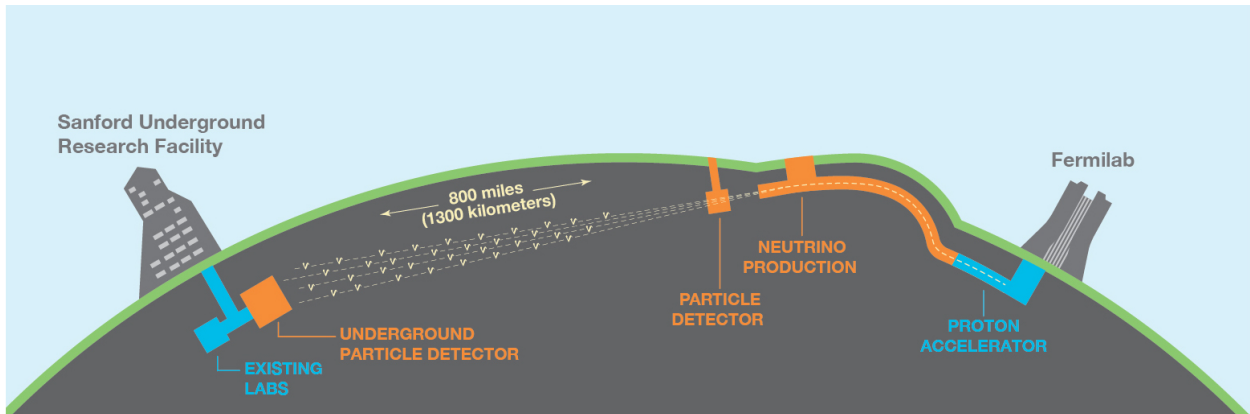


Figure 3.1: Overview of the future DUNE experiment. Toward the right is the neutrino beam facility and the near detector complex at Fermilab in Batavia, IL. Toward the left is far detector complex 1300km away at Sanford Underground Research Facility in Lead, SD [25].

503 The FD complex seeks to measure the number and flavor of neutrinos after they have had
504 a chance to oscillate after traveling some distance. By measuring the rates at the FD, the
505 oscillation probabilities (and more specifically, the parameters describing these probabilities)
506 are probed directly. The ND, on the other hand, provides constraints on flux and neutrino
507 cross section uncertainties within the models used for the oscillation analyses.

508 The planned DUNE FD will be comprised of four modules. Two of the modules (including
509 the first to be installed) will be 10kt active volume single-phase (SP) Liquid Argon Time
510 Projection Chambers (LArTPCs). This detector technology is the same as ProtoDUNE-SP
511 and will be explained in detail in Chapter 4. One other module will be a dual-phase LArTPC
512 (which is slightly different to the SP technology, but will not be explored here), while the

513 final module’s design is still to be determined. One common trait between all four modules
514 is that their sensitive volumes will be 10kt of liquid argon. This argon will serve as both the
515 target and detection medium for DUNE’s neutrino beam.

516 The DUNE ND complex will be comprised of multiple detector subsystems. Included in
517 these subsystems is a set of small, modular LArTPCs known as ArgonCube. It is necessary to
518 have a portion of the near detector’s target be argon in order to cancel neutrino interaction
519 model uncertainties between the near and far detectors. This part of the detector will
520 also be allowed to move lateral to the incident neutrino beam. Because the far detector
521 is located at an angle of 0° with respect to the beam direction (it is “on-axis”), gathering
522 data “off-axis” provides independent measurements of the neutrino beam. This off-axis data
523 reduces systematic uncertainties surrounding the neutrino beam model. Other subsystems
524 in the ND complex include a gaseous argon TPC downstream of the LArTPC portion (to
525 help measure muons which punch through the back of the LArTPC), a fine-grained plastic
526 scintillator detector (which remains on-axis to monitor the stability of the beam), and an
527 electromagnetic calorimeter surrounding the previous two subsystems (which will assist in
528 measuring all of the final state energy within the neutrino interactions).

529 **3.3 The Role of Pion Interaction Systematic Uncertainties**

To precisely measure the oscillation parameters, DUNE will attempt to discern the flavor and energy of the neutrinos interacting within the detector. Equation 2.10 (repeated here), shows the importance of successfully determining these quantities. Misidentification of the flavor will of course change the overall interaction rates of the various neutrino flavors, thus the extracted oscillation probability. Misestimation of the energy will change where in the energy distribution of interactions an event lies, thus distorting the energy spectrum of

events, and further distorting the apparent oscillation probability.

$$\begin{aligned}
 P(\nu_a \rightarrow \nu_b) &= \delta_{ab} \\
 &- 4 \sum_{j>k} \Re \left[U_{bj}^* U_{aj} U_{bk} U_{ak}^* \right] \sin^2 \left(\frac{\Delta m_{jk}^2 L}{4E} \right) \\
 &\pm 2 \sum_{j>k} \Im \left[U_{bj}^* U_{aj} U_{bk} U_{ak}^* \right] \sin \left(\frac{\Delta m_{jk}^2 L}{2E} \right)
 \end{aligned} \tag{2.10}$$

530 Both of these quantities are inferred from the final state particles resulting from the inter-
 531 action. Figure 3.2 shows an example ν_μ interaction with multiple hadrons in the final state highlighting how complicated the final state of the interaction can be.

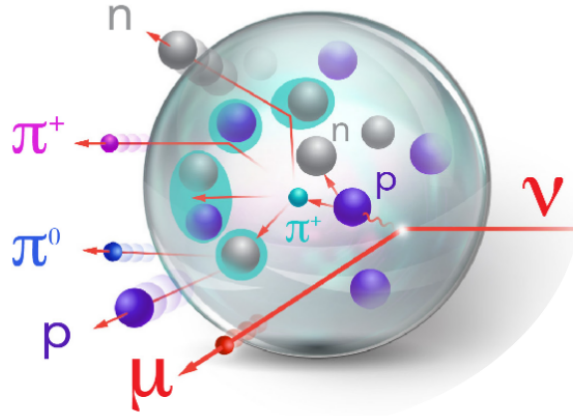


Figure 3.2: Cartoon of a ν_μ interaction with multiple hadrons in the final state [26].

532

533 For determining the flavor, reconstruction software attempts to identify the outgoing
 534 leptons from CC interactions (μ^\pm and e^\pm from muon and electron neutrinos respectively).
 535 Specifically for ν_e CC events, the e^- will produce an electromagnetic shower at the interaction
 536 vertex. A background to this interaction is a ν_μ NC interaction with a π^+ in the final state.
 537 This π^+ can potentially strike a nearby Ar nucleus and create a π^0 in a charge exchange
 538 interaction. This π^0 will promptly decay into two photons which will shower similarly to the
 539 e^- . This could cause this event to be wrongly selected as a ν_e CC event. Corrections for
 540 this type of background is taken from simulation, and any uncertainty on the rate of π^+ -Ar

541 charge exchange interactions will translate to an uncertainty on the true number of ν_e events
 542 (and thus limit the precision of the oscillation measurements).

543 Similarly, smearing between true and reconstructed neutrino energy will be influenced on
 544 the modeling of π^+ interactions. In DUNE, the neutrino energy is estimated by the energy
 545 of the final state particles using a calorimetric energy reconstruction given in Equation 3.1.
 546 Here, E^l is the energy of the outgoing lepton, T_i^{Nucleon} is the kinetic energy of any final
 547 state protons or neutrons, and E_i^π is the total energy of pions in the final state¹. The rest
 548 mass of any pion must be included since some of the incident neutrino energy must be used
 549 to produce the pion. A π^+ in the final state could undergo an absorption interaction on a
 550 nearby nucleus and produce a proton. The reconstruction software could fail to identify that
 551 there was a pion in the final state and the rest mass of the pion could be lost in Equation 3.1.
 552 Again, simulation is used to account for this type of effect (and similar other effects), and any
 553 uncertainty in the rate of π^+ -Ar absorption limits the resolution of oscillation measurements.

$$E_{\text{Reco}}^\nu = E^l + \sum_i T_i^{\text{Nucleon}} + \sum_i E_i^\pi \quad (3.1)$$

554 These two examples are not an exhaustive list of regions where uncertainty on the rates
 555 of these interactions will add to DUNE's total systematic uncertainty. Rather, they are
 556 illustrative of the goal of the analysis presented in this thesis. Measurements of these inter-
 557 action cross sections will provide constraints within DUNE's oscillation analyses, and will
 558 reduce DUNE's systematic uncertainty. This is an important task, as DUNE's systematic
 559 uncertainty budget is limited to 2% in order to achieve the physics goals laid out in this
 560 chapter [24]. An example of how pion scattering data can be used for the benefit of neutrino
 561 experiments is given by T2K's use of world π^+ scattering data to constrain the nuclear model
 562 used within their neutrino interaction simulation [27].

¹Other particles such as kaons have been ignored for this example, but, in general, could be present in the final state.

CHAPTER 4

THE PROTODUNE-SP DETECTOR

563

564 Currently, the single-phase ProtoDUNE detector (ProtoDUNE-SP) is the world's largest
565 active Liquid Argon Time Projection Chamber (LArTPC). This detector is designed to be
566 a prototype of DUNE's single phase far detector, located in CERN's North Area. Detector
567 installation and integration began in 2017 and finished Summer 2018. This was followed by
568 a commissioning phase (including its charged particle beam line commissioning) in the late
569 Summer & early Fall of 2018. After commissioning, cosmic ray data and beam line data
570 was taken up to the CERN long shut down¹. Since then, cosmic ray data taking has been
571 ongoing.

572 Section 4.1 describes the general operation principles of LArTPCs. Section 4.2 describes
573 the specific design of ProtoDUNE-SP. Section 4.3 provides a description of the characteri-
574 zation of data taken by the TPC. Section 4.4 describes the reconstruction of events in the
575 TPC. Section 4.5 highlights the calibration of the detector. Finally, Section 4.6 describes
576 the Monte Carlo simulation of events within the detector. The beam line will be described
577 separately in Chapter 5.

578 4.1 LArTPC Principles

579 The detection principles of LArTPCs are based on the detection of ionization electrons
580 and scintillation light produced by charged particles passing through the liquid argon (LAr).
581 The argon sits between a set of anode wires and a cathode, which create a (nominally)
582 uniform electric field. The ionization electrons drift along the electric field toward the anode
583 wires. These wires are instrumented with electronics and detect signals produced by the
584 drifting ionization. A configuration can be achieved such that several planes of wires can

¹During this time, the Super Proton Synchrotron, from which the ProtoDUNE-SP beam line originates, was shut down to allow for upgrades.

585 measure the ionization. In such a configuration, two wire planes sit in front of a third. The
 586 electric field lines terminate on the third plane, meaning the drifting electrons ultimately
 587 deposit onto this wire. Thus this plane is called the “collection plane.” Before collection,
 588 the electrons drift past the two other “induction planes.” Bipolar signals are induced on
 589 these planes as the electrons drift first toward and then away from these wires on their way
 590 toward the collection plane. If these are oriented in different directions to the collection
 591 plane, the combination of signals provide a 2D projection of the charged particle’s position
 592 as it traversed the LAr. The third dimension is given by the time which the ionization took
 593 to finish drifting. As the drift velocity is constant and known from the electric field, one can
 594 measure the initial position as:

$$x = t_{\text{drift}} * v_{\text{drift}} = (t_f - t_0) * v_{\text{drift}} \quad (4.1)$$

595 where x is the lateral position of the track, t_f is the readout time, t_0 is the time at the
 596 start of charge drift, and v_{drift} is the known drift velocity. These principles are shown in
 597 Figure 4.1. This shows a neutrino interaction producing two charged particles. These go on
 598 to ionize the LAr, and the ionization electrons drift against the electric field created by the
 599 anode wires and cathode plane. Signals are produced on the wires: the plane labeled ‘V’
 600 shows bipolar signals created by induction; the plane labeled ‘Y’ shows unipolar collection
 601 signals.

602 By ionizing the LAr as it travels through the TPC, the charged particle loses energy.
 603 Thus, by measuring the amount of ionization (the size of the signals produced on the wires),
 604 one can measure the energy lost by the particle during its traversal of the LAr. This allows
 605 LArTPCs to provide calorimetric energy measurements of the particles it detects. The
 606 ProtoDUNE-SP event display shown in Figure 4.2 highlights this capability. In this, a beam
 607 particle enters the TPC from the left of the figure. It travels through the LAr until it
 608 undergoes an interaction with an Ar nucleus, producing two visible particle tracks. The
 609 strength of the signals is shown by the color of the tracks. The incident beam particle

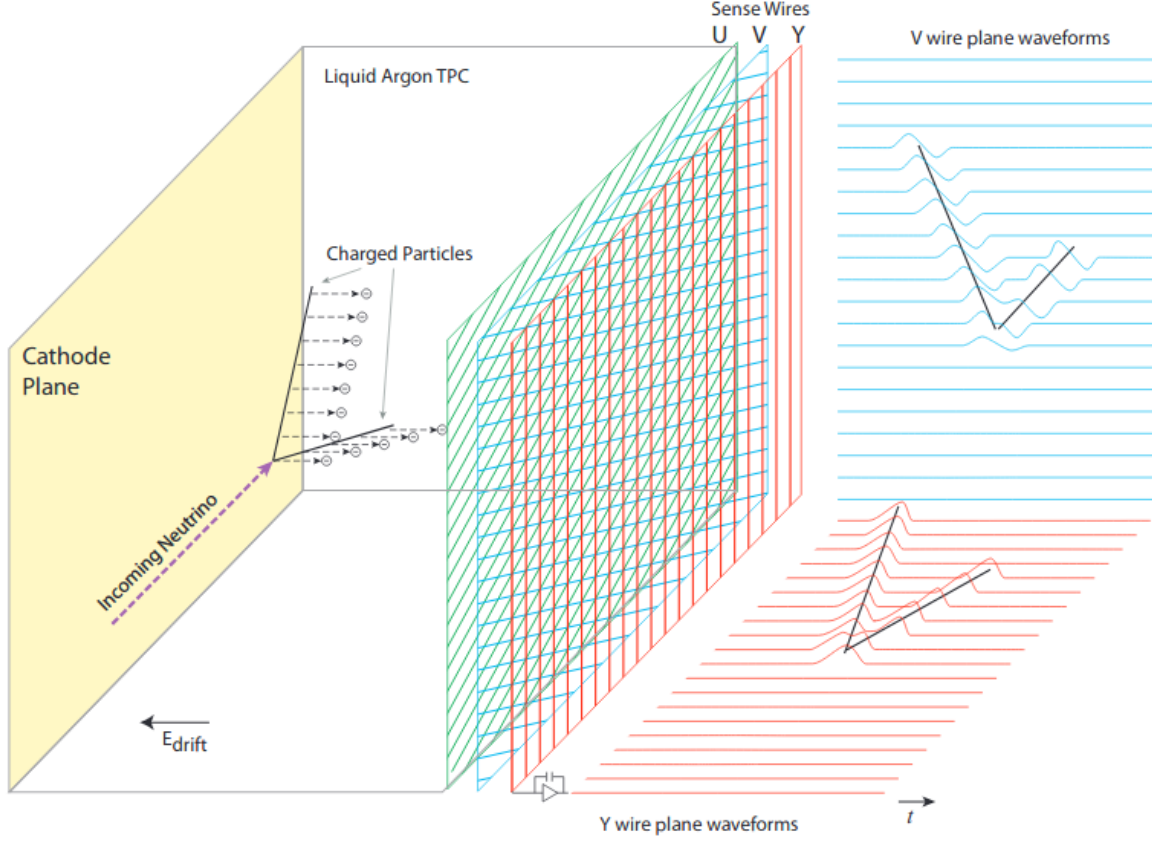


Figure 4.1: Design and operating principles of a LArTPC[28].

610 is a beam π^+ candidate, and deposits considerably less energy per unit distance than the
 611 products of its interaction.

612 Detecting the deposited energy provides Particle Identification (PID) of the charged
 613 particles because the mean rate of energy loss is well described by the Bethe formula[29]
 614 shown in Equation 4.2.

$$\left\langle -\frac{dE}{dx} \right\rangle = Kz^2 \frac{Z}{A} \frac{1}{\beta^2} \left(\frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 W_{max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right) \quad (4.2)$$

615 Here, z is the charge number of the incident particle, Z and A are the atomic number and
 616 mass of the material through which the particle is traveling, and I is the mean excitation
 617 of the material. $K = 4\pi N_A r_e^2 m_e c^2$ where N_A is Avogadro's number, and m_e and r_e are
 618 the mass and classical radius of the electron. W_{max} is the maximum energy transfer to an

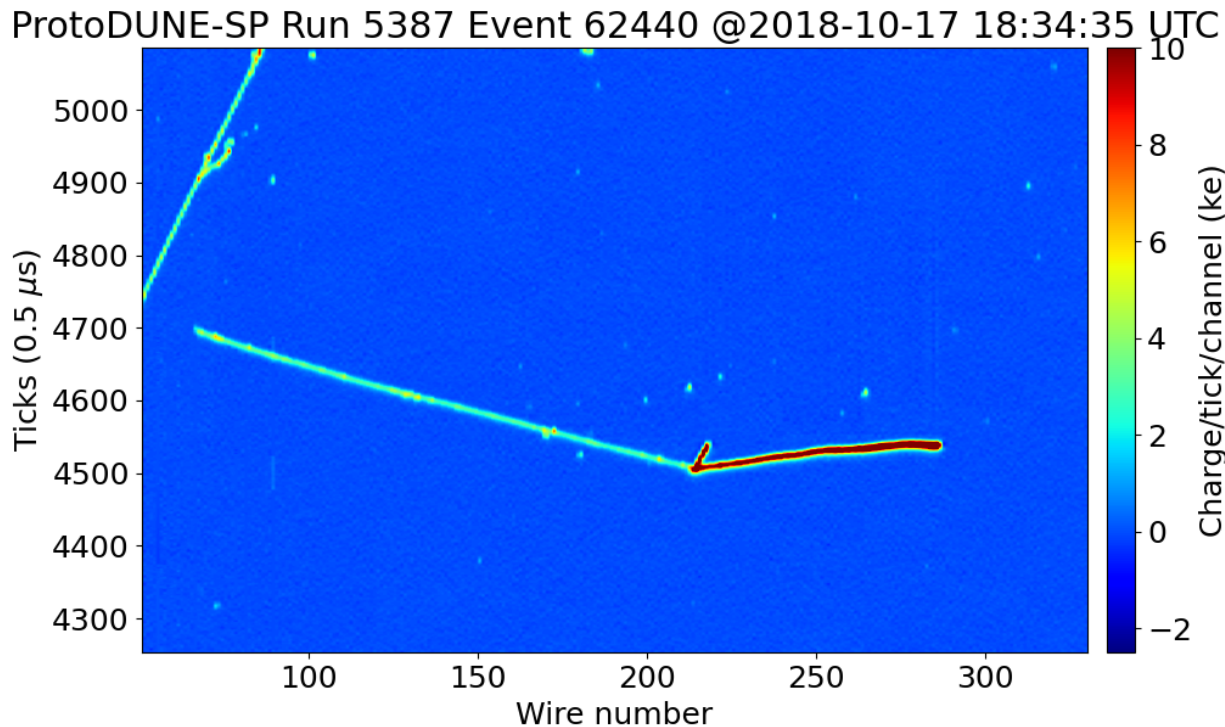


Figure 4.2: ProtoDUNE-SP event display showing a candidate beam π^+ entering from the left and undergoing an interaction with an Ar nucleus.

619 atomic electron for a single collision by a particle of mass M and is given by Equation 4.3.

$$W_{max} = \frac{2m_e c^2 \beta^2 \gamma^2}{1 + 2\gamma m_e/M + (m_e/M)^2} \quad (4.3)$$

620 The main dependence on the incident particles comes from the factor $1/\beta^2$ [29]. For particles
 621 at the same energy, heavier particles will have a smaller β and thus will deposit a larger
 622 amount of energy per unit length. PID can be performed on particles by observing how
 623 much energy they deposited along their travel through the LAr. Thus, the products of
 624 the interaction in Figure 4.2 appear to be protons, making this event a candidate for π^+
 625 absorption. The exact technique used in this analysis to separate π^\pm from protons will be
 626 described in Chapter 7.

627 In addition to ionization, the charged particles create scintillation light by the excitation
 628 and subsequent radiative decay of argon excimers. Scintillation light from LAr is produced
 629 isotropically in a narrow band around 128 nm and has a large yield of 24,000 photons per

630 MeV deposited at a drift field of 500 V/cm (ProtoDUNE-SP’s operating drift field value)[30].
631 It is produced in both a fast ($\sim 5\text{ns}$) and slow ($\sim 1.3 - 1.4 \mu\text{s}$) component at a ratio of 1:3[30].
632 The LAr is transparent to its own scintillation light[28], allowing photon detectors within the
633 LAr to collect the light produced by the charged particles. This provides important timing
634 and triggering capabilities for neutrino experiments.

635 This section provided an overview of LArTPC detection principles. However, certain com-
636 plications arise in normal operating situations. The next subsections will provide overviews
637 of the following effects in LArTPCs which modify this simple interpretation of tracking:
638 recombination of ionization electrons(4.1.1); attenuation of ionization electrons due to LAr
639 impurities (4.1.2); the Space Charge Effect created by the accumulation of positive Ar in the
640 bulk of the LAr(4.1.3).

641 4.1.1 Recombination

642 Recombination is the effect by which ionized electrons thermalize with the Ar and then
643 quickly attach to the positive Ar^{1+} ions created by the charged particle. This modifies
644 the charge observed by the wires and must be accounted for in calibration in order to
645 accurately measure the energy deposited by charged particles. There has been limitation
646 in theoretical treatments of recombination to provide a global description of data [31]. The
647 preferred models for LArTPCs with similar electric fields to ProtoDUNE-SP are the Birks
648 and Box models. Both models are based off of the principle recombination effect arising from
649 ionization electrons attaching to other Ar^{1+} ions created by the charged particle (as opposed
650 to reattaching to its original atom) and both depend on the electric field. These models differ
651 in that the Box model neglects electron diffusion and ion mobility during recombination and
652 uses “Box model” boundary conditions rather than Birks’ cylindrical assumptions [31]. The
653 ICARUS experiment found good agreement to fits of the Birks model [32], while ArgoNeuT
654 achieved good agreement with a “modified Box model” [31] which enabled another parameter
655 to vary in order to achieve agreement to the Birks model at low dE/dx [31]. ProtoDUNE-SP

656 also adopted this modified Box model.

657 **4.1.2 Ionization Attenuation**

658 Impurities in the LAr, such as water and O₂, can capture ionization electrons as they drift
659 toward the anode plane. This reduces the final amount of collected charge, and is simply
660 modeled as an exponential decay as in Equation 4.4. Here, Q_C is the collected charge, Q_0 is
661 the initial charge deposited, t_d is the drift time, and τ is the “drift electron lifetime.” This
662 lifetime is lowered by the presence of impurities.

$$Q_C = Q_0 e^{-t_d/\tau} \quad (4.4)$$

663 **4.1.3 Space Charge Effect**

664 The Ar¹⁺ ions created by the charged particles drift toward the cathode, but at a much
665 slower velocity. As such, if enough positive ions are created, positive charge can build up
666 in the bulk of the LArTPC. This accumulated charge can distort the electric fields, causing
667 the so-called Space Charge Effect (SCE). This is especially the case for LArTPCs on Earth’s
668 surface such as ProtoDUNE-SP. These surface detectors are subject to a large cosmic ray
669 flux, which constantly replenishes the positive charge. This large accumulation of charge
670 causes the field lines to be bend toward the center of the TPC, resulting in distorted particle
671 tracks. Through modifying the electric field, the SCE also changes recombination. The
672 specifics of SCE in ProtoDUNE-SP will be discussed in Section 4.5.1.

673 **4.2 The ProtoDUNE-SP Detector**

674 With a total of 770 tonnes of LAr (420 tonnes are within the instrumented volume),
675 ProtoDUNE-SP is the largest LArTPC ever constructed [30]. It provided a test bed for many
676 components and engineering challenges of the single phase technology that will comprise the
677 first DUNE far detector. It was designed to satisfy stringent requirements and achieve

678 improved levels of LArTPC performance required by DUNE, and it surpassed these in many
679 cases [30]. This section describes the design of the ProtoDUNE-SP components including
680 the following: the cryostat surrounding the TPC and the LAr purification system (4.2.1); the
681 TPC components (4.2.2); the Cold Electronics (CE) used to readout the TPC signals (4.2.3);
682 the photon detector system used to readout scintillation light (4.2.4); the cosmic ray tagger
683 (4.2.5); the Data Acquisition (DAQ), timing, and triggering systems (4.2.6). The detector
684 will be described in terms of a right-handed coordinate system with y as the vertical axis
685 pointing up, z horizontal and pointing approximately along the beam axis, and x horizontal
686 and pointing along the electric field.

687 4.2.1 Cryostat and Purification

688 The cryostat, cryogenics, and purification system serve the role of keeping the argon in
689 a liquid state with as little impurities as possible in order to avoid signal attenuation as
690 described in Section 4.1.2. The TPC is encased in a membrane cryostat, which is formed of a
691 corrugated membrane that holds the liquid and gaseous (from boil-off) argon with insulation,
692 fireproofing, and supports outside of this [33]. The internal dimensions are 8.5m x 7.9m x
693 8.5m, making this the largest LAr cryostat ever constructed [30]. The membrane contains
694 several openings to allow installation of detector elements, electrical/signal feedthroughs, the
695 support structure for the TPC (which is suspended within the membrane), and cryogenic
696 systems [34].

697 Due to ProtoDUNE-SP's larger drift distance (3.6m), a higher purity of LAr had to
698 be achieved in order to limit the attenuation of ionization during drift. The purification
699 systems used for ProtoDUNE-SP were built off those developed for ICARUS, MicroBooNE
700 (another LArTPC neutrino experiment at Fermilab), and a LAr purity demonstrator based
701 at Fermilab [30]. This purification system is the largest to date, and, along with the rate of
702 recirculation and avoidance of leaks in the cryostat, reached an equivalent oxygen contam-
703 ination of a few parts per trillion (ppt) [30]. This is in line with DUNE's requirement for

704 <100 ppt contamination in its single phase far detector [35].

705 4.2.2 TPC

706 The TPC of ProtoDUNE-SP is an active volume of 7.2m x 6.0m x 6.9m separated by a
707 cathode at $x = 0$ into two drift volumes each of drift distance 3.6m and a drift field of
708 500 V/cm. The cathode is formed of six Cathode Plane Assemblies (CPAs) biased at -180
709 kV. Each side contains three Anode Plane Assemblies (APAs) opposite the cathode which
710 contain the instrumentation wires and CE used to readout the wires. Surrounding the top
711 and bottom and sides parallel to the drift field is the Field Cage (FC) that provides (in
712 addition to the APAs and CPAs) electrostatic boundary conditions to achieve the intended
713 drift field. Penetrating into the $x < 0$ drift volume (henceforth called the “beam side”) is the
714 Beam Plug which minimizes the energy loss and interactions of beam particles with inactive
715 material. This layout is shown in Figure 4.3.

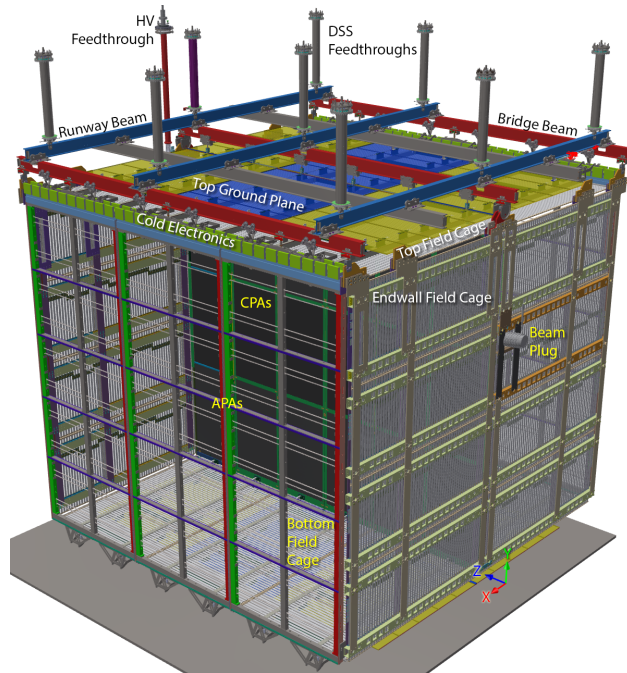


Figure 4.3: Diagram of the ProtoDUNE-SP TPC components [34].

716 The CPAs are 1.15m wide and 6.1m and consist of three vertically stacked cathode panels.

717 In order to avoid an electrical breakdown of the TPC which could destroy the CE, the CPAs
718 are constructed of heavily resistive materials which give them a very long discharge time. The
719 panels are constructed from a fire-retardant fiberglass-epoxy composite and are laminated
720 on both sides with a Kapton film [30].

721 The APAs are formed of a rectangular stainless steel frame 6.1m high, 2.3m wide, and
722 76mm thick. Bonded directly over each side of the frame is a bronze wire mesh with 85%
723 optical transparency that provides a grounded shield plane for four sets of wires on each side
724 of the frame. Each successive wire plane is 4.75mm above the previous, with the inner most
725 plane also 4.75mm above the mesh. The inner most plane is the X plane and is oriented
726 vertically. Above the X plane is the V layer oriented at -35.7° from vertical, proceeded
727 by the U plane oriented at $+35.7^\circ$ from vertical. Finally, the Grid (G) plane lies above
728 the U plane and is oriented vertically. The G plane serves as a protective shield against
729 electrostatic discharge and is not read out. The rest of the wires are connected to front-end
730 CE and serve as the main instrumentation wires. The voltages of the wire planes ($V_{G=-665}$
731 V, $V_{U=-370}$ V, $V_{V=0}$ V, $V_{X=+820}$ V) are chosen such that the field lines terminate on the
732 X plane, thus designating the X plane as the collection plane. The V and U planes are
733 thus the induction planes. The X and G planes both have a wire pitch of 4.79mm, but
734 are staggered from each other by half a wire pitch (meaning the G plane wires sit above
735 and between two X plane wires). The V and U planes both have a wire pitch of 4.67mm.
736 Each side has separate X and G planes, while the V and U planes are wrapped once around
737 the APA. The angle of the V and U planes is such that 1) each wire crosses only a given
738 collection wire on each side only once and 2) an integral number of CE boards reads out
739 one APA. The first point serves to reduce ambiguities in track reconstruction. A diagram
740 of an APA with a limited number of wires displayed is shown in Figure 4.4. Additionally,
741 electron diverters were installed between the APAs on the beam. These were formed of two
742 vertical electrode strips mounted on insulating board that, with voltages applied between
743 the electrodes, modified the local drift field such that electrons drifted away from the gaps

744 and toward the active area. During operation, high currents were drawn from the diverters’
 745 power supplies due electrical shorts in the cold volume. They were therefore left unpowered
 746 during operation, and, due to a resistive path to ground, the outer electrode was grounded.
 747 This was not the intended voltage, as it then collected charge near the gaps between APAs
 748 and distorted tracks crossing between APAs [30].

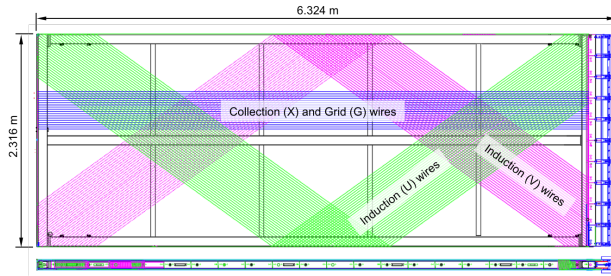


Figure 4.4: Diagram of an APA with its wire planes labeled. The bronze wire mesh is not shown. As it is shown, it is oriented on its side. The right side of the figure is the top of the APA when it is oriented vertically. The connections to the front-end CE boxes can be seen on the right side [34].

749 The Field Cage covers the remaining four sides of the drift volumes not covered by the
 750 APAs or cathode plane. It provides the remaining electrostatic boundary conditions to
 751 create uniform electric fields in the drift volumes. The top and bottom are comprised of six
 752 FC assemblies each, while four end wall panels each consisting of four assemblies oriented
 753 parallel to the x direction (the nominal drift direction). The assemblies are made of parallel
 754 metal profiles connected to each other by a resistive divider chain to provide the voltage
 755 gradient, I-beams that form an insulating support structure, and ground planes for the top
 756 and bottom assemblies. The ground planes prevent (at the top) a high electric field entering
 757 the gaseous argon and (at the bottom) a high electric field reaching the cryostat floor and
 758 cryogenic services [34].

759 On the beam side FC wall closest to the beam ($z \approx 0$), a beam plug is installed. This
 760 plug displaces the LAr and reduces the mass through which the beam particles must travel
 761 before reaching the TPC. This then reduces the energy loss and interactions upstream of
 762 the active volume. It is formed of a series of alternating fiberglass and stainless steel rings,

763 forming a cylinder capped by low mass fiberglass plates. It extends about 5 cm inside the
764 field cage boundary. A printed circuit board acting as a mini field cage covers the inside face
765 of the plug in order to reduce drift field distortions. It is filled with nitrogen at a pressure of
766 1.3 bar to balance against the hydrostatic pressure of LAr at its positioned height. The beam
767 plug can be seen on the right side of Figure 4.3. In addition to the beam plug, the cryostat
768 warm structure and insulation are modified to further reduce upstream interactions [30].

769 4.2.3 Cold Electronics

770 Each APA has a total of 2560 sense wires, resulting in a total of 15,360 channels to be
771 read out. 20 Front End Mother Boards (FEMBs) are located directly on top of each APA
772 and within the LAr to read out the sense wires. By being placed close to the wires, the
773 capacitance of each channel is reduced, thus reducing the noise recorded by the electronics.
774 The CE collect the signals from the APA wires, then amplify, shape, and digitize them
775 before transmitting them to Warm Interface Boards (WIBs). These interface electronics
776 then handle transmitting these signals to the DAQ.

777 The FEMBs consist of an analog motherboard containing eight 16-channel analog Front-
778 End (FE) ASICs that provide the amplification and shaping of the signals, and eight 16-
779 channel Analog to Digital Converter (ADC) ASICs. These ASICs are both custom circuits
780 designed by Brookhaven National Laboratory (BNL) [36]. In addition to the analog mother-
781 board is a mezzanine card containing a commercial Altera Cyclone IV FPGA which provides
782 clock and control signals to the two sets of ASICs. The FEMB layout can be seen in Figure
783 4.5.

784 The FE ASICs provided amplification with a programmable gain of 4.7, 7.8, 14, and 25
785 mV/fC and a 5th-order anti-aliasing shaper with programmable peaking time of 0.5, 1, 2, and
786 3 μ s. It also included options for enabling AC coupling, selectable baseline adjustment for
787 operating at 200 mV for unipolar pulses on the collection plane or 900mV for bipolar pulses on
788 the induction planes, and a selectable pre-amplified leakage current of either 100, 500, 1000,

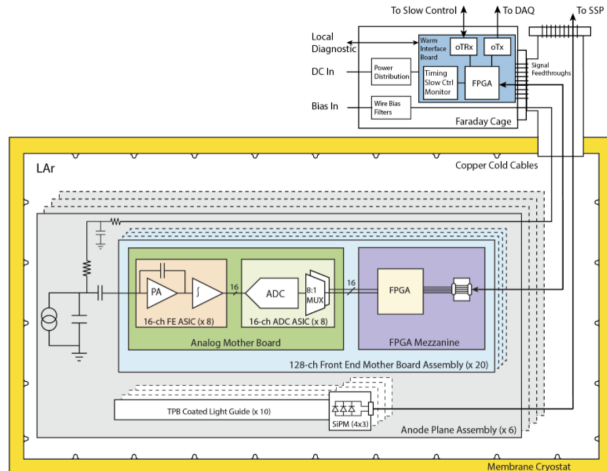


Figure 4.5: Diagram of the Cold Electronics in ProtoDUNE-SP [36].

789 or 5000 pA [36]. These ASICs also contained an internal, programmable pulse generator for
 790 electronics calibration. At normal running conditions, the FE ASIC gain is set to 14 mV/fC
 791 and the peaking time is set to 2 μ s for all channels [30]. At cryogenic temperature, the
 792 FE ASIC packaging puts stress on the ASIC chip causing a channel dependent non-uniform
 793 lowering by up to 150 mV of the 200 mV collection mode baseline [36]. In addition to this,
 794 large input charge caused the FE ASICs to saturate. Due to this, the baselines for both
 795 collection and induction plane channels were set to 900 mV [30].

796 The ADC ASICs have 16 12-bit digitizers operating at speeds up to 2 MHz and an 8:1
 797 multiplexing stage resulting in a pair of parallel serial readout lines that send output signals
 798 to the FPGA. At cryogenic temperature, the ADC ASIC suffered from an issue caused by
 799 failures in transistor matching. This effect is hard to simulate at LAr temperatures and is
 800 not present at room temperature. The mismatch between transistors effected the transition
 801 between the six most significant bits and six least significant bits in the ADC’s “domino”
 802 architecture, causing the ADC output to prefer 0 and multiples of 63 in the dynamic range
 803 of the ADC [36]. This issue, referred to as “sticky codes,” was corrected for after data-taking
 804 and will be further discussed in Section 4.3.

805 The signals from the ADC ASICs are collected by the FEMB’s FPGA, which further

806 serializes the 16 pairs of data streams into four 1.25 Gbps links to the WIBs. The FPGA also
807 provides a calibration pulse to the FE ASICs as a cross-check for electronics calibration [36].

808 The WIBs serve as the interface between the CE and DAQ, and are each controlled by
809 an Altera Arria V GT FPGA. Data cables from all FEMBs on a given APA feed through
810 a signal flange to a Warm Interface Electronics Crate (WIEC). The WIECs each contain
811 one Power and Timing Card (PTC) which is connected to both a 48 V power supply and
812 the detector timing system via a bidirectional fiber optical link. The PTC is connected
813 to a Power and Timing Backplane (PTB) also housed in the WIEC. The PTB steps down
814 the power and then fans out the power and clock signals from the PTC to five WIBs also
815 contained in the WIEC. Each WIB distributes power to and controls up to 4 FEMBs. The
816 WIB FPGA reorganizes and transmits FEMB data over fiber optical links to the DAQ. It
817 also includes a real-time digital diagnostic readout on a Gb Ethernet link and an on-board
818 component that can provide independent clocking to the FEMBs. These two components
819 allowed for installation and checkout tests to be performed on the FEMBs before they were
820 connected to the timing system and DAQ [36].

821 **4.2.4 Photon Detectors**

822 To collect scintillation light produced by charged particles in the LAr, 10 bar-shaped photon
823 detectors 8.6 cm in height, 2.2 m in length, and 0.6 cm thick were embedded in each APA
824 frame. Three different designs of photon detection technology were used in order to test
825 options for use in DUNE's far detector modules. In each, the ~ 128 nm scintillation photons
826 were converted into visible light using wavelength shifters. This visible light is trapped within
827 the photon detectors and eventually collected by an array of silicon photomultipliers [30].

828 **4.2.5 Cosmic Ray Tagger**

829 Located upstream and downstream (relative to z /beam direction) of the ProtoDUNE-SP
830 cryostat is a cosmic ray tagger (CRT) used to provide triggers from cosmogenic muons. The

831 CRT is formed of scintillation counters recycled from the outer veto of the Double Chooz
832 experiment that coarsely measure the x and y position of cosmic muons which pass through
833 it. Coincidence hits registered upstream and downstream of the detector can be used to
834 form tracks that can then be matched to reconstructed tracks in the TPC and provide
835 calibration [30].

836 4.2.6 Data Acquisition, Timing, Triggering

837 The DAQ reads in data from the TPC, photon detectors, and CRT. Two readout solutions
838 were employed for the TPC as tests for the DUNE far detector readout: RCE [37] and
839 FELIX [38]. During the beam run, one APA (located on the $x > 0$ side) used FELIX, while
840 the other 5 APAs used RCE. artDAQ [39] was used as the software framework that controlled
841 the data-flow including event building, configuration, and writing of data to disk [30].

842 The timing system provides a 50 MHz clock to all subsystems of the detector. It also
843 serves to distribute triggers created by the Central Trigger Board (CTB). The CTB is a hard-
844 ware triggering system that forms trigger words based on the status of individual subsystems
845 (CRT, photon detectors, beam instrumentation). These words are sent to the timing system
846 which ultimately makes the readout decisions. Various trigger conditions can be created by
847 creating requirements of subsystem statuses (active vs. inactive). When these requirements
848 are met, the CTB sends off its trigger words to the timing system, which then determines if
849 an event should be formed. If so, it issues the trigger to the DAQ and the various readout
850 systems [30]. Importantly, the CTB can create beam-on and beam-off triggers based on if
851 the beam instrumentation recorded a particle passing through the beam line. This way, TPC
852 events containing a beam particle can be easily identified and used for analysis.

853 Each triggered readout of the detector, also known as an “event,” consists of 3ms of data
854 taking: 6000 consecutive samples taken at a rate of 2MHz from each ADC. The event is
855 built from data taken in by the DAQ starting 250 μ s before the trigger time. This collects
856 signals from charge deposited in the detector before the trigger, but that arrive within the

857 time of the event. Coinciding data from the photon detectors and CRT are saved in the
858 output stream as well, and matching to beam instrumentation (which will be described in
859 Chapter 5) is done after data taking.

860 4.3 TPC Characterization

861 Before performing analysis, several data preparation steps are required to convert the
862 waveforms in units of ADC to units of charge, as well as to mitigate readout issues. The first
863 step is to determine the pedestal of each channel, as voltage offsets are introduced at the
864 input of the front end amplifiers and these vary on a channel-by-channel basis. Additionally,
865 for a given channel, the pedestal varies from one TPC event to the next. As such, the
866 pedestal is evaluated separately for each channel and each event [30]. The pedestals are
867 determined by finding the mean of all (typically) 6000 samples in an event for each channel.

868 For each channel in the event, its pedestal is subtracted from all ADC samples in the
869 waveform. This difference is then multiplied by the channel's gain. This gain g is determined
870 by using the 6-bit DAC included in the FE ASIC to inject a known amount of charge Q .
871 For the collection plane, the integral of the ADC signal over the pulse A is related to the
872 input charge as $Q = gA$. Special runs were taken where the DAC injected known amounts
873 of charge. For each charge setting, the mean of the ADC integral of the resulting waveforms
874 were determined. A line constrained to pass through 0 was fit to a set of these mean
875 values near charge inputs typical to operation (up to several overlapping Minimum Ionizing
876 Particles). The slope of this line is proportional to g for that channel [30]. Figure 4.6 shows
877 an example of this.

878 In addition to the gain calibration, readout issues are identified and mitigated. The
879 first readout issue is the aforementioned sticky codes issue. ADC values subject to sticking
880 as well as the channels which exhibit the issue were initially identified by scanning a few
881 waveforms and the pedestal histograms for every channel. The channels with particularly
882 prevalent sticky codes are identified, and the list of known sticky codes is used to mitigate the

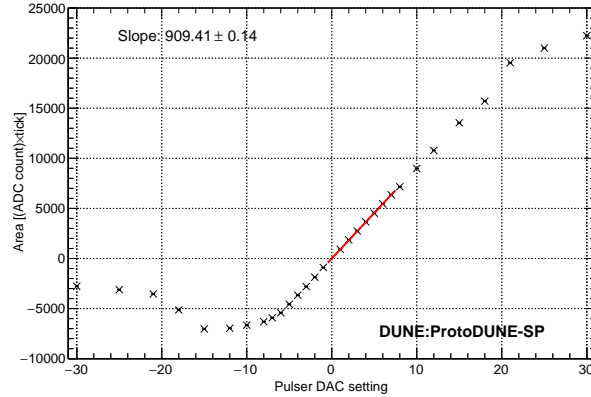


Figure 4.6: Example of gain determination for one channel in PDSP. The slope of the line divided by the charge level of each DAC step ($Q_s = 3.43\text{fC} = .4\text{ ke}$) gives the gain of the channel [30].

883 issue in less problematic channels. The mitigation works on these channels known to exhibit
 884 sticky code issues by replacing any ADC sample at a sticky code with a value taken from
 885 interpolation between the nearest non-sticky neighbors [30]. An example of this is shown in
 886 Figure 4.7.

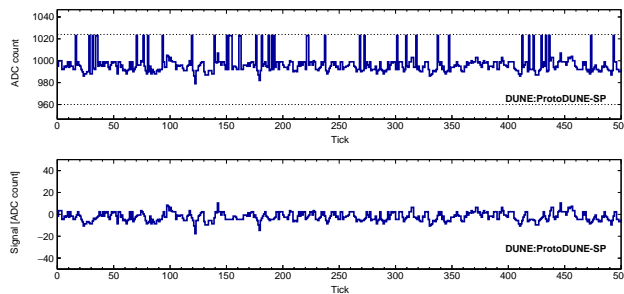


Figure 4.7: Example of ADC waveform before (top) and after (bottom) pedestal subtraction and sticky code mitigation. The spikes are samples which have stuck to the code represented by the upper horizontal dashed line. They are removed and replaced by interpolating to the nearest non-sticky neighbors [30].

887 In addition to the sticky code mitigation, preparations to remove tails resulting from AC
 888 coupling in the CE and correlated noise are also performed [30].

889 4.4 Event Reconstruction

890 After events are recorded, and an amount of data preparation is performed (described
891 previously in Section 4.3), reconstruction software builds up a description of what happened
892 during the event. This is done in a two step process: 1) hit finding which identifies local-
893 ized charge deposits on wires and 2) pattern recognition which separates collections of hits
894 into objects representing particle tracks and showers, and which also attempts to associate
895 particles together in a hierarchy representing series of interactions.

896 Ideally, charge depositions on the wires should form (possibly overlapping) Gaussian-
897 shaped signals when read out by the electronics. Thus, a hit-finding algorithm attempts to
898 identify these separate depositions of charge by fitting Gaussian peaks to the waveform in
899 a given wire. Each Gaussian peak thus represents one reconstructed hit or, in other terms,
900 a localized deposition of charge in the detector. An example of this is given in Figure 4.8,
901 where three hits have been reconstructed to the shown waveform.

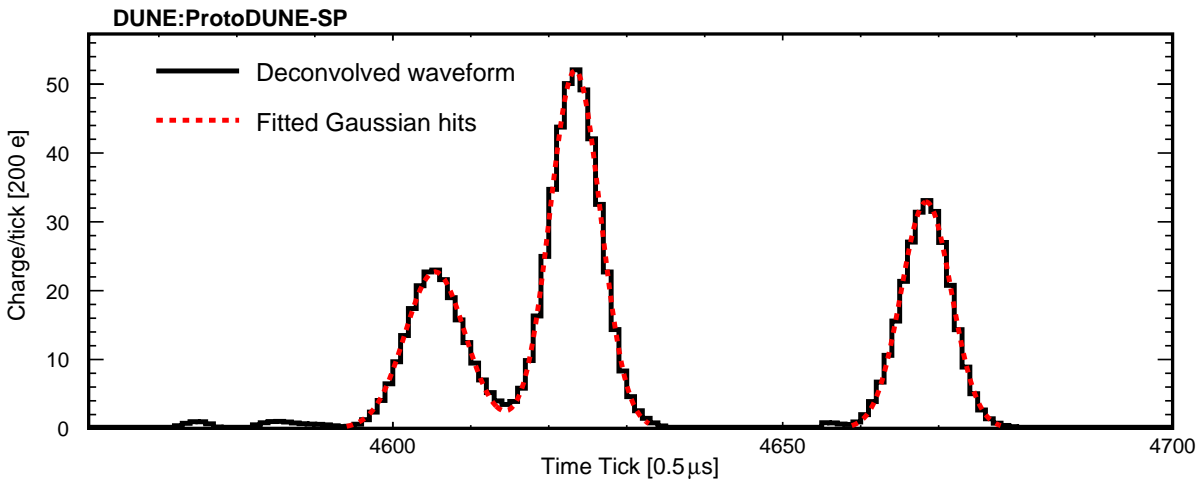


Figure 4.8: Example of three hits reconstructed to a single wire’s waveform [30].

902 Because induction wires are wrapped around the APA, charge on either side of the APA
903 can create a signal on a given wire. Thus, a disambiguation must be performed to determine
904 which side of the APA the signal came from. Collections one wire from each plane are formed

905 by identifying signals that arrived within a narrow time window. Sometimes, multiple pairs
906 of induction wires can be matched to the collection plain. To determine which ones were
907 truly paired, the algorithm attempts to minimize the difference in charge between that on the
908 collection plane wire and on the induction plane wires. Simulation shows that this assigns
909 >99% of hits to correct wire segments [30].

910 The second part of reconstruction is pattern recognition, which is performed by the Pan-
911 dora framework [40]. This software has been successfully used in other LArTPC experiments
912 such as MicroBooNE [41]. The first step in the pattern recognition is to perform two dimen-
913 sional clustering of the reconstructed hits in each view. It then attempts to match sets of
914 2D clusters between the views, with care taken to resolve ambiguities. Afterwards, 3D hits
915 are created. Then, in order to provide a detailed description of events, particle interaction
916 hierarchies are created. Pandora then attempts to pick out particles originating from the
917 beam line. All clusters are reconstructed first under a cosmic ray hypothesis. Clear cosmic
918 ray candidates are then identified and removed. After these cosmic particles are removed,
919 Pandora attempts to divide the detector into 3D regions containing all hits produced by a
920 given particle interaction. These regions could contain cosmic rays that were not previously
921 identified as such or particles that originated from the beam. Parallel reconstruction chains
922 (one for cosmic rays and the other for test-beam particles) are then performed on these
923 detector regions. The reconstruction for the beam particles is intended to resolve intricate
924 hierarchies of particles such as from hadronic interactions. After the dual reconstruction is
925 performed, a boosted-decision-tree algorithm tries to identify which region (if any) originated
926 from the incoming beam [30]. The full reconstructed hierarchy (links between parent and
927 child particles) in the beam region are made available for analysis. Further pattern recog-
928 nition tries to identify whether the reconstructed particles were track-like (such as pions,
929 protons, muons) or shower-like (electromagnetic showers from electrons or photons). Recon-
930 structed track and shower objects are created for the corresponding particles. These provide
931 information such as track length or shower direction (depending on the object) to users.

932 4.4.1 Hit Classification Using Machine Learning

933 In addition to the track/shower discrimination from Pandora, a machine learning-based clas-
934 sification was developed. A convolutional neural network (CNN) was trained to classify hits
935 into track-like, shower-like, empty, or Michel-like categories. The track-like category repre-
936 sents hits coming from particles like pions and muons, the shower-like category represents
937 hits from electron or photon showers, and the empty category represents hits resulting from
938 noise. The Michel-like category represents electrons which originate from the decay of muons
939 in the LAr, which appear different to electron showers. The output of the network is a set of
940 scores representing how similar to each category the hit appears. The Michel-like category
941 can overlap with the track-like and shower-like categories, and was not used for this analy-
942 sis. The other three category scores are constrained to sum to one such that the hit can be
943 classified as only one of these categories (that with the highest score).

944 The network uses as input 48×48 pixel² images created from wire readout data with the
945 hit in question at the center. Each pixel is filled with the ADC value from the readout
946 data. One axis of the image represents the wire which recorded the hit, while the other
947 axis is the time coordinate (which has been downsampled by taking an average over time
948 samples). The readout data used as input has been prepared according to the procedure
949 described in 4.3. MC simulation was used to train the network by identifying whether the
950 hit was due to charge deposited by a particle (or was created by noise) and what type of
951 particle created the hit. Further information on network architecture and training can be
952 found in Reference [42]. The analysis presented in this thesis utilized the CNN scores as an
953 alternate track/shower discrimination technique. Scores for the full reconstructed particle
954 were calculated by averaging over all hits in the particle. Cuts can be placed on these average
955 scores to categorize full particles as tracks or showers. The use of this within this analysis is
956 described further in Chapter 6

957 4.5 Detector Calibration

958 In order to conduct useful physics analyses such as the measurement presented here, the
959 relationship between deposited energy in the detector and the response of the detector must
960 be determined. Several effects that must be taken into account have already been described
961 in Sections 4.1.3, 4.1.1, and 4.1.2. This section serves to describe the procedures taken to
962 calibrate for these effects.

963 4.5.1 Space Charge Effect in ProtoDUNE-SP

964 As previously described, the steady flux of cosmic rays produces a buildup of charge from the
965 slowly drifting Ar^{1+} ions produced by ionization. This so-called Space Charge Effect leads
966 to a persistent distortions of the drift field. These alter the drift paths of ionization electrons
967 and also affect the amount of prompt charge recombination, resulting in spacial distortions
968 of reconstructed tracks and modified reconstructed dE/dx of tracks. The spatial distortions
969 of reconstructed tracks is evident in Figure 4.9, where the end points of cathode-crossing
970 cosmic rays are pulled inward from the edges of the detector (the dashed lines).

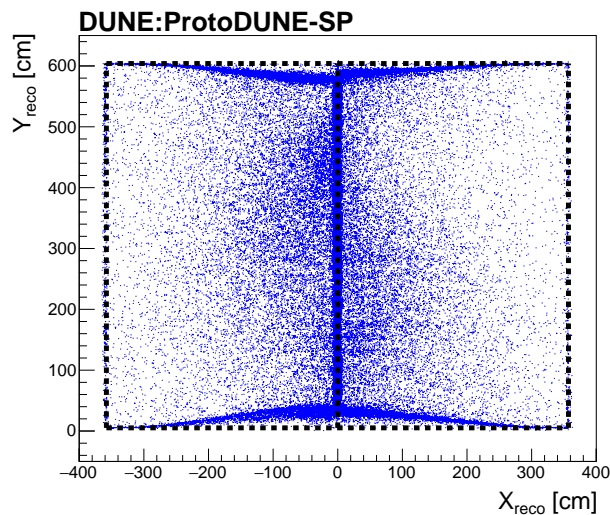


Figure 4.9: Projection of reconstructed track end points from cathode-crossing cosmic ray muons. The presence of SCE causes the end points to deviate from the boundaries of the TPC volumes represented by the dashed lines [30].

971 Figure 4.10 shows the distortions normal to four of the detector faces for events piercing
 972 the respective face within data events. These provide the magnitude of spatial distortions
 973 at these points of the detector in data events. A simulation of SCE was developed for
 974 ProtoDUNE-SP. This is shown in Figure 4.11, which is analogous to Figure 4.10. Data-MC
 975 discrepancies can be seen here that possibly stem from incorrect values of the Ar^{1+} drift
 976 velocity (amounting to a different amount of accumulated charge) and/or unsimulated flow
 of the liquid argon.

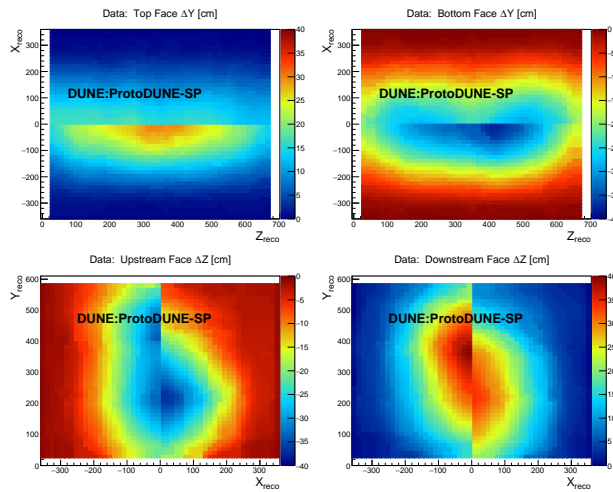


Figure 4.10: Spatial distortions normal to four detector faces from data events. Top: upstream & downstream relative to the z-direction. Bottom: Upper & lower faces relative to the y-direction. The reconstructed location of the end points of cathode-crossing tracks that pierce through the respective face show the distortions perpendicular to that face at the reconstructed 2D location [30].

977

978 In order to overcome the inability of the simulation to reproduce the SCE seen in data, a
 979 data-driven simulation of space charge was implemented. This consisted of creating a set of
 980 both spatial and electric field distortion maps to modify the nominal simulated distortions.
 981 These maps can also be used to correct for SCE in both data and MC by recovering the
 982 original positions and also accounting for the modified electric field. These are created as
 983 follows:

984 1. The ratio of the data to simulation map are taken for each of the relevant faces of the

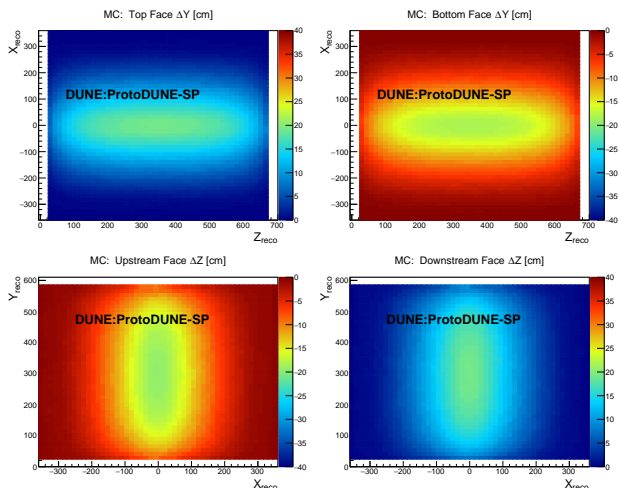


Figure 4.11: Spatial distortions normal to four detector faces from MC events. Top: upstream & downstream relative to the z -direction. Bottom: Upper & lower faces relative to the y -direction. The reconstructed location of the end points of cathode-crossing tracks that pierce through the respective face show the distortions perpendicular to that face at the reconstructed 2D location [30].

- 985 detector. This produces a 2D map of scale factors at each of these faces.
- 986 2. Spatial distortions in the y -direction are calculated by linearly interpolating the scale
 987 factor maps between the top and bottom faces. The same is done for z -direction dis-
 988 tortions by interpolating the scale factor maps between the upstream and downstream
 989 faces. The x -direction distortions are then taken as the average between the distortions
 990 in y and z . This creates a 3D map with scale factors in all three directions. These are
 991 used to rescale the magnitudes of the spatial distortions maps.
- 992 3. The resulting distortions in each 3D map are then reversed in order to form maps that
 993 can be used to correct for the spatial distortions in both data and MC. The correction
 994 repositions the reconstructed ionization charge depositions to their original locations.
- 995 4. The gradient of the spatial distortion along the local drift direction (determined from
 996 the reversed maps) and the known drift velocity are used to form 3D electric field
 997 distortion maps.

998 These data-driven maps are used to modify the reconstructed position of ionization charge
999 in simulation as well as to improve the prediction of prompt recombination effects [30].

1000 4.5.2 Electron Lifetime

1001 As discussed in Section 4.1.2, impurities in the liquid argon can capture drifting electrons
1002 before they reach the instrumentation wires. This reduces the amount of charge reaching
1003 the collection plane wires and is measured as an exponential decay as a function of drift time
1004 as in Equation 4.4, where τ is the drift electron lifetime. A larger τ corresponds to a higher
1005 liquid argon purity.

1006 The electron lifetime can be measured by fitting the dQ/dx of cosmic ray collection plane
1007 hits as a function of drift time. To do this, cosmic rays that pass through the CRT and the
1008 front and back faces of the TPC were selected. The CRT was used to measure the initial
1009 time t_0 at which the track traveled through the TPC. The difference between the time the
1010 hit collected on the wire and t_0 was used as the drift time. The most probable value of
1011 dQ/dx for hits in slices of 100 μ s drift times was fit according to Equation 4.5 to extract the
1012 lifetime. Two example fits, taken at the beginning and end of the beam data run are shown
1013 in Figure 4.12. The later data shows a higher lifetime resulting from higher purity [30].

$$\frac{dQ(t)_{\text{MPV}}}{dx} = \frac{dQ_{0,\text{MPV}}}{dx} \exp(-(t_{\text{hit}} - t_0)/\tau) \quad (4.5)$$

1014 4.5.3 Energy Calibration

1015 Reconstructed dQ/dx is affected by electronics gain variations, SCE, and attenuation. Previ-
1016 ous sections describe the calibrations for these. Additional effects have also been calibrated
1017 out via a two-step process laid out here: first to equalize the detector response (using a
1018 sample of throughgoing cosmic rays), then with a determination of the absolute energy scale
1019 (using a sample of stopping cosmic rays).

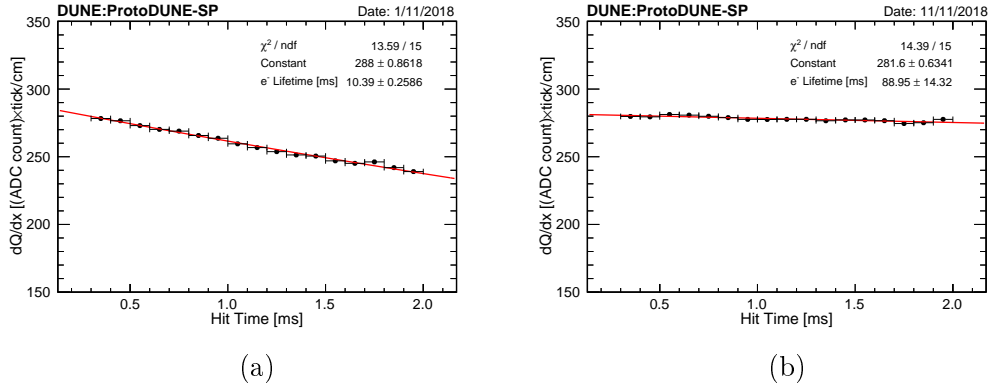


Figure 4.12: Fits to the drift electron lifetime τ for data collected at two different periods of time. Left is an earlier period with a lower period and shows a lower lifetime (10.39 ± 0.2586 ms) compared to the right (88.95 ± 14.32 ms) [30]

1020 The equalization step accounts for nonuniformities from various effects that depend sep-
 1021 arately on x and $y - z$ positions of the hit. Effects that depend on $y - z$ position include
 1022 non-uniform wire response from nearby dead channels, detector features such as the electron
 1023 diverters, and transverse diffusion. This portion of the equalization step is done separately
 1024 for each half of the detector on either side of the central cathode and as a function of y and
 1025 z . The median dQ/dx value of hits in a given $y - z$ bin is determined and compared to the
 1026 median dQ/dx value of the half of the detector wherein the hit lies ($x > 0$ or $x < 0$) to
 1027 obtain a correction factor defined as such:

$$C(y, z) = \frac{(dQ/dx)_{YZ}^G}{(dQ/dx)_{YZ}^L} \quad (4.6)$$

1028 where the numerator is the global median dQ/dx value on that side of the detector and the
 1029 denominator is the median value on that $y - z$ bin.

1030 Following this, effects that depend on x position such as longitudinal diffusion are equal-
 1031 ized. Similar to the corrections in the $y - z$ plane, the median dQ/dx value of hits in an x
 1032 bin are compared to the median value of all hits in the detector. This produces a correction
 1033 factor that depends on x position as such defined in Equation 4.7.

$$C(y, z) = \frac{(dQ/dx)_{YZ}^G}{(dQ/dx)_{YZ}^L} \quad (4.7)$$

1034 Finally, the dQ/dx values are normalized to the average value at the two anodes using
 1035 the following factor:

$$N_Q = \frac{(dQ/dx)^A}{(dQ/dx)^G} \quad (4.8)$$

1036 where the numerator is the average of the mean values at either anode, and the denominator
 1037 is the mean value over the whole TPC. Thus, the dQ/dx of every hit in an event is equalized
 1038 according to Equation 4.9.

$$(dQ/dx)_C = N_Q C(y, z) C(x) (dQ/dx) \quad (4.9)$$

1039 Next the measured dQ/dx must be translated to the energy loss of the particle per unit
 1040 length dE/dx using a sample of cosmic muons that stop in the detector. dQ/dx values in
 1041 the minimum ionizing region (120 to 200 cm from the end of the track) are converted to
 1042 dE/dx using Equation 4.10 from the modified Box model [31] and fit to values predicted
 1043 by Landau-Vavilon theory [43] as a function of residual range (the distance along the track
 1044 from the hit to the end of the track).

$$\frac{dE}{dx} = \left(\exp \left(\frac{(dQ/dx)_C \beta' W_{\text{ion}}}{C_{\text{cal}} \rho \mathcal{E}} - \alpha \right) \right) \left(\frac{\rho \mathcal{E}}{\beta'} \right) \quad (4.10)$$

1045 In Equation 4.10, W_{ion} is the amount of energy required to ionize an Argon atom (equal
 1046 to 23.6×10^{-6} MeV/electron), ρ is the density of liquid argon at ProtoDUNE-SP operating
 1047 temperature (equal to 1.38 g/cm^3), \mathcal{E} is the local electric field at the location of the hit, α and
 1048 β' are modified Box model parameters and were measured by ArgoNeuT with values of 0.93
 1049 and $0.212 \text{ (kV/cm)(g/cm}^2\text{)/MeV}$ respectively [31]. Finally, C_{cal} is a calibration constant
 1050 that accounts for electronics gain and ADC conversion, and corrects for any residual effects
 1051 not explicitly calibrated previously and is the parameter of interest in the fit [30].

1052 The normalization factor N_Q , equalization maps $C(y, z)$ and $C(x)$, and calibration con-
 1053 stant C_{cal} are measured separately for MC and each run of data, and are applied during
 1054 analysis when extracting the values of dE/dx for each hit considered.

1055 4.6 Monte Carlo Simulation

1056 The simulation of test beam events in the TPC begins with the simulation of test beam
1057 particles generated within the beam line. A dedicated Geant4 [44] simulation of the beam
1058 line transports particles from their production point toward the face of the ProtoDUNE-SP
1059 TPC. More details can be found in Reference [45]. The rest of the PDSP simulation chain is
1060 based in the analysis framework LArSoft [46]. The beam line simulation results are passed
1061 to an event generator module that creates particles to be simulated by Geant4. The events
1062 are created when a “primary” particle (such as a π^+) travels through two triggering planes
1063 and reaches the outside of the PDSP cryostat structure. Additionally, checks are performed
1064 for when particles interact or decay (if applicable) in the beam line such that events are
1065 also created if some downstream particle (for example a μ^+ from a π^+ decay) reaches the
1066 cryostat. Without this check, the rate of test beam muons was severely underpredicted
1067 by the event generator in early simulation productions. The set of simulation used in this
1068 analysis included this hierarchy check. Each event created by this event generator is assigned
1069 a primary particle: either the original particle or the last-extant particle which reached the
1070 cryostat structure (i.e. the μ^+ described above). These primary particles serve as the main
1071 particles considered in the analysis. Other particles originating from the beam line which are
1072 “in time” with the beam are passed on to the next stage of the simulation. These additional
1073 particles are added in if they are within 4.5ms of the primary particle, similar to what can
1074 occur in events in data. Cosmic-ray particles as simulated by CORSIKA [47] are overlaid on
1075 the event and passed on to the Geant4 step as well.

1076 All particles generated by the beam-based event generator and the overlaid cosmics are
1077 then passed to Geant4 to simulate their transport through the detector. It also facilitates
1078 the interaction of hadrons with the detector material (via the Bertini Cascade Model [48]),
1079 thus serving as the signal interaction model of the analysis. The full geometry of the detector
1080 is considered, allowing for particles to interact and lose energy within the uninstrumented
1081 portion of the detector geometry (i.e. the steel cryostat structure, insulation, etc.). As

1082 charged particles travel through the LAr portion of the detector, ionization is created which
1083 is then passed on to the drift simulation step of the simulation.

1084 The drift simulation transports the ionization electrons produced during the Geant4
1085 simulation stage along field lines toward the wire planes. The nominal electric field map
1086 used within the simulation is distorted according to the data-driven SCE maps discussed
1087 earlier in Section [4.5.1](#). The full electronics response to the ionization drift and collection
1088 onto the wires is simulated, creating waveforms which are then passed to the reconstruction
1089 chain described earlier.

CHAPTER 5

PROTODUNE-SP BEAM LINE

1090

1091 Test beam particles are delivered to ProtoDUNE-SP from an extension of the existing H4
1092 beam line in the CERN North Area. This beam line is known as the H4-VLE (very low
1093 energy) beam line as it supplies particles (π^+ , μ^+ , e^+ , K^+ , and p) in the momentum range
1094 $0.3 - 7 \text{ GeV}/c$. Within the North Area Secondary Beam facility, protons from the CERN
1095 Super Proton Synchrotron (SPS) impinge on a beryllium target to create a beam of secondary
1096 particles. These particles are transported through the H4 beam line before impinging on a
1097 secondary target to create the test beam for ProtoDUNE-SP. These test beam particles are
1098 momentum-selected¹ and transported through the H4-VLE beam line toward ProtoDUNE-
1099 SP.

1100 5.1 Beam Line Instrumentation

1101 The H4-VLE beam line is instrumented with a set of various devices to aid in particle
1102 identification (PID), momentum reconstruction, and tracking the beam. The layout of the
1103 beam line is shown in Figure 5.1. The instrumentation consists of scintillating planes (XBTF)
1104 for triggering and time of flight (TOF) measurements; scintillating fiber monitors (XBPF) for
1105 profiling, tracking, and momentum reconstruction; and Cherenkov detectors (XCET) as part
1106 of the PID process. Throughout the beam line are bending magnets which direct the beam
1107 toward ProtoDUNE-SP – with one also being used as part of a momentum spectrometer.

1108 5.1.1 Fiber Monitors

1109 The XBPF profile monitors [49] are comprised of a set of 192 square scintillating fibers of
1110 width 1 mm set side-by-side to provide a measurement of a beam particle's position in one
1111 direction. Two can be placed in perpendicular orientations to provide a 2D measurement of

¹Nominal momentum settings consist of 0.3, 0.5, 1, 2, 3, 6, and 7 GeV/c

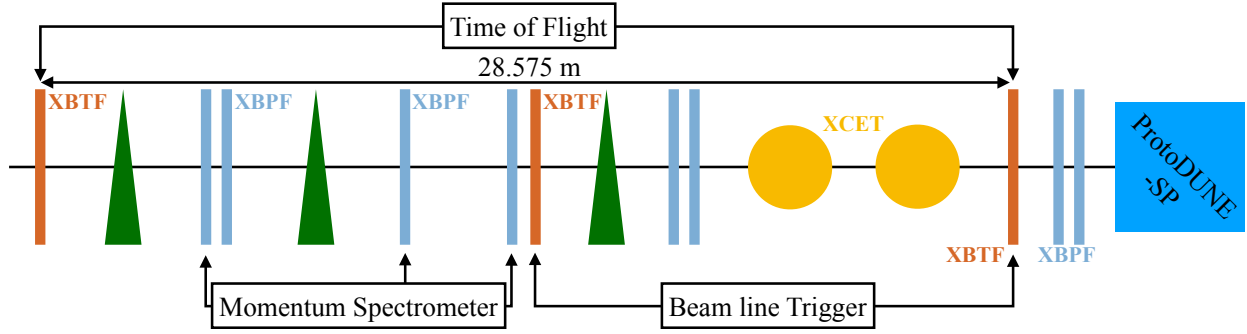


Figure 5.1: Diagram of the H4-VLE beam line instrumentation layout. XBTF (orange lines) are scintillating planes used for triggering and TOF measurement; XBPF (blue lines) are scintillating fiber monitors used for tracking and momentum reconstruction; XCET (orange circles) are Cherenkov detectors used for PID (sometimes in conjunction with the TOF); the green triangles are bending magnets throughout the beam line.

1112 the particle's position. Each fiber is connected to an individual Hamamatsu S13360-130
 1113 silicon photomultiplier (SiPM) on one end². Figure 5.2 shows a photograph of a prototype
 1114 XBPF module taken from [49]. Further discussions of these devices and their readout are
 1115 found in this reference as well as Reference [45]. The XBPF data was packaged such that,
 1116 for each trigger in the beam line, the statuses (on/off) of the 192 fibers were separated into
 1117 six 32-bit words. Two examples of this decoding is given in Figure 5.3.

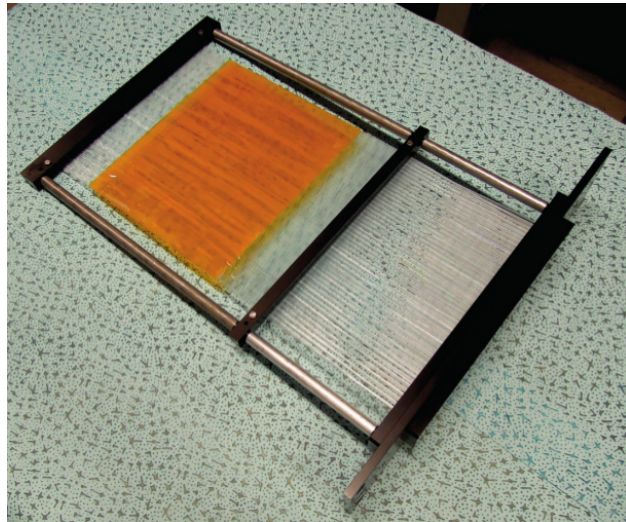


Figure 5.2: XBPF module. Taken from Reference [49].

²On the other end of the set of fibers is an aluminized mylar mirror.

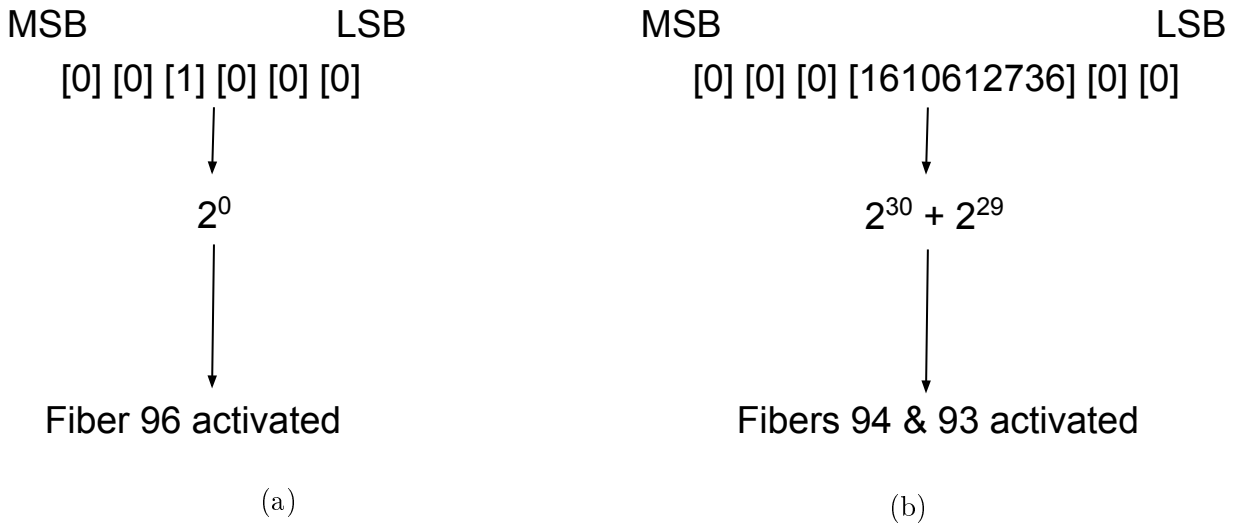


Figure 5.3: Two examples of XBPF data decoding. The most significant bit (MSB) and least significant bit (LSB) are labeled at the top of each example. a) The 0th fiber in the fourth 32-bit word is active. Thus the active fiber is $(0 + 3 * 32) = 96$. b) Two fibers are active: the 30th and 29th fibers in the third 32-bit word. Thus fibers $(30 + 2 * 32) = 94$ and $(29 + 2 * 32) = 93$ are active.

1118 The last two sets of XBPF devices (shown immediately before the XCET devices and
 1119 after the last XBTF plane in Figure 5.1) were used for tracking the particle as it entered
 1120 into the TPC. 2D positions were reconstructed in both sets of XBPFs and used to create
 1121 a trajectory between these points along the beam direction. This trajectory was further
 1122 projected to the face of the active TPC to give the reconstructed position at the beam
 1123 window.

1124 These projected trajectories were used within analysis to cut out events considered as
 1125 background to our pion sample. The difference in position between this reconstructed beam
 1126 point and the start of the reconstructed TPC track, as well as the angle between the recon-
 1127 structed beam trajectory and the starting angle of the reconstructed TPC track, were used
 1128 to exclude various backgrounds (i.e. cosmic or particles from “upstream” interactions before
 1129 the start of the active TPC volume).

1130 5.1.1.1 Issues with XBPFs

1131 In Winter 2019, two issues were identified within the data obtained by the XBPF during the
1132 initial beam run. In the first, the rate of fiber activations for the upper half of the fibers in
1133 the first XBPF was higher than the lower half. This can be seen in Figure 5.4 where the
1134 number of activations for each fiber for each selected event in the first XBPF are plotted. In
1135 talks with the device experts, this was determined to be caused by a configuration issue in
1136 the ASICs controlling the readout of this XBPF. This amounted to a higher efficiency in the
1137 upper half of fibers in this device. However, this was not an issue in ProtoDUNE-SP data
1138 analysis as this effect was suppressed by the lower trigger rate of the ProtoDUNE-SP detector
1139 compared to the trigger rate of the beam line (a subset of beam line particles triggered the
1140 detector).

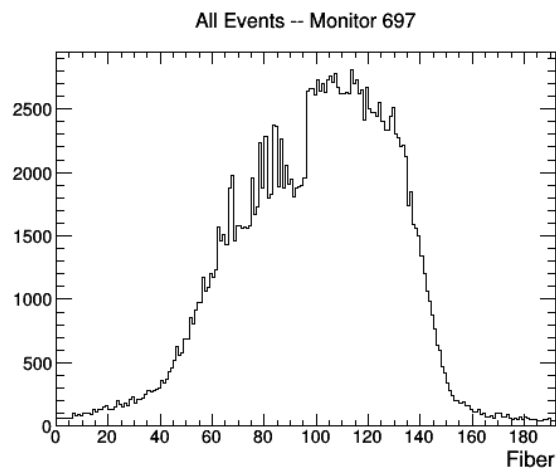


Figure 5.4: Active fibers in the first XBPF device from every event from a 1 GeV/ c run. Note, multiple fibers can be active in any one event. The jump in rate at fiber 96 is due to a configuration problem in the readout electronics.

1141 The second issue identified was due to a bug in the software controlling the data acquisi-
1142 tion for the XBPFs, and occurred in all XBPF devices. In this issue, systematically repeated
1143 hits were being recorded in the last 64 fibers of each XBPF. This can be seen in Figure
1144 5.5, where a bump is present starting near fiber 128 of the second XBPF device. Figure 5.6
1145 highlights this issue, as it shows the number of times a fiber was activated in two subsequent

1146 events. A large spike in this rate can be seen starting at fiber 128. In discussions with the
 1147 device expert, this was determined to be due to a software bug, in which the data in the
 1148 last two words was not being cleared between events in the XBPF devices. This caused a
 1149 “hangover” in the apparent activation of fibers in these two words. This resulted in extra
 1150 reconstructed hits seen during analysis, which led to ambiguity in the reconstructed momen-
 1151 tum and incident tracks. An attempt to mitigate this was implemented by simply scanning
 1152 the last two words of each event for repeated fibers, and then masking the repeated fibers
 1153 (in the second event). The results for this are shown in Figure 5.7

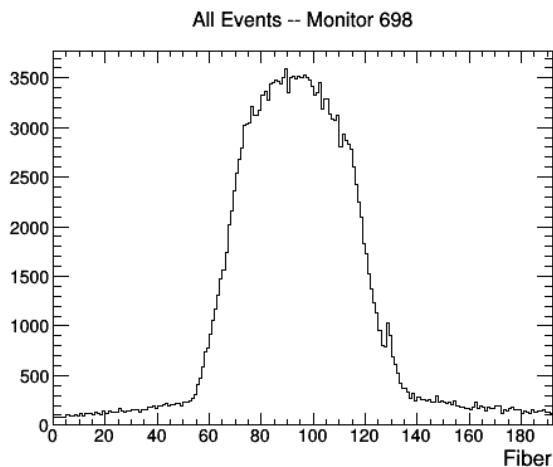


Figure 5.5: Active fibers in the second XBPF device from every event from a 1 GeV/c run. Note, multiple fibers can be active in any one event. A small bump can be seen starting around fiber 128.

1154 5.1.2 Momentum Reconstruction Using XBPFs

1155 Within Figure 5.1, three XBPFs are labeled as “Momentum Spectrometer.” Coincident
 1156 signals in these three monitors were used to measure the deflection of the test beam particle
 1157 by the bending magnet which the monitors surround. The angle of deflection is then used,
 1158 along with the known magnetic field, to reconstruct the particle’s momentum. A diagram of
 1159 this measurement technique is shown in Figure 5.8. The lateral position (χ_1, χ_2, χ_3) of the
 1160 particle – given by the activated fiber in each of the three XBPFs – is used with the known

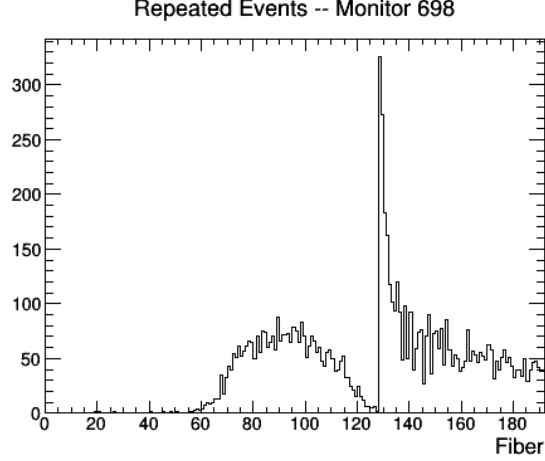


Figure 5.6: Rate of repeated fiber activations in the second XBPF device from every event from a 1 GeV/ c run. The large jump at fiber 128 highlights the issue.

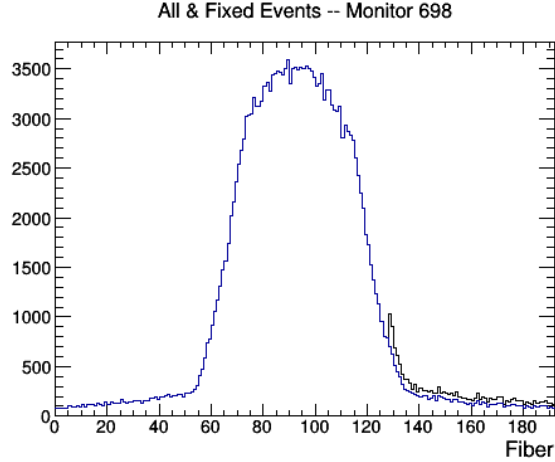


Figure 5.7: Active fibers in the second XBPF device from every event from a 1 GeV/ c run before (black) and after (blue) the mitigation procedure.

1161 distances between each monitor (L_1 , L_2 , L_3) in Equations 5.1 and 5.2 to reconstruct the
 1162 momentum. In these, $M \equiv \alpha + \chi_1$, $\alpha = \frac{\chi_3 L_2 - \chi_2 L_3}{L_3 - L_2} \cos \theta_0$, $\Delta L \equiv L_3 - L_2$, and $\Delta \chi \equiv \chi_2 - \chi_3$.
 1163 θ_0 is the nominal bending angle of the beam and is equal to 120.003 mrad [45]. This
 1164 measurement has a nominal 2% resolution according to Monte Carlo studies [45].

$$\cos \theta = \frac{M[\Delta L \tan \theta_0 + \Delta \chi \cos \theta_0] + L_1 \Delta L}{\sqrt{[M^2 + L_1^2][(\Delta L \tan \theta_0 + \Delta \chi \cos \theta_0)^2 + \Delta L^2]}} \quad (5.1)$$

$$p = \frac{299.7924}{\theta} \times \int_0^{L_{\text{mag}}} (Bdl) \quad (5.2)$$

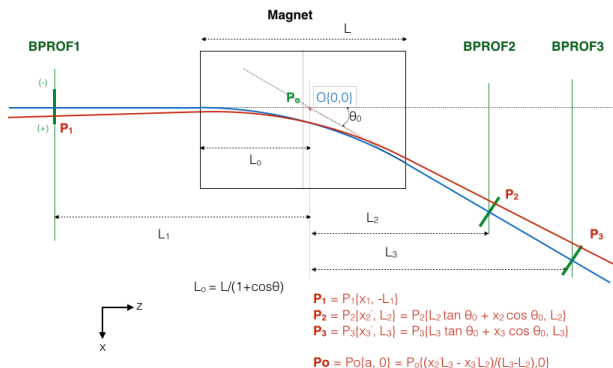


Figure 5.8: Momentum Spectrometer technique

1165 Shortly after commissioning, a 5% offset in the reconstructed momentum was observed,
 1166 and this was determined to originate from a bulk shift of the fibers in the third profiler of
 1167 the spectrometer. Monte Carlo studies determined the fiber shift to be 1.45 ± 0.18 mm in the
 1168 plane perpendicular to the beam. This was used as a systematic uncertainty within this
 1169 analysis and will be discussed further in Chapter 8.

1170 5.1.3 Scintillating Planes

1171 The XBTF scintillating planes are of similar design to the XBPF – sets of 192 fibers arranged
 1172 side-by-side and set perpendicular to the beam direction – but without individual readout
 1173 of the fibers. Instead, the fibers are bundled into two groups, which are read out by two
 1174 separate Hamamatsu H11934-200 photomultiplier tubes (PMTs). Figure 5.9 shows the
 1175 bundled nature of the XBTF fibers.

1176 The first and third (last) XBTFs – as shown in Figure 5.1 – are used for measure the TOF
 1177 of the test beam particle over a distance of 28.575 m. The second (middle) and last XBTFs
 1178 are used as a trigger for the rest of the beam line instrumentation as well as a prerequisite
 1179 for triggering beam-type events within the ProtoDUNE-SP detector.

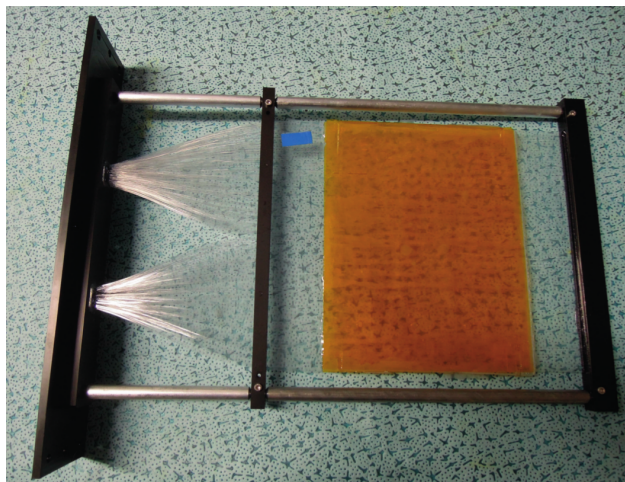


Figure 5.9: XBTF module. The bundling of the two sets of fibers can be seen on the left [49].

1180 5.1.3.1 Issue with XBTFs

1181 A ~ 4 ns “jitter” can be seen in the left plot of Figure 5.11b where a small bump in the TOF
1182 distribution exists around 100 ns. The cause for the issue was never identified. Furthermore,
1183 it has little effect on the analysis, due to the cuts used for PID (see below).

1184 5.1.4 Cherenkov Devices

1185 The two Cherenkov devices each consist of a 1.9 m long tube filled with the radiator gas (CO₂)
1186 followed by a stainless steel enclosure. This enclosure houses a PMT at the bottom to collect
1187 the Cherenkov light and a curved mirror to guide Cherenkov light toward the PMT. The fill-
1188 pressure of the two devices were set to two different values to allow for discrimination between
1189 certain particle types. Figure 5.10 [45] shows the Cherenkov threshold pressure of CO₂ at
1190 various momenta for different particle types, as well as the maximum possible pressure value
1191 for the two XCET devices. Consider an example setup at 3 GeV/ c momentum. One device
1192 can be set above the electron threshold but below the μ/π thresholds, while the other can be
1193 set above the μ/π threshold but below the K/p threshold in order to distinguish positrons,
1194 muons/pions, and kaons/protons. Further discussion of Cherenkov devices within the beam

1195 line PID algorithm will be given in the following section.

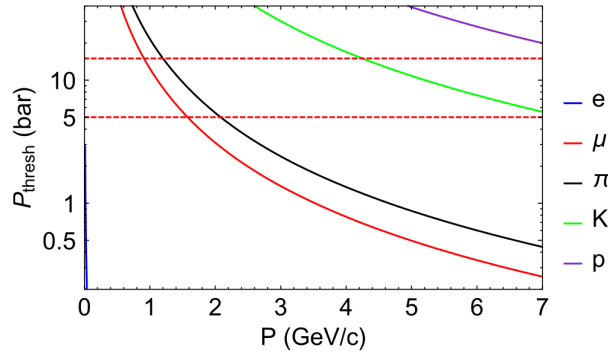


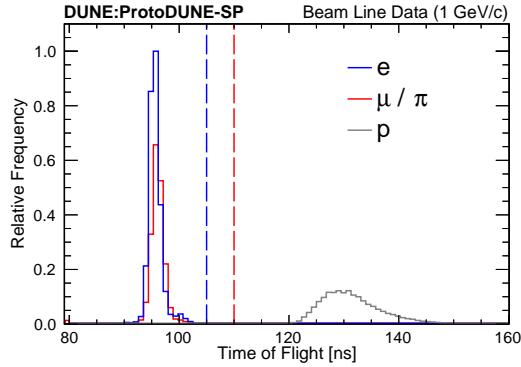
Figure 5.10: CO2 Cherenkov threshold pressures across ProtoDUNE's beam momentum range for the various particles present in the beam line. The dashed red lines show the maximum pressures for the two Cherenkov devices present in the beam line. Taken from Reference [45].

1196 5.2 Beam Line PID

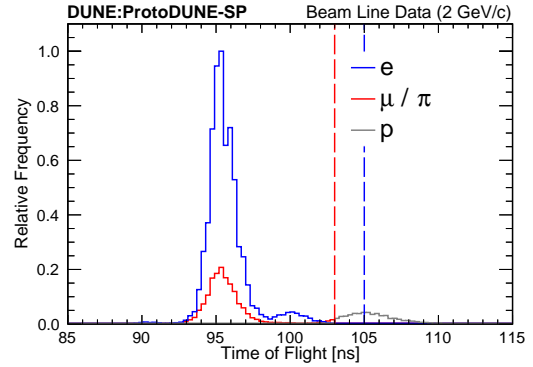
1197 As mentioned above, the Cherenkov devices and TOF as measured by the XBTFs were
1198 used for PID of the beam line particles. Table 5.1 shows the conditions of the Cherenkov
1199 devices and TOF value used for the PID algorithm across the various nominal momentum
1200 settings. As shown in this table, for nominal momenta below 3 GeV/c, one Cherenkov device
1201 is used to distinguish e from the other particles, and the TOF is then used to distinguish
1202 μ/π from p . At 3 GeV/c, both Cherenkov devices are used to separate e , μ/π , and K/p .
1203 Finally, at 6 – 7 GeV/c, the two Cherenkov devices are used to separate $e/\mu/\pi$, K , and p .
1204 Figure 5.11 demonstrates this for the various beam momentum settings.

		Momentum (GeV/c)			
		1	2	3	6 - 7
e	TOF (ns)	0, 105	0, 105	–	–
	Low-p Status	1	1	1	1
	High-p Status	–	–	1	1
μ / π	TOF (ns)	0, 110	0, 103	–	–
	Low-p Status	0	0	0	1
	High-p Status	–	–	1	1
K	TOF (ns)	–	–	–	–
	Low-p Status	–	–	0	0
	High-p Status	–	–	0	1
p	TOF (ns)	110, 160	103, 160	–	–
	Low-p Status	0	0	0	0
	High-p Status	–	–	0	0

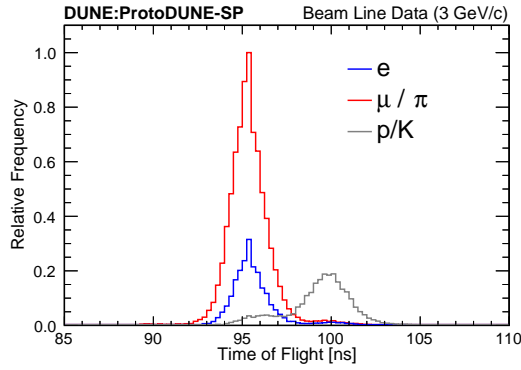
TABLE 5.1: A summary of beam line instrumentation logic used in the identification of particle types. Each cell reflects how a particular type of instrumentation is used at a given reference momentum. When time of flight is used, the values of the lower and upper cuts are given in nanoseconds. In the case of the high-pressure Cherenkov (“High-p Status”) and the low-pressure Cherenkov (“Low-p Status”), zero and one represent the absence and presence of a signal respectively. When a given piece of instrumentation is not involved in a logic decision at a given momenta, a dash is used.



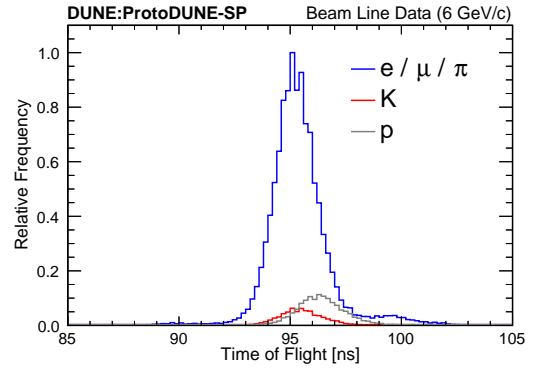
(a) Nominal beam momentum = 1 GeV/c. Vertical lines represent the time of flight cuts used for electrons (blue), and muons/pions (red).



(b) Nominal beam momentum = 2 GeV/c. Vertical lines represent the time of flight cuts used for electrons (blue), and muons/pions (red).



(c) Nominal beam momentum = 3 GeV/c.



(d) Nominal beam momentum = 6 GeV/c.

Figure 5.11: Time of flight distributions for different reference momenta, separated by particle using the PID techniques listed in table 5.1. The distributions are normalized such that the maximum height is equal to 1. Taken from Reference [30].

CHAPTER 6

EVENT SELECTION

1205

1206 This section describes the characterization of reconstructed data and MC events. Included
1207 is a set of data-MC comparisons detailing the cuts used in the selection. The data shown
1208 here is from Run 5387 of the initial ProtoDUNE-SP running period in the Fall of 2018.

1209 For data, an event is included in the set if it passes the following criteria:

1210 1. It is an event that was triggered by the beam line

1211 2. It follows the π/μ beam line selection.

1212 3. It has singular hits in each beam profile monitor. This is to eliminate ambiguity in the
1213 beam line momentum and tracking reconstruction.

1214 For MC, due to the lack of fully simulated beam line instrumentation, the only require-
1215 ment is that the simulated event was generated from a (primary) π^+ or μ^+ in the beam
1216 line simulation. Only μ^+ and π^+ are considered because at 1 GeV/ c (the beam momentum
1217 used for this analysis), the beam line PID can distinguish π^+ and μ^+ from e^+ and p , but
1218 not from each other. The criteria for the beam line PID can be seen in Table 5.1. Figure 5.11a
1219 shows that the protons are well separated by the TOF cut used to select μ^+/π^+ . The MC
1220 events have been normalized to the number of data events that pass the aforementioned data
1221 criteria.

1222 6.1 Truth Definitions

1223 The MC events which pass the above criteria are separated into the following categories
1224 based on truth information of the primary beam particle:

1225 1. Muons: The primary beam particle was a μ^+ .

- 1226 2. Upstream Interaction: The primary beam π^+ did not reach the TPC Fiducial Volume
1227 (FV).
- 1228 3. Past FV: The primary beam π^+ extended past the FV in the z -direction¹.
- 1229 4. Background Interaction: The primary beam π^+ interacted within the FV, and that in-
1230 teraction was a background (not Absorption or Charge Exchange) inelastic interaction.
1231 This includes any inelastic interaction between the primary π^+ and an Ar nucleus with
1232 an outgoing π^\pm above a momentum threshold of 150 MeV/ c . Note: this threshold is
1233 discussed further below.
- 1234 5. Absorption: The primary beam π^+ interacted within the FV in an Absorption inter-
1235 action. This is the first type of signal event and is defined as a π^+ which interacted
1236 with an Ar nucleus and resulted in no outgoing π^\pm (above threshold) or π^0 .
- 1237 6. Charge Exchange: The primary beam π^+ interacted within the FV in a Charge Ex-
1238 change interaction. This is the second type of signal event and is defined similarly to
1239 Absorption, but with any number of π^0 present.
- 1240 7. Other: The primary beam π^+ ended within the FV, and did not interact inelastically
1241 (i.e. it decayed in flight or came to a stop and then decayed at rest).

1242 The signal categories have been defined to occur within the FV (defined for primary par-
1243 ticles ending before $z = 222$ cm). This is due to the fact that the grounded electron diverters
1244 created electric field distortions (as described in Section 4.2.2) which caused reconstructed
1245 tracks to break in their vicinity. A data-driven simulation of the electric field distortions was
1246 implemented, which attempted to account for this effect in MC. The track-breaking effect
1247 can be seen in the Figure 6.1, which shows the reconstructed endpoint of beam tracks in the

¹Reminder: using a right-handed coordinate system, the z -direction is horizontal and follows the beam direction, the x -direction is horizontal and points away from the wires on the beam side TPCs, and the y -direction is vertical and points up.

1248 TPC in the z direction. The legend shows the truth categories described in the previous section. Note that the exact effect is not perfectly modeled by the simulation, and a systematic

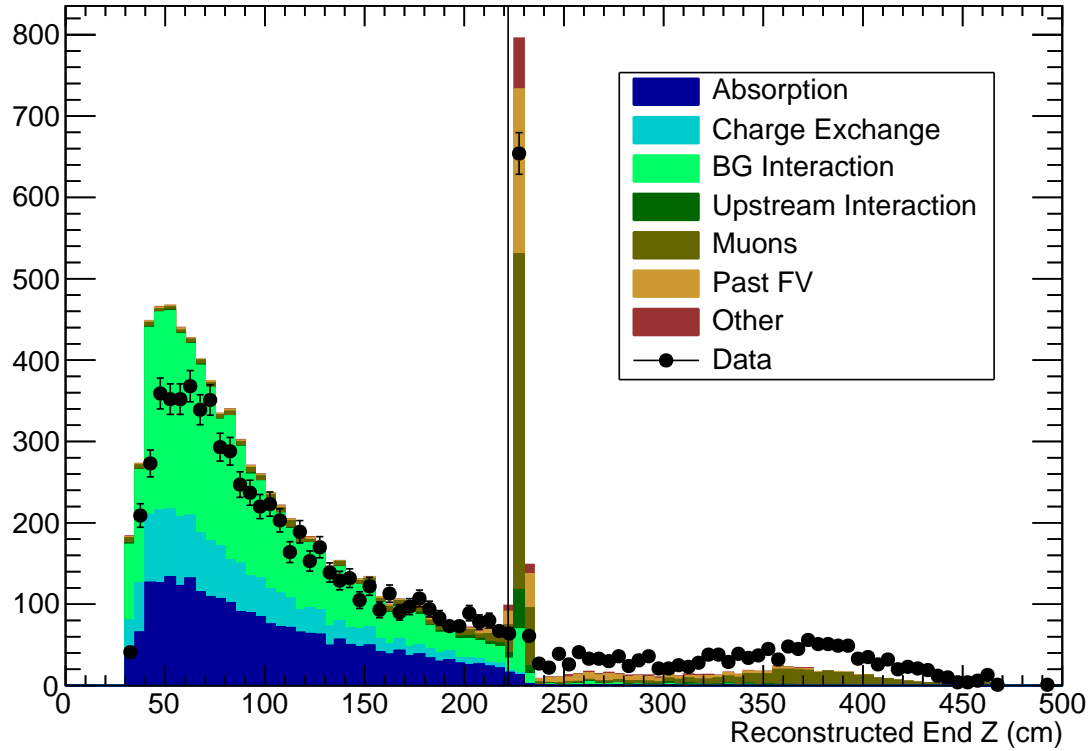


Figure 6.1: Reconstructed endpoint of beam tracks within the TPC. The vertical line represents the FV cut at 222 cm. The spike immediately after the FV cut is the track-breaking effect from the grounded electron diverters.

1249

1250 uncertainty on the strength of this effect in MC was implemented. This will be discussed
 1251 further in Section 8.3.

1252 Additionally, as mentioned above, the signal definitions are defined to have no charged pi-
 1253 ons above a momentum threshold of 150 MeV/ c . This is due to the inefficiency to reconstruct
 1254 charged pions exiting the primary interactions which are below this threshold.

1255 6.2 Event Selection

1256 The events that pass the previously stated criteria are then categorized according to
1257 the results of the TPC reconstruction. Ultimately, attempts are made to distinguish π^+
1258 from μ^+ and then to distinguish absorption and charge exchange interactions from other π^+
1259 interactions and stopping π^+ . Every event is accounted for and characterized into one of the
1260 following categories:

- 1261 1. The event contained no Pandora-reconstructed beam track *or* it did not leave enough
1262 hits on the collection plane wires. If, after the Pandora reconstruction described in
1263 Section 4.4 is performed, either no beam object was found or the beam object was
1264 reconstructed as a shower, the event is placed in this category. Also, events are placed
1265 here if there are not enough hits on the collection plane wires, as these are used in later
1266 cuts and in binning the events.
- 1267 2. The event contained a reconstructed beam track, but it was not considered consistent
1268 with coming from the beam. This is done in order to pick out events in which the pion
1269 interacted upstream of the TPC FV or if Pandora erroneously reconstructed a cosmic
1270 particle as the beam track. This is described in Section 6.3.
- 1271 3. The event was consistent with the incident track, but it extended past the FV cut in
1272 the z -dimension (222 cm).
- 1273 4. The event remained in the FV, but it was rejected by the combined absorption and
1274 charge exchange selection. The selection criteria for this and the following two cate-
1275 gories is presented in Section 6.4.
- 1276 5. The event passed the combined absorption/charge exchange selection, and was distin-
1277 guished as an absorption interaction.
- 1278 6. The event passed the combined absorption/charge exchange selection, and was distin-
1279 guished as a charge exchange interaction.

1280 6.3 Beam Cuts

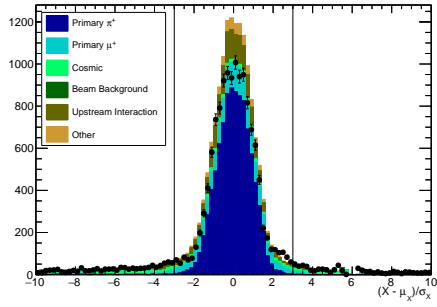
1281 Sometimes, the wrong particle is identified as the beam particle by the Pandora recon-
1282 struction. As mentioned above, these could come from a cosmic muon or particle resulting
1283 from an interaction before the active volume of the detector. Information about the starting
1284 position and direction of the reconstructed track identified as beam is used to separate these
1285 out and place them in their own category. For the cuts in position, the mean (μ) and width
1286 (σ) of the beam track distribution in x , y , and z is found (using SCE-corrected informa-
1287 tion). Any track that is at least 3σ away from the mean in any direction is categorized as
1288 inconsistent with the beam. Additionally, the direction of the track is taken from the vector
1289 connecting the SCE-corrected start and end points of the track. The means of the angles
1290 relative to the Cartesian planes ($\bar{\theta}_x, \bar{\theta}_y, \bar{\theta}_z$) are found. Then, for each beam track, the cosine
1291 of the angle between its direction and the mean direction (defined by the mean angles) is
1292 found. This is defined in Equation 6.1. Any track which has $\cos(\theta) < 0.95$ is considered
1293 inconsistent with the beam.

$$\cos(\theta) = \cos(\theta_x) * \cos(\bar{\theta}_x) + \cos(\theta_y) * \cos(\bar{\theta}_y) + \cos(\theta_z) * \cos(\bar{\theta}_z) \quad (6.1)$$

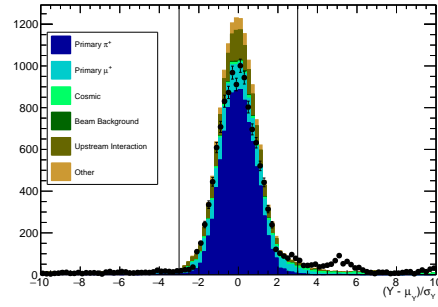
1294 Figure 6.2 shows the distributions of the position (relative to the μ and σ) of the beam in
1295 each direction, as well as two views of the $\cos(\theta)$ distribution. The 3σ cuts in position and
1296 the $\cos(\theta) < 0.95$ cut in direction are shown as the vertical black lines. As can be seen in
1297 these plots, the cosmic particles and upstream interactions tend to have extreme angles and
1298 positions.

1299 6.4 Absorption and Charge Exchange Selection

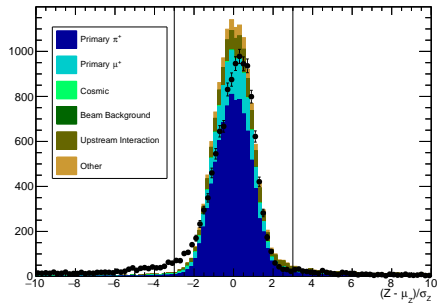
1300 By analyzing reconstructed particles that have been associated to the TPC beam track as
1301 daughter particles, the tracks ending within the FV are separated between two categories: 1)
1302 absorption or charge exchange 2) other events. Because both absorption and charge exchange
1303 events contain no charged pion (above threshold) in the final state, the selection strategy is



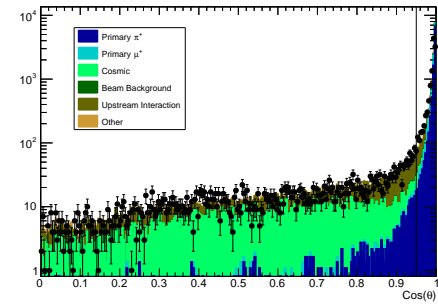
(a) SCE-corrected start x position reconstructed TPC track relative to mean and width of all beam tracks.



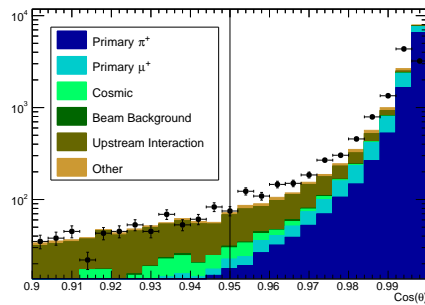
(b) SCE-corrected start y position reconstructed TPC track relative to mean and width of all beam tracks. Note that a bump in the distribution exists towards the right of the plot. This was recently identified as caused by a known detector effect present in data. Further investigation and possible treatments in MC simulation is ongoing.



(c) SCE-corrected start z position reconstructed TPC track relative to mean and width of all beam tracks.



(d) Cosine of the angle between reconstructed beam track and the mean track.



(e) Cosine of the angle between reconstructed beam track and the mean track. Zoomed-in to highlight cut region.

Figure 6.2: Distributions used to determine consistency with the beam line. The vertical black lines represent the cut values used.

1304 to identify events with a charged pion daughter.

1305 The daughter particles are separated between track-like (ideally from μ , π , p , etc.) and
1306 shower-like (ideally from e , γ) objects using the results of the CNN described in Section
1307 4.4.1. For daughter particles of beam tracks, the CNN-based track/shower discrimination
1308 performed better than Pandora’s native track/shower discrimination, and so was used for this
1309 analysis. Each hit in an event receives a set of scores produced by the CNN that encodes the
1310 degree to which it appears to be produced by a track-like particle or a shower-like particle. For
1311 each associated daughter, the scores from all of its hits are averaged to produce aggregated
1312 scores for the reconstructed particle. The Pandora reconstruction software was configured to
1313 reconstruct both a track-like and shower-like object for each reconstructed particle cluster
1314 in an event, so that analyzers could use alternate track/shower discrimination (such as the
1315 CNN method described here) and access the information accordingly.

1316 At time of writing, only the calibration for collection plane hits was at a suitable state,
1317 and so only these hits were used to calculate the aggregated scores. A cut on the track-
1318 like score of the daughter particle at 0.3 (shown in Figure 6.3) was used to separate the
1319 daughters into shower-like and track-like. Here, the track score of every reconstructed particle
1320 associated as a daughter particle to the primary reconstructed TPC particle. The MC has
1321 been categorized by the true particle that created the reconstructed particle object. The
1322 fields “Daughter+” and “Daughter++” represent particles that are products of reinteractions
1323 of final state particles and so on. The field “Self” refers to segments of the true primary
1324 particle that were associated as a daughter (i.e. the track ended early). Finally, the field
1325 “ γ ” represents photons emitted by the nucleus following a primary interaction (i.e. from
1326 nuclear de-excitation), while “ $\pi^0\gamma$ ” represents photons truly originating from the decay of a
1327 π^0 created within a primary interaction.

1328 If the daughter is considered track-like, an attempt is made to identify charged pions
1329 by identifying particles that appear to be a Minimum Ionizing Particle (MIP). This MIP-
1330 like determination is done first by looking at the energy deposited per unit length by the

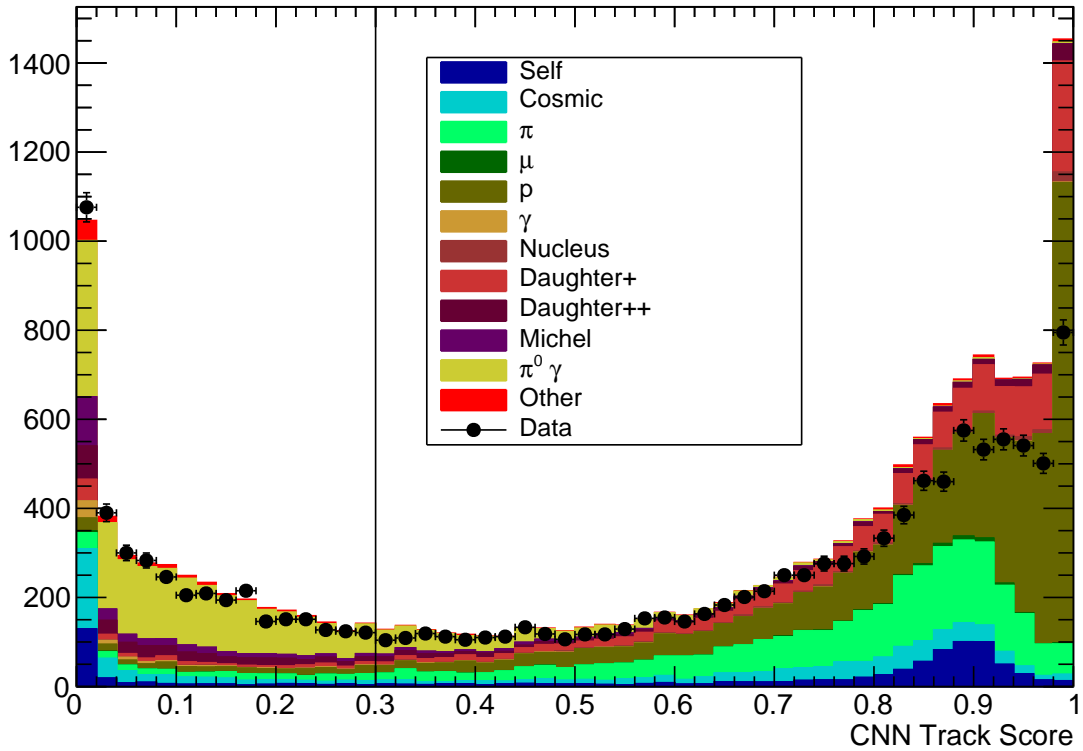
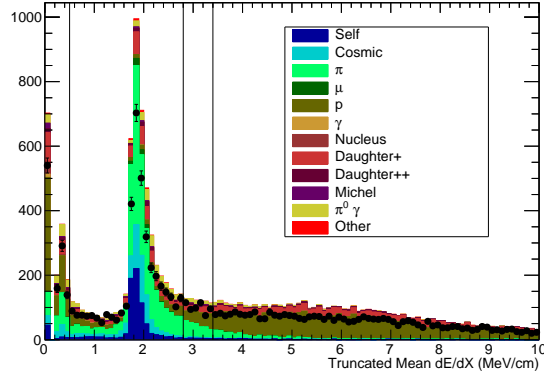
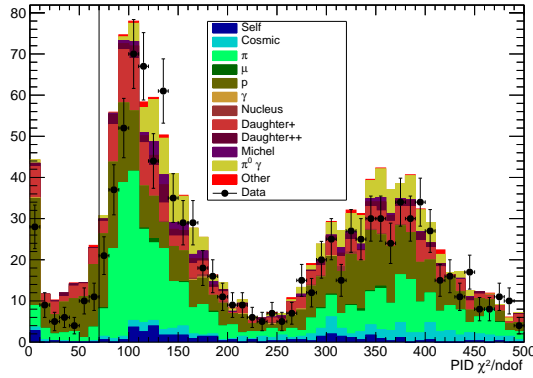


Figure 6.3: CNN Track scores of all reconstructed particles associated as daughters to the primary beam.

1331 reconstructed track. For this, the truncated-mean dE/dx (defined to be the total energy
 1332 deposited by a reconstructed hit divided by the track pitch of that hit) is used in order to
 1333 exclude the large energy deposits from stopping particles. In its calculation, the lowest 16%
 1334 and highest 16% of hits in a track are ignored. The distribution of the truncated mean
 1335 dE/dx for all daughter tracks is shown in Figure 6.4a. Particles are immediately considered
 1336 MIP-like if they fall between 0.5 and 2.8 MeV/cm, and are considered not MIP-like if they
 1337 are above 3.4 MeV/cm. For other particles (those that fall below .5 MeV/cm or between
 1338 2.8 and 3.4 MeV/cm), another step is done in the selection. This step consists of comparing
 1339 the dE/dx of each hit in the track to the expectation value for protons and producing a χ^2
 1340 value. Ideally, protons should have a low χ^2 and pions should have a high χ^2 . This is shown
 1341 in Figure 6.4b. These particles are considered MIP-like if they have a χ^2 above 70.



(a) Truncated mean dE/dX of all reconstructed tracks associated as daughters to the primary beam.

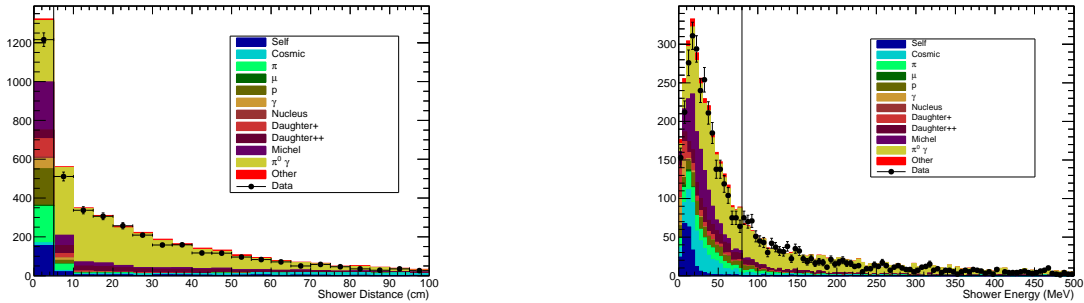


(b) PID χ^2 value of all reconstructed tracks associated as daughters to the primary beam.

Figure 6.4: Distributions used for combined absorption and charge and exchange selection. Vertical black lines represent the cuts used. Note that the events were separated into multiple regions of truncated mean dE/dX , as indicated by the multiple black lines on the left plot.

1342 If any of the daughter particles appears MIP-like, it is considered to be a charged pion
 1343 originating from the primary interaction, and the event is rejected from the absorption and
 1344 charge exchange selection.

1345 Following the combined absorption and charge exchange selection, these interactions are
 1346 separated by attempting to identify showers originating from the decay of π^0 daughters. A
 1347 daughter shower is considered as coming from a π^0 decay if it is at least 5 cm away from
 1348 the end of the primary track and has at least 80 MeV of energy. These cuts are chosen
 1349 to exclude any activity around the interaction vertex which originated from lower-energy
 1350 pions and protons or nuclear de-excitation photons from both the primary and downstream
 1351 interactions. This can be seen in Figure 6.5.



(a) Distance between the end of the reconstructed beam track and start of showers associated as daughters to the primary track.

(b) Total deposited energy of all showers associated as daughters to the primary track.

Figure 6.5: Distributions used to separate absorption from charge exchange. The black vertical lines represent the cuts used.

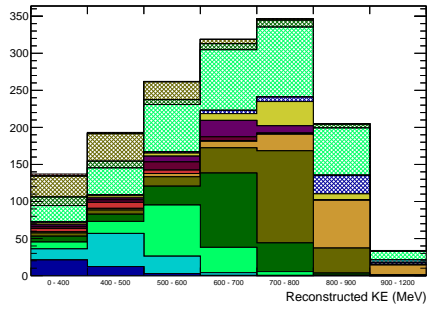
Channel	Efficiency	Purity
Absorption	0.53	0.52
Charge Exchange	0.23	0.80

TABLE 6.1: Efficiency and purity of the signal categories.

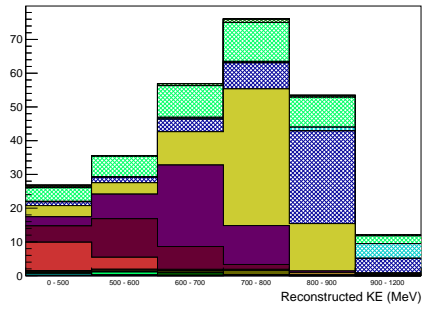
1352 6.5 Binning

1353 The bin variables used for the event selection categories described in the previous sec-
1354 tion are as follows. Events that fall in the first two categories (no beam track and events
1355 that fail the beam cuts) are each placed in single, unitless bins. Events that end past the
1356 FV cut are binned according to their SCE-uncorrected ending position in z . Finally, the
1357 three “interaction” categories (absorption, charge exchange, and other) are binned according
1358 to their ending kinetic energy. This is determined by first calculating their reconstructed
1359 kinetic energy using the reconstructed beam line momentum, assuming they are pions, and
1360 then subtracting the energy of each collection plane hit up to but not including the last.
1361 Occasionally, large hits from large amounts of vertex activity or from crossing cosmic tracks
1362 saturate the cold electronics, resulting in seemingly enormous reconstructed energy deposits
1363 on the order of a few hundred to a thousand MeV. Thus, any hit above 80 MeV is ignored in
1364 the calculation. This value was chosen such that the saturated electronics hits are skipped,
1365 but truly large energy deposits like from overlapping hits are kept.

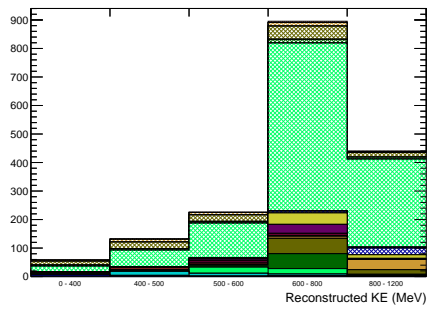
1366 This binning is shown in Figure 6.6, where the reconstructed distributions from the
1367 nominal MC are shown. Additionally, these distributions are broken down by their true
1368 category (i.e. a signal interaction in a given energy bin or a muon). These are displayed as
1369 the stacked histograms, where the different colors represent the specific true category. The
1370 bin edges for the interaction distributions were chosen based on the smearing between true
1371 and reconstructed kinetic energy shown as the spread in the different colored portions of the
1372 stacks. The purity and efficiency for the absorption and charge exchange selections is also
1373 shown in Table 6.1.



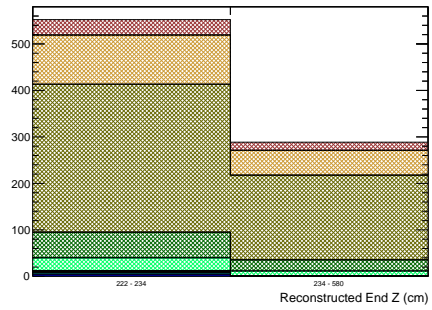
(a) Selected Absorption



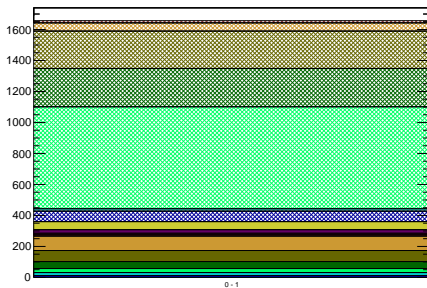
(b) Selected Charge Exchange



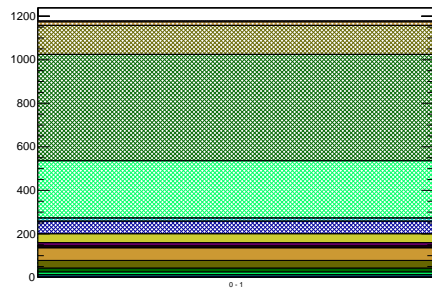
(c) Selected other interactions.



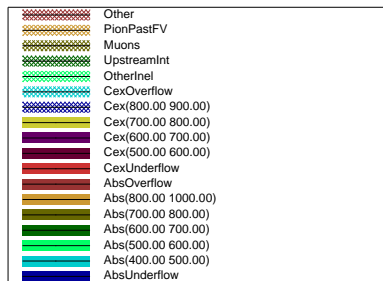
(d) Selected as ending past FV



(e) Failed beam cuts



(f) No reconstructed beam track



(g)

Figure 6.6: Reconstructed distributions of events from the nominal MC. The distributions are broken down by true categories, shown in 6.6g.

1374 6.6 Selected MC Event Displays

1375 This section provides some examples of successes and failures in the event selection within
1376 the MC sample used in the analysis. Shown in the following figures are reconstructed event
1377 displays in the view of the collection plane near the beam entrance. The x-axis is the wire
1378 number which is equivalent to the position in z. The y-axis is the time (or tick) at which
1379 the charge reached the wire plane. This is equivalent to the horizontal position away from
1380 the wire plane.

1381 The first example, shown in Figure 6.7, is a true absorption interaction correctly selected
1382 as absorption. The pion is shown as the yellow track entering from the left, and it interacts
1383 with a nucleus. The interaction produces two protons. These are correctly identified as
1384 proton-like tracks by the event selection and are shown as the light blue and pink tracks in
1385 the display.

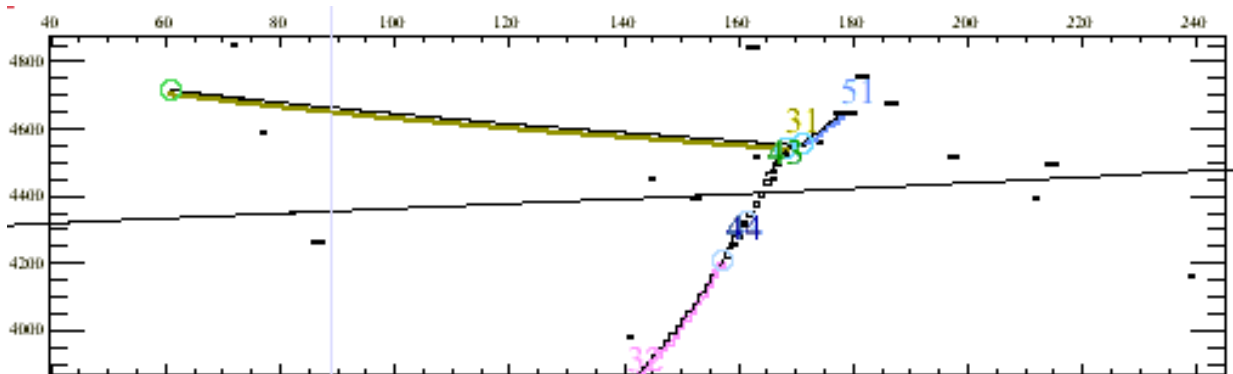


Figure 6.7: MC absorption event correctly identified as absorption.

1386 Next, Figure 6.8 shows a true charge exchange event incorrectly identified as an absorp-
1387 tion event. Again, the beam enters from the left shown as the tan track, and interacts with
1388 a nucleus. A very energetic proton exits the interaction and travels toward the lower right
1389 of the display shown as the green track. This proton is correctly identified as a proton.
1390 However, a π^0 also exits the interaction. Near the vertex, one of the γ s produced by the
1391 decay of this is identified as a small shower (represented by the black rectangles near the
1392 vertex). Its reconstructed energy is too low to be identified as resulting from a π^0 . The other

1393 γ is not identified. The pink track extending from the top to bottom of the plot is a cosmic
 1394 muon.

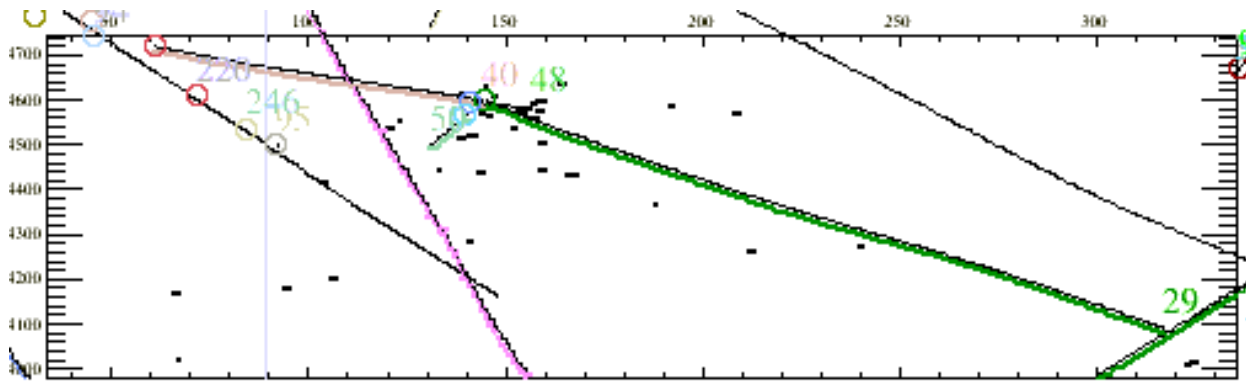


Figure 6.8: MC charge exchange event incorrectly identified as absorption.

1394

1395 The third example in Figure 6.9 is a background inelastic event (one a charged pion in
 1396 the final state) selected as absorption. The beam pion enters from the left (shown as the
 1397 red track), and strikes a nucleus. Both a π^+ and π^0 exit the interaction. The π^0 promptly
 1398 decays, and the resulting showers are not associated to this event as daughters. The π^+ is
 1399 reconstructed as the tan track exiting the interaction, but it does not appear to be a pion
 1400 when its calorimetry information is checked in the event selection procedure. The light blue
 1401 track extending from the top right to the bottom left is a cosmic muon.

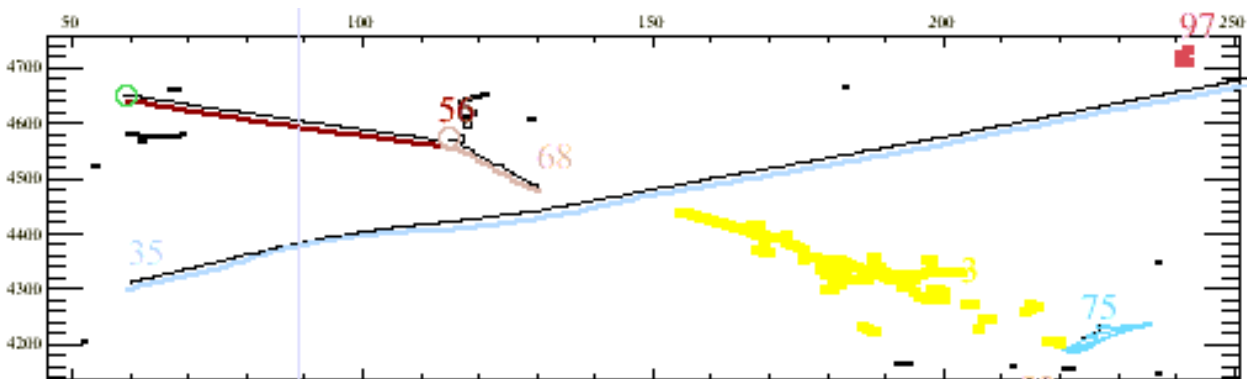


Figure 6.9: MC background inelastic event incorrectly identified as absorption.

1402 The next example in Figure 6.10 shows the pion as a tan track entering from the top left
 1403 before interacting with a nucleus. A resulting proton is reconstructed as the light blue track
 1404 heading toward the bottom of the figure. A π^0 exits the interaction and promptly decays.

1405 The resulting photons are reconstructed as the red and yellow showers, and identified as
1406 such.

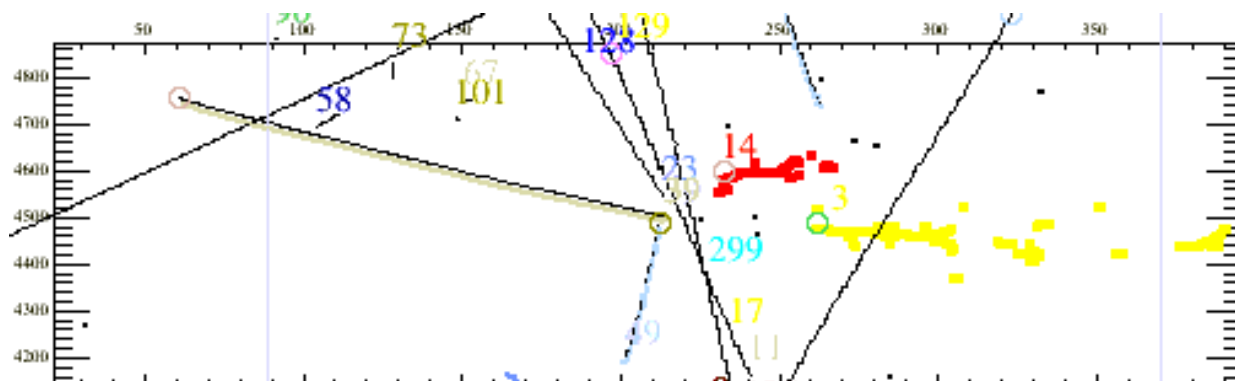


Figure 6.10: MC charge exchange event correctly identified as charge exchange.

1407 Figure 6.11 shows an absorption event misidentified as charge exchange. The beam pion is
1408 reconstructed as the pink track and interacts with a nucleus. A neutron exits the interaction
1409 before itself interacting and resulting in a proton track (the tan track toward the right).
1410 A proton also exits the interaction, but was reconstructed as a shower (represented by the
1411 black boxes at the end of the track). This proton appears as a π^0 shower and so the event
1412 is selected as charge exchange.

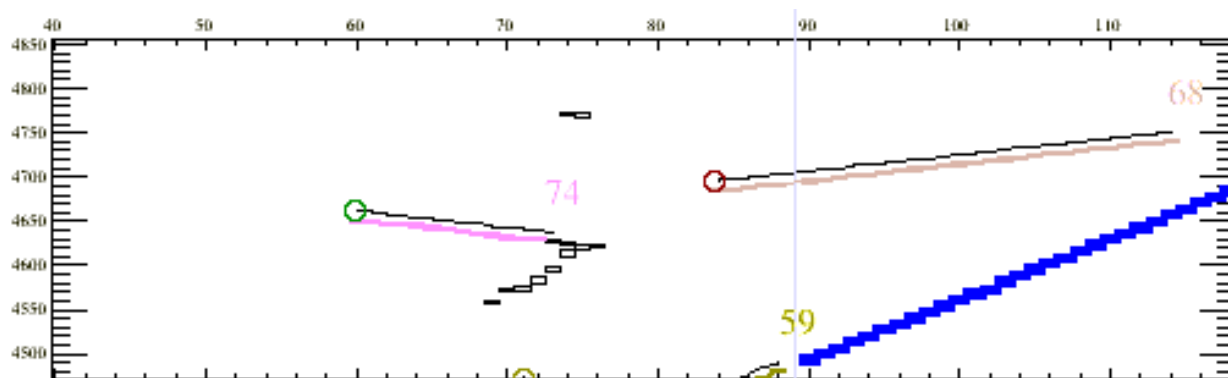


Figure 6.11: MC absorption event incorrectly identified as charge exchange.

1413 Figure 6.12 shows a background inelastic misidentified as charge exchange. The pion,
1414 reconstructed as the tan track, enters from the left and ends in an inelastic interaction. A
1415 high energy π^+ exits the interaction and reinteracts nearby the primary interaction, resulting

1416 in a charge exchange event. The π^0 from the secondary interaction decays, and a shower is
 1417 reconstructed and associated to the primary interaction.

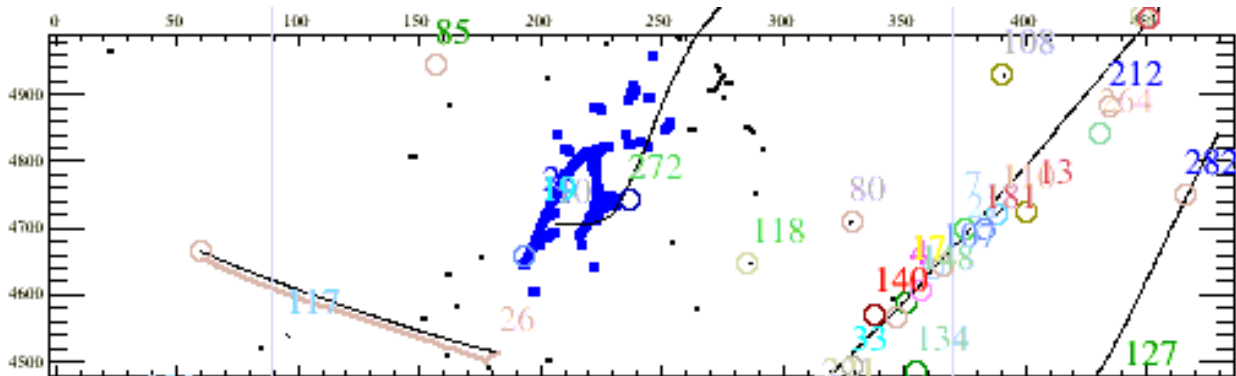


Figure 6.12: MC background inelastic event incorrectly identified as charge exchange.

1418 Finally, Figure 6.13 shows a muon that is misidentified as a background inelastic inter-
 1419 action. The muon, reconstructed as the yellow track, enters from the top left of the plot and
 1420 its track is ended prematurely. The remainder of the muon is reconstructed as a MIP-like
 1421 track (the red track) and associated as a daughter to the primary track. The blue and yellow
 1422 tracks toward the left of the figure extending from top to bottom are cosmic muons.

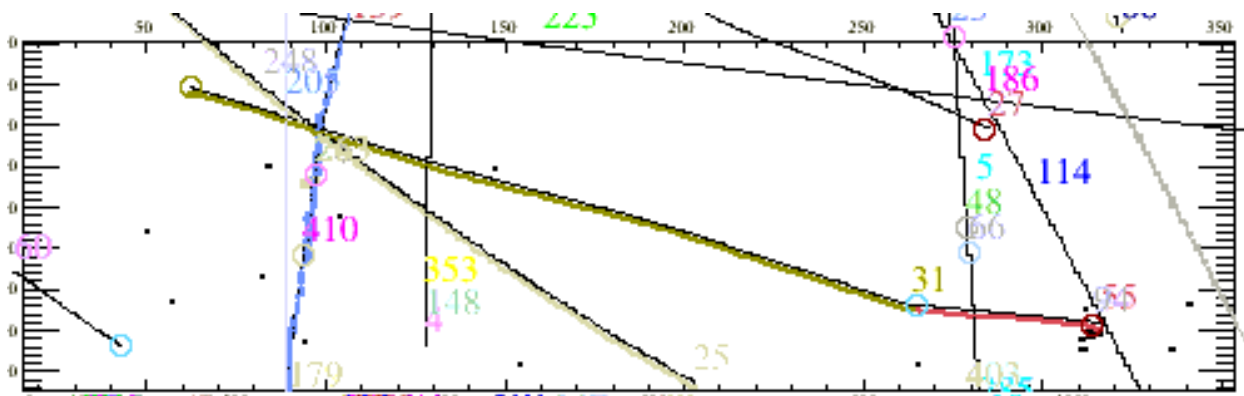


Figure 6.13: MC muon incorrectly identified as a background inelastic interaction.

1423 6.7 Selected Data Event Displays

1424 This section provides example events in 5 of the 6 selection categories (all but the category
 1425 “no-track” category) used in the fit to data. The dataset containing these events, Run 5809,
 1426 is different from the one used to display the event selection cuts.

1427 The first example in Figure 6.14 is a selected absorption event. The pion candidate enters
 1428 from the left, and appears to interact with a nucleus. The reconstruction does not associate
 1429 any tracks as daughters to this primary particle. Despite this, there appear to be a pair of
 1430 heavily-ionizing protons exiting the interaction. A cosmic muon crosses the primary track
 in a nearly-vertical trajectory, and a pair of cosmic muons appear toward the right.

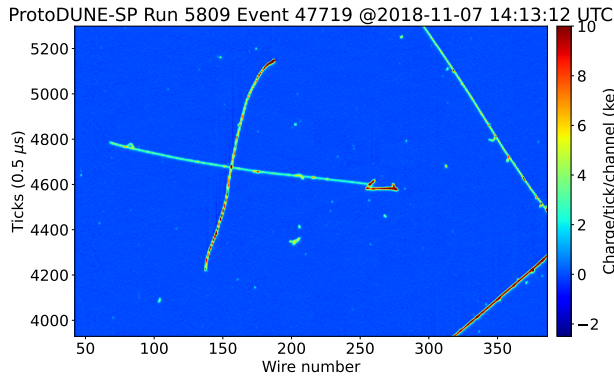


Figure 6.14: Selected absorption event.

1431 The second example shown in Figure 6.15 is a selected charge exchange event. The pion
 1432 candidate enters from the left, and appears to interact with a nucleus. Clearly seen after the
 1433 interaction is an apparent shower structure resulting from the decay of a π^0 .

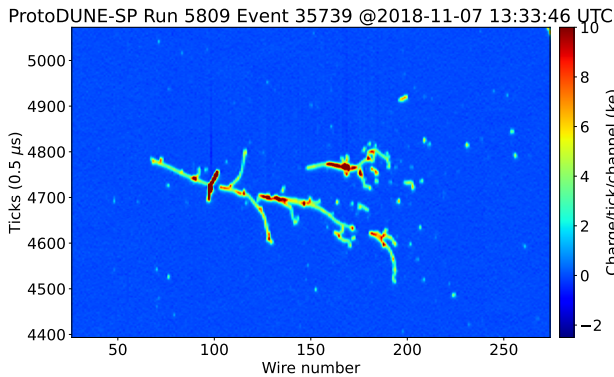


Figure 6.15: Selected charge exchange event.

1434 The third example in Figure 6.16 is a selected background inelastic interaction. The pion
 1435 candidate enters from the left, and results in an interaction with multiple particles exiting.
 1436 A daughter pion candidate travels from the interaction toward the top right of the plot.
 1437

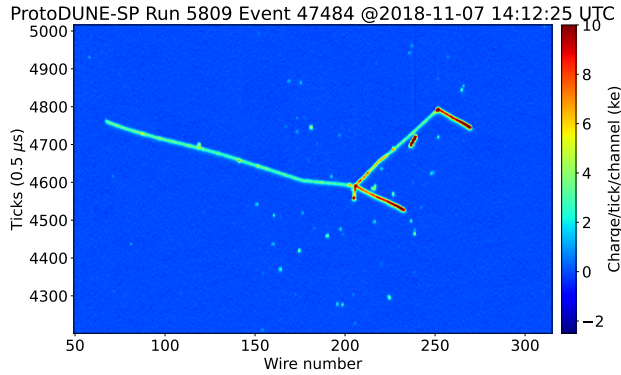


Figure 6.16: Selected background inelastic interaction event.

1438 The fourth example in Figure 6.17 is a π^+/μ^+ candidate extending past the fiducial
 1439 volume. The primary particle appears to come to a stop near wire number 700. A break in
 1440 the particle's ionization track is seen near wire number 500. This is the dead region caused
 1441 by the grounded electron diverters. Additionally, a cosmic muon is seen crossing the primary
 track.

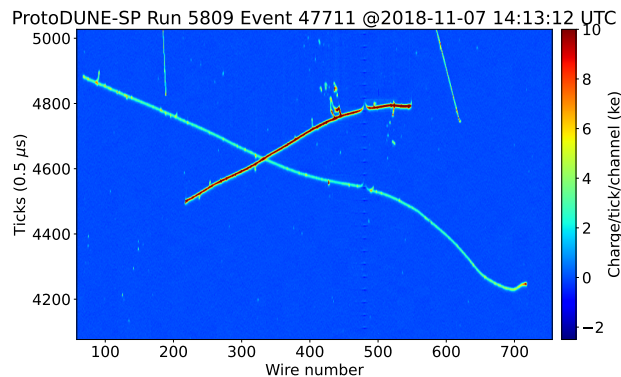


Figure 6.17: Event selected as extending past the fiducial volume.

1442

1443 Finally, in Figure 6.18 is another π^+/μ^+ candidate that extends past the fiducial volume.
 1444 This time, however, the reconstruction (not shown) ends near the grounded electron diverters.
 1445 The remainder of the primary particle's ionization to the right of the grounded electron
 1446 diverters is reconstructed as a separate track, and is associated as a daughter to the primary
 1447 track.

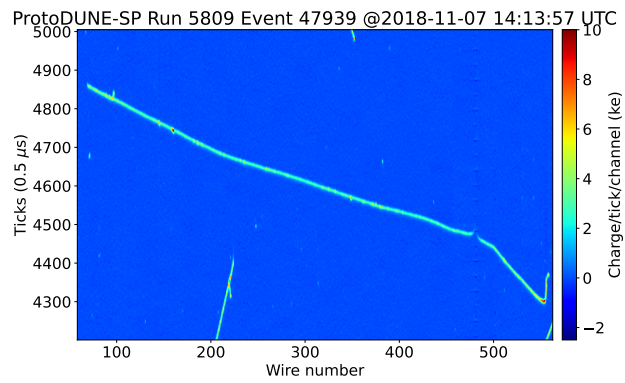


Figure 6.18: Event selected as extending past the fiducial volume, and specifically ending near the electron diverter region.

CHAPTER 7

1448

CROSS SECTION MEASUREMENT TECHNIQUE

1449 This analysis measures π^+ - Ar absorption and charge exchange cross sections using beam-
1450 triggered events in ProtoDUNE-SP. The measurement employs a fit which extracts the
1451 number of signal (absorption and charge exchange) interactions as well as the number of
1452 background events (incident muons, non-signal interactions, stopping pions) from this data.
1453 Truth-level information (information representing the exact results of the simulation, rather
1454 than reconstructed information taken from a simulated detector response) is then used to
1455 extract the cross section according to a technique derived from the Liquid Argon in A Test
1456 Beam experiment (LArIAT) [50]. That technique, known as the "Thin Slice Method" (de-
1457 scribed in Section 7.2) was used to measure hadron cross sections using a LArTPC wherein
1458 the detection medium (LAr) also serves as the target. This method is distinct from mea-
1459 surements using thin targets. This chapter first describes these thin target cross section
1460 measurements, as well as the Thin Slice Method. It then specifies how the Thin Slice
1461 Method is used on truth information to extract the cross section from simulation. It then
1462 describes the statistical fit used to interpret the data.

1463 7.1 Thin Target Cross Section Experiment

1464 Historically, hadron scattering experiments have been performed by firing a beam of
1465 particles onto a thin piece of material as a target. By counting the number of interactions,
1466 the cross section for an interaction can be measured as a function of the incident energy
1467 (since the target is thin, a negligible amount of energy is lost before an interaction, and the
1468 cross section is measured at the incident beam energy). A simple cartoon of the experimental
1469 setup can be seen in Figure 7.1. Here, a beam of pions of width A and flux Φ impinges on a
1470 target of thickness t . After passing through the target, N_{Inter} pions have interacted, while
1471 N_{Surv} have passed through without interacting. The cross section can be extracted from

1472 Equation 7.1.

$$\frac{N_{Inter}}{\Phi A} = \frac{N_{Inter}}{N_{Inc}} = 1 - e^{-nt\sigma} \quad (7.1)$$

1473 Here, σ is the cross section for the relevant interaction, n is the number density of atoms
 1474 in the target material, and N_{Inc} is the number of incident pions as given by $\Phi \times A$. This
 1475 can be slightly simplified by expanding the exponential term around t as such:

$$\frac{N_{Inter}}{\Phi A} = \frac{N_{Inter}}{N_{Inc}} \approx 1 - (1 - nt\sigma + \mathcal{O}(t^2)) = nt\sigma \quad (7.2)$$

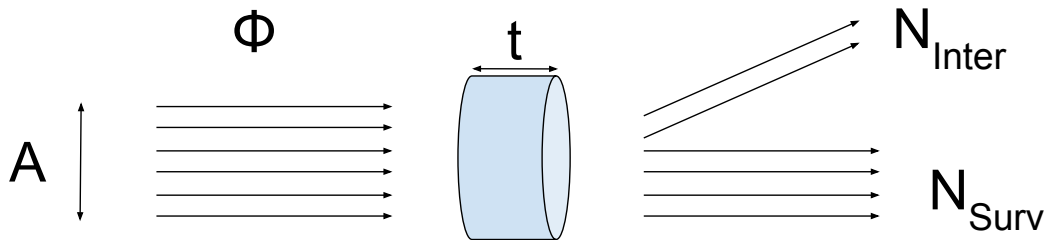


Figure 7.1: Cartoon of a thin target scattering experiment.

1476 7.2 The Thin Slice Method

1477 By virtue of being a LArTPC, ProtoDUNE-SP is not thin, and thus cannot be used for
 1478 the simple thin target experiment as described above. However, LArIAT [50] used a method
 1479 they called the Thin Slice Method to mock-up a series of multiple thin target experiments in
 1480 an extended volume of LAr in order to measure hadronic cross sections in an extended LAr
 1481 volume. The segmentation created by the collection plane wires allows analyzers to treat an
 1482 extended volume of LAr as if it were multiple thin targets stacked in front of one another.

1483 This can be seen in Figure 7.2, where a cartoon of a pion track in a LArTPC is shown. The
 1484 vertical dashed lines represent the collection wires of the TPC, and the red dot represents the
 1485 point at which the pion interacts. One can treat every slice the pion passes through (up to
 1486 and including the slice which contains the interaction) as a separate thin target experiment.
 1487 In each of these, the pion enters the slice and either interacts, or decays. From this, one can
 1488 count the number of incident pions (N_{Inc} as described above) by counting the number of
 1489 times a pion enters a slice (it passes by a new wire) and the number of interactions (N_{Inter})
 1490 to extract the cross section.

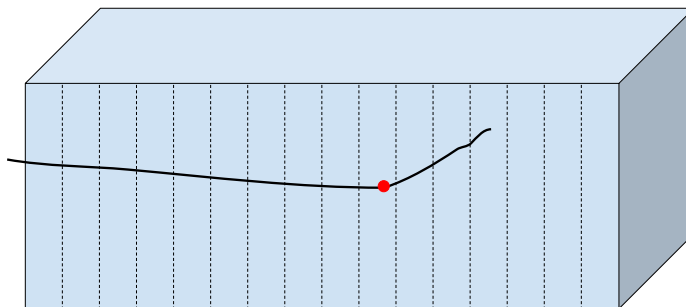


Figure 7.2: Cartoon of the thin slice method applied to a pion track within a LArTPC. The red point represents a hadronic interaction.

1491 If the energy of the pion is known as it enters each slice, then energy-dependence is added
 1492 to Equation 7.2, as reflected in Equation 7.3. Here, the thickness t is the width of the wire
 1493 spacings.

$$\frac{N_{Inter}(E)}{N_{Inc}(E)} = nt\sigma(E) \tag{7.3}$$

1494 Mechanically, this calculation is achieved by using two histograms thus called “Incident”
 1495 and “Interacting” which respectively represent the denominator and numerator of Equation

1496 7.3. As the pion enters into a new slice, the Incident histogram is filled at the corresponding
 1497 energy. This is done for the entire pion track up to the end, meaning a track can contribute
 1498 multiple entries in the histogram. For example, in Figure 7.2, the pion track will contribute
 1499 an entry for every section of Ar up to and including the interaction point (represented by
 1500 the red dot). If the pion undergoes an interaction of interest, the Interacting histogram is
 1501 filled according to the energy of the pion as it entered the final slice (this will be the same
 1502 energy for the final entry into the Incident histogram). A demonstration of this is shown in
 1503 Figure 7.3.

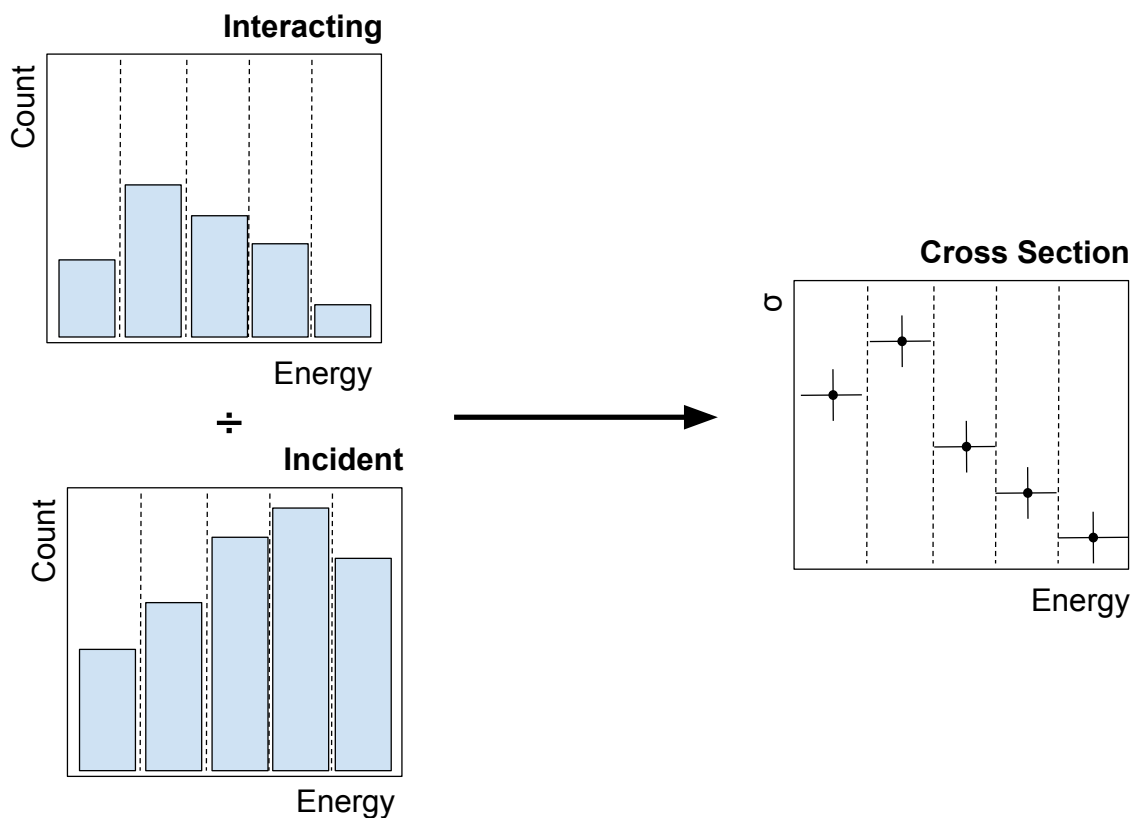


Figure 7.3: Demonstration of the cross section calculation using Equation 7.3.

1504 7.3 Thin Slice Method on Truth Information

1505 The previous section described how the Thin Slice Method could be used on reconstructed
1506 information to determine hadronic cross sections. The measurement presented in this thesis
1507 is slightly different, but is generally based on this method. Rather than using reconstructed
1508 information to determine the Incident histogram, it is taken directly from truth information
1509 from ProtoDUNE-SP Monte Carlo simulation. This simulation will be modified by perform-
1510 ing a fit to collected data. This fit, known as a “template fit” which is described later in
1511 Section 7.4, will vary the number of signal and background interactions (binned in true end-
1512 ing kinetic energy) within the MC. This will of course change any true Interacting histogram
1513 created from this information. It will, in turn, also change the true Incident histogram, as
1514 the number of slices (equivalently, the distance traveled by the pion) depend on the pion’s
1515 starting and ending energy. In this way, the varied MC which best describes the data can be
1516 used to extract a varied cross section. This section describes the procedure used to extract
1517 the cross section from truth information.

1518 The ProtoDUNE-SP MC simulation contains a set of π^+ and μ^+ created by the beam
1519 impinging on the detector. Pions that interact before the start of the LAr are ignored and
1520 do not contribute to the Incident distribution. For all other pions (those that enter into
1521 the TPC), their energy at the initial TPC point (E_0) is used as an entry in the Incident
1522 distribution. Using a uniform spacing¹, the energy deposited by the pion as it was simulated
1523 by Geant4 is separated into “slices”. The energy at each slice boundary crossed by the pion
1524 is calculated by summing the energy deposited in the previous slice and subtracting that
1525 from the previous incident energy. Thus, the energy as the pion crosses slice boundary i is
1526 equal to $E_{i-1} - \delta E_{i-1,i}$ where $\delta E_{i-1,i}$ is the energy deposited between slice boundaries $i - 1$
1527 and i . This is demonstrated in Figure 7.4, where the labels E_i represent the energy of the
1528 pion as it crosses each slice boundary. All of the energies after E_0 are then given an entry

¹Note, the width of the spacing to extract the cross section from truth info is arbitrary. For this analysis, the wire spacing (.47974cm) was used.

1529 in the Incident distribution as well. This occurs for every pion that reaches the TPC, and
 1530 along each pion up to some fiducial volume edge. Then, for each pion ending in a signal
 1531 interaction within the fiducial volume, the energy of the pion at its interaction point is used
 1532 as an entry in the Interacting histogram. The resulting Interacting and Incident histograms
 are used as in Equation 7.3 and Figure 7.3 to compute the cross section.

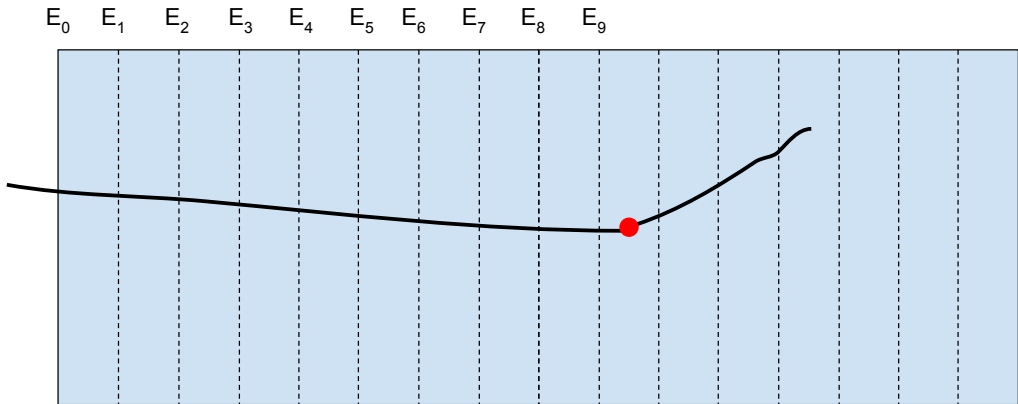


Figure 7.4: Cartoon diagram showing a pion track split up into multiple slices and the energy denoted at each slice boundary.

1533

1534 It is instructive to consider this measurement technique under a varied cross section
 1535 model. If the cross section is higher over the momentum range of the simulated pions, more
 1536 interactions will occur (N_{Inter} will be higher). The pions will (on average) travel through
 1537 less slices before they interact, and thus contribute less events to the Incident histogram. For
 1538 an overall lower cross section, the inverse is true: less interactions occur, and the pions travel
 1539 further on average (creating more entries in the Incident histogram). This line of thought can
 1540 be extended to more complicated variations in shape as well. The number of interactions at a
 1541 given energy will change, and so too will the entries in the Incident distribution. This serves
 1542 as the guiding principle used in this measurement: if one is able to measure the number
 1543 of interacting pions at a given energy (and equally importantly, the number of pions that

do not interact), the Thin Slice Method can extract cross sections using truth information from a varied Monte Carlo simulation that best describes the data. The following section describes the fit strategy used to interpret the data in terms of a varied Monte Carlo sample in order to extract the absorption and charge exchange cross sections in this manner.

7.4 Fit Strategy

This analysis uses a binned maximum likelihood fit to 1 GeV/c momentum ProtoDUNE-SP beam line events to estimate the number of signal and background interactions in the data set. The fit results in a set of varied MC which best matches this data and from which the signal cross sections are extracted. A set of signal parameters ($\vec{\theta}$) and nuisance (also called systematic) parameters (\vec{p}) controlled by the fit vary simulated π^+ and μ^+ events. The fit attempts to find the set of parameters that best describe the data by maximizing the likelihood $L(\vec{\theta}, \vec{p}; \vec{n})$ to observe a set of events \vec{n} given the model parameters $\vec{\theta}$ and \vec{p} . Additionally, we include constraints to the nuisance parameters represented by predictions of their central values \vec{q} and the uncertainties on these predictions represented by a covariance matrix V_{Cov} . As such, the likelihood L is made of two components: a statistical term and a systematic term:

$$L(\vec{\theta}, \vec{p}; \vec{n}) = L_{\text{Stat}}(\vec{\theta}, \vec{p}; \vec{n}) L_{\text{Syst}}(\vec{p}; \vec{q}, V_{\text{Cov}}) \quad (7.4)$$

For compatibility with the fitting routines (discussed later) in finding the best fit parameters and their uncertainties, the minimum of twice the negative log-likelihood ($-2 \ln L$) is found instead of the maximum likelihood². Additionally, minimizing this value is equivalent to minimizing twice the negative of the natural logarithm of the likelihood ratio λ [29]. The likelihood ratio is defined as

$$\lambda = L(\vec{\theta}, \vec{p}; \vec{n}) / L(\vec{\theta}_T, \vec{p}_T; \vec{n}) \quad (7.5)$$

²The fitting routines implemented in ROOT work by minimizing rather than maximizing some value.

1565 where $\vec{\theta}_T, \vec{p}_T$ represents the true, unknown underlying model. Plugging Equation 7.4 into
 1566 this results in the following.

$$\lambda = L_{\text{Stat}}(\vec{\theta}, \vec{p}; \vec{n}) L_{\text{Syst}}(\vec{p}; \vec{q}, V_{\text{Cov}}) / L_{\text{Stat}}(\vec{\theta}_T, \vec{p}_T; \vec{n}) \quad (7.6)$$

1567 Where there is no true value of L_{Syst} shown in the denominator, as it is trivially equal to
 1568 one. $-2 \ln \lambda$ is thus defined as

$$-2 \ln \lambda = -2 \ln (L_{\text{Stat}}(\vec{\theta}, \vec{p}; \vec{n}) / L_{\text{Stat}}(\vec{\theta}_T; \vec{n})) - 2 \ln L_{\text{Syst}}(\vec{p}; \vec{q}, V_{\text{Cov}}). \quad (7.7)$$

1569 In this fit, we are seeking to categorize a fixed number of events (a set of beam line-
 1570 triggered events) based on the results of ProtoDUNE-SP reconstruction described in Sec-
 1571 tion 4.4. As such, the likelihood L_{Stat} is the multinomial likelihood as defined in Equation
 1572 7.8.

$$L_{\text{Stat}}(\vec{\theta}, \vec{p}; \vec{n}) = N! N^N \prod_j y_j(\vec{\theta}, \vec{p})^{n_j} / n_j! \quad (7.8)$$

1573 Here, $y_j(\vec{\theta}, \vec{p})$ and n_j are the number of predicted and measured events in reconstructed bin
 1574 j , and $N = \sum_j n_j = \sum_j y_j(\vec{\theta}, \vec{p})$ is the total number of beam line events. As stated before,
 1575 $L_{\text{Stat}}(\vec{\theta}_T, \vec{p}_T; \vec{n})$ depends on some true underlying model denoted by $\vec{\theta}_T, \vec{p}_T$. This model is
 1576 unknown, but $L_{\text{Stat}}(\vec{\theta}_T, \vec{p}_T; \vec{n})$ is estimated using the measured events as shown in Equation
 1577 7.9.

$$L_{\text{Stat}}(\vec{\theta}_T, \vec{p}_T; \vec{n}) = N! N^N \prod_j n_j^{n_j} / n_j! \quad (7.9)$$

1578 From this, the statistical portion of $-2 \ln \lambda$ is defined as follows.

$$-2 \ln \lambda_{\text{Stat}} = 2 \sum_j n_j \ln \frac{n_j}{y_j} \quad (7.10)$$

1579 The systematic term $-2 \ln \lambda_{\text{Syst}}$ is a constraint term that assumes the systematic param-
 1580 eters \vec{p} are Gaussian distributed around their central values \vec{q} and whose uncertainties are
 1581 described by a covariance V_{Cov} :

$$-2 \ln \lambda_{\text{Syst}} = \sum_{i,j} (p_i - q_i)(V_{\text{Cov}}^{-1})_{ij}(p_j - q_j). \quad (7.11)$$

1582 With this, the full statistic minimized by the fit is given by Equation 7.12.

$$-2 \ln \lambda = 2 \sum_j n_j \ln \frac{n_j}{y_j} + \sum_{i,j} (p_i - q_i)(V_{\text{Cov}}^{-1})_{ij}(p_j - q_j) \quad (7.12)$$

1583 A crucial step in the analysis is the extraction of true information (the set of true events
 1584 from which the number of signal interactions and slices which form the cross section calcula-
 1585 tion as in Section 7.3) from reconstructed quantities. In general, this is known as “unfolding”
 1586 and is a common problem within High Energy Physics [51]. Several unfolding techniques
 1587 exist, each with their own benefits and drawbacks (typically, a balance is made between bi-
 1588 ased results, bin-to-bin correlations, uncertainty, and smoothness) [52]. The fit done within
 1589 this analysis, known as a template fit, performs the role of unfolding. Included in the set of
 1590 parameters are a set of “template weights” assigned to the MC signal events which vary the
 1591 normalization of signal events in a given true energy bin, and which also have a subsequent
 1592 effect on the predicted reconstructed distributions. The fit simultaneously varies the tem-
 1593 plate weights and the other parameters, then compares the resulting predicted reconstructed
 1594 distributions to the measured distributions until it converges at a minimum $-2 \ln \lambda$ value.

1595 The role of the template parameters is highlighted in Equation 7.13, which shows the
 1596 relationship between the true and reconstructed events as predicted by MC. \hat{y}_i represents the
 1597 number of events in true bin i for the indicated true category (absorption, charge exchange,
 1598 muon background, or pion backgrounds). The events have a chance ϵ^k to be selected as
 1599 some selection category k when the reconstructed information is passed through the event
 1600 selection (described in Section 6). Reconstruction effects smear the events from some true

1601 bin i to some reconstructed bin j in selection category k . This is represented by $t_{i,j}^k$ which
1602 can be thought of as a “smearing matrix.” In general, ϵ^k and $t_{i,j}^k$ depend on the true category
1603 they act on. \hat{y}_i , ϵ^k and $t_{i,j}^k$ all depend on some subset of the fit parameters $\vec{\theta}$ and can be
1604 modified at each step in the fit. A parameter f^μ is used to vary the normalization of muons
1605 in the sample, as this is uncertain. Lastly, c_i^{Abs} , c_i^{Cex} are the template parameters that
1606 control the normalization of absorption and charge exchange events in true bin i . The sums
1607 extend over the number of true bins n_T for the different true categories. Since the number
1608 of impinging π^+/μ^+ is known and static, the fit is constrained as in Equation 7.14.

$$y_j^k = \sum_i^{n_T} c_i^{\text{Abs}} \hat{y}_i^{\text{Abs}} \epsilon^k t_{i,j}^k + \sum_i^{n_T} c_i^{\text{Cex}} \hat{y}_i^{\text{Cex}} \epsilon^k t_{i,j}^k + \sum_i^{n_T} f^\mu \hat{y}_i^\mu \epsilon^k t_{i,j}^k + \sum_l^{n_{\pi\text{BG}}} \sum_i^{n_T} \hat{y}_i^l \epsilon^k t_{i,j}^k \quad (7.13)$$

$$N = \sum_j y_j = \sum_j n_j \quad (7.14)$$

1609 In addition to the constraint on the overall number of incident particles, the number of
1610 incident particles in bins of true initial momentum (where it was generated by the beam
1611 event generator module), is also held constant. This has been omitted from Equation 7.14
1612 for clarity.

1613 Thus, the fit changes $\vec{\theta}$ and \vec{p} until the measured and predicted reconstruction distribu-
1614 tions best match. The result of the fit is a set of best-fit parameters $\vec{\theta}_0$ and \vec{p}_0 and their
1615 covariance which will be used for error propagation as described in the Section 7.5. The
1616 best-fit parameter values produce a set of modified MC events that can be used as in Section
1617 7.3 to extract cross sections.

1618 The fit uses the MIGRAD [53] routine of the Minuit2 [54] minimizer library within
1619 ROOT [55] to find the maximum likelihood ratio. The MIGRAD routine estimates the
1620 gradient of the likelihood ratio surface at each fit point and follows the gradient until it
1621 reaches the best-fit point. After finding the best-fit point, the HESSE routine within Minuit2
1622 is called. This computes the Hessian matrix: the second derivative of the $-2 \ln \lambda$ surface

1623 around the best fit point. The Hessian matrix is inverted to create the covariance matrix
 1624 which describes the post-fit uncertainties and correlations of the fit parameters.

1625 7.5 Error Propagation

1626 The output of the fit – the best-fit parameters $\vec{\phi}_0$ ³ and their associated covariance matrix
 1627 Σ – can be used to propagate the post-fit errors to the extracted cross sections. First, the
 1628 Cholesky decomposition [56] of the post-fit covariance matrix is computed. This representa-
 1629 tion of the covariance matrix (shown in Equation 7.15) is the product of an upper triangular
 1630 matrix R with positive diagonal elements and its transpose R^T .

$$\Sigma = R^T R \quad (7.15)$$

1631 A random set of fit parameters $\vec{\phi}_t$ (also known as a “throw”) can be generated by multiplying
 1632 a random unit Gaussian vector \vec{r}_t by R and adding this to the best-fit parameter values $\vec{\phi}_0$,
 1633 as shown in Equation 7.16. $\vec{\theta}_t$ will be randomly distributed with the same covariances of the
 1634 post-fit covariance matrix [56].

$$\vec{\theta}_t = \vec{\phi}_0 + R\vec{r}_t \quad (7.16)$$

1635 This procedure is repeated on the order of 1000 times to generate an ensemble of throws.
 1636 Each set of thrown parameters is used to calculate the cross section as described in Section
 1637 7.3. The cross section covariance matrix V is computed as in Equation 7.17, where V_{ij} is
 1638 the covariance between bins i and j , σ_{it} is the cross section in bin i for throw t and σ_{i0} is
 1639 the best-fit cross section in bin i . Note: the bins i, j include both absorption and charge
 1640 exchange to account for the covariances between these channels.

$$V_{ij} = \frac{1}{N} \sum_t^N (\sigma_{it} - \sigma_{i0})(\sigma_{jt} - \sigma_{j0}) \quad (7.17)$$

³The set of parameters $\vec{\phi}$ includes both the signal parameters $\vec{\theta}$ and systematic parameters \vec{p} .

1641 If any parameter is thrown into an unphysical region (i.e. for the template parameters,
1642 below zero), the throw is repeated until all parameters are within their allowed regions. This
1643 may results in truncated Gaussian distributions for any parameters that experience this
1644 issue. If this truncated area is small, the distribution is considered valid and has a negligible
1645 effect on the cross section covariance.

1646 This throwing procedure makes an assumption that the likelihood surface around the best
1647 fit point is distributed according to a multivariate Gaussian. If this assumption holds, the
1648 covariance matrix from the fit describes a multidimensional contour with constant χ^2 around
1649 the best-fit point which represents the probable spread of fit parameters. Additionally,
1650 the cross section covariance created by this propagation procedure describes a constant- χ^2
1651 contour centered around the best-fit cross section point [57][58].

CHAPTER 8

SYSTEMATIC UNCERTAINTIES

1652

1653 This chapter describes the systematic uncertainties and their implementation within the
1654 analysis. The uncertainties discussed stem from the dE/dx calibration, the reconstructed
1655 beam line momentum, the modeling (via Geant4) of the hadrons as they pass through the
1656 detector, the effect of the electron diverters on reconstructed track, and differences in the
1657 rate of both events without a reconstructed track and those failing the beam cuts. These
1658 uncertainties are parameterized within the fit and are constrained by a covariance within the
1659 systematic term given in Equation 7.11.

1660 8.1 dE/dX Calibration

1661 Section 4.5.3 describes how the measured charge per unit distance dQ/dx is translated
1662 into the energy deposited per unit distance dE/dx (which is used for the energy measure-
1663 ments of particles in this analysis). Part of this dE/dx extraction is the determination of a
1664 calibration constant C_{cal} by analyzing stopping muons. There is some uncertainty in what
1665 this calibration constant is, and as such, it has been implemented as a systematic parameter
1666 in the fit.

1667 As C_{cal} is varied within the fit, it has two large effects. The first is to change the MIP-
1668 like separation of daughter tracks during the event selection as described in Section 6.4, and
1669 the second is to migrate events between bins since more apparent energy will be accounted
1670 for in the energy reconstruction. This parameter was first implemented within the fit by
1671 rescaling the dE/dx in each step of the fit where the prediction histograms are refilled before
1672 comparing to data. This caused instability within the fit, as events would fail to migrate bins
1673 until the parameter was turned enough. This “threshold” behavior caused discontinuities in
1674 the $-2\ln\lambda$ surface, and so a different approach was opted for. Instead of implementing
1675 this effect directly on the events, a weighting scheme was implemented, where varied MC

1676 samples were created for various values of C_{cal} . In each bin of the prediction histograms,
 1677 the ratio to nominal was taken to form a weight for that bin and C_{cal} . These weights were
 1678 then interpolated between in order to form a smoothly-varying surface that could be used
 1679 within the fit. Each step of the fit, the events are given a weight which depends on the bin
 1680 the event falls into and the value of C_{cal} for that fit step.

1681 8.2 Beam Momentum

1682 Section 5.1.2 describes how the beam line instrumentation reconstructs momentum using
 1683 sets of fiber monitors surrounding a bending magnet in the beam line. This section also
 1684 mentions a bulk shift to the fibers in one of the monitors that affected the reconstructed
 1685 momentum. This shift was found to be $1.45 \pm 0.18 \text{mm}$. In addition to the uncertainty in the
 1686 shift is an estimated 1% uncertainty on the magnetic field. Recalling Equation 5.2 (repeated
 1687 here), these systematic uncertainties affect the reconstructed momentum as such: the fiber
 1688 shift varies θ , which can then cancel out a variation in B .

$$p = \frac{299.7924}{\theta} \times \int_0^{L_{\text{mag}}} (Bdl) \quad (5.2)$$

1689 These parameters would then be degenerate within the fit, and so these effects were combined
 1690 into a single momentum rescaling parameter c_p . The prior uncertainty on c_p is given by the
 1691 shifts to p due to variations in both parameters added in quadrature. The effect of the
 1692 variation to B is trivially 1%. For the effect of the fiber shift, the nominal beam line
 1693 MC simulation was ran with the fibers in the third monitor shifted by its 1σ uncertainty
 1694 (0.18mm). This results in an average 0.7% shift in the reconstructed momentum. The
 1695 uncertainty on c_p is thus given in Equation 8.1.

$$\sigma_{c_p} = \sqrt{.007^2 + .01^2} = .012 \quad (8.1)$$

1696 Within the analysis, the effect of this scaling parameter is to change the difference between

1697 true and reconstructed momentum r (defined in Equation 8.2).

$$r = \frac{p_{\text{Reco}} - p_{\text{True}}}{p_{\text{True}}} \quad (8.2)$$

1698 The beam simulation show this is Gaussian distributed with mean μ and width σ . Some
1699 variation to c_p will then result in a distribution with varied μ' and σ' . An event can then
1700 be given a weight according to its value of r and the μ' and σ' resulting from the value of
1701 c_p within one step of the fit. This weight is given in Equation 8.3, which is the ratio of two
1702 Gaussian distributions.

$$w = \frac{\sigma}{\sigma'} \exp\left(\frac{(r - \mu)^2}{2\sigma^2} - \frac{(r - \mu')^2}{2\sigma'^2}\right) \quad (8.3)$$

1703 The dependences of μ' and σ' on c_p were found from studies of the beam line MC simulation
1704 and used within the fit to form the weights as defined in 8.3.

1705 During fit validation, it was found that this beam momentum parameter created insta-
1706 bility in the fit due to its tendency to create extremely large weights for certain events at
1707 large parameter variations. This made it difficult to properly assess the post-fit error of the
1708 other parameters. As such, this parameter was chosen to be fixed during fits. Its pre-fit
1709 uncertainty was propagated to the cross section uncertainties by adding it in quadrature to
1710 post-fit parameter covariance matrix.

1711 8.3 Electron Diverter Effect

1712 As shown in Figures 6.1 and 8.1, the simulation of the grounded electron diverters (which
1713 causes tracks to prematurely break), differs from data. To account for the uncertainty in the
1714 strength of the track-breaking effect, a simple weighting scheme was developed to artificially
1715 vary the track-breaking strength.

1716 The weighting scheme varies the fraction of tracks ending above 222 cm, which end in
1717 the “track-breaking” region of 222–234 cm. This fraction, f , is defined as

$$f = \frac{N_{\text{Break}}}{N_{>222}} = \frac{N_{\text{Break}}}{N_{>234} + N_{\text{Break}}} \quad (8.4)$$

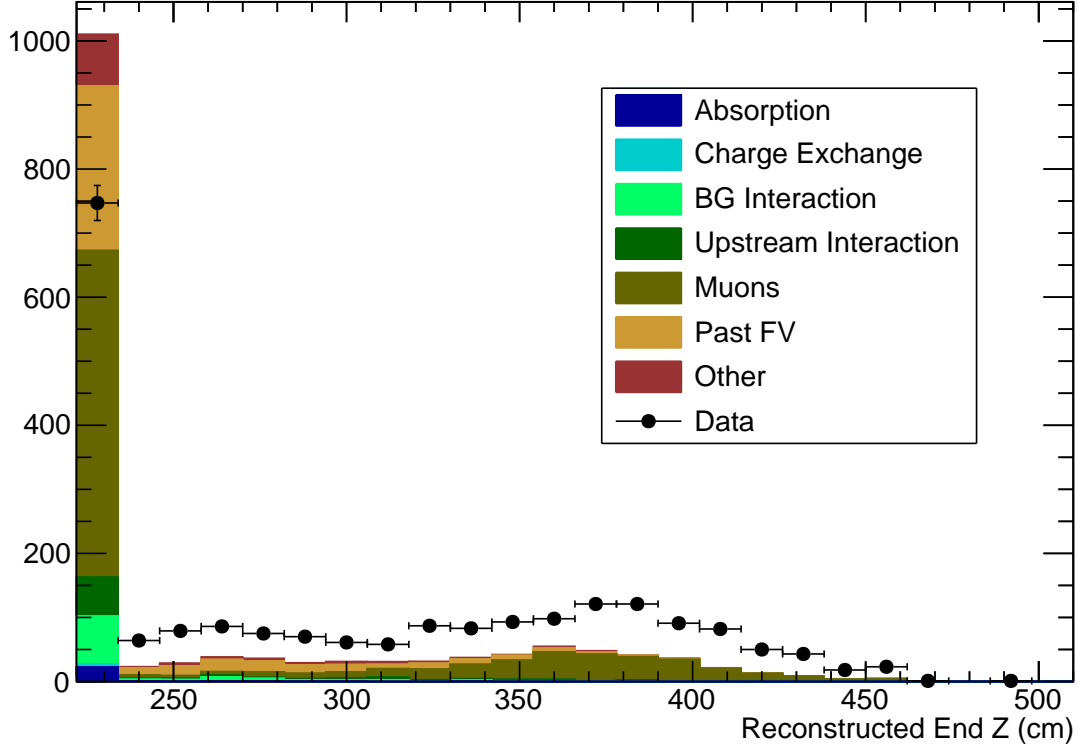


Figure 8.1: Enhanced view of the reconstructed endpoint of beam tracks within the TPC.

1718 where N_{Break} is the number of broken tracks (ending between 222 and 234 cm), and $N_{>222}$
 1719 and $N_{>234}$ are the number of tracks above 222 and 234 cm respectively. The probability
 1720 for a track ending above 222 cm to break is thus f , while the probability for a track to not
 1721 break is $1 - f$.

1722 Consider some variation as such: $f \rightarrow f' = cf$. Each track ending above 222cm is thus
 1723 given a weight as follows, depending on if was or was not broken.

$$W_{\text{Break}} = \frac{f'}{f} = \frac{cf}{f} = c \quad (8.5)$$

$$W_{>234} = \frac{1 - f'}{1 - f} = \frac{1 - cf}{1 - f} \quad (8.6)$$

1724 The nominal value of f , the fraction of events ending in the electron diverter region, in MC
 1725 is 0.6133. The central value of the scale factor c was set to 0.50 as taken from comparisons

1726 between MC and data shown here. The uncertainty on this was set naively to 20%.

1727 **8.4 Beam Efficiencies**

1728 It was found that the Pandora had an apparent difference between data and MC in
1729 the efficiency for identifying beam particles in the TPC. Additionally, shape differences in
1730 the position and direction of reconstructed beam tracks (possibly due to inaccuracies in
1731 mapping SCE as described in Section 4.5.1) created a difference in the fraction of events
1732 passing the beam cuts. These two uncertainties were parameterized as efficiency-like effects
1733 by varying the numbers of events in the following categories: 1) no reconstructed beam
1734 track, 2) reconstructed beam track that fails the beam cuts, 3) reconstructed beam track
1735 that passes the beam cuts. Let the fraction of events categorized as such be represented by
1736 f_1 , f_2 , and f_3 respectively. These fractions sum to one ($f_3 = 1 - f_2 - f_1$) and can be varied
1737 as follows.

Consider some variation to these fractions (these are, in effect, variations to the two efficiency-like effects):

$$\begin{aligned} f_1 &\rightarrow f'_1 = c_1 f_1 \\ f_2 &\rightarrow f'_2 = c_2 f_2 \\ f_3 &\rightarrow f'_3 = 1 - c_2 f_2 - c_1 f_1 \end{aligned} \tag{8.7}$$

Similar to the previous section, the events are given weights according to how they are categorized:

$$\begin{aligned} W_1 &= \frac{f'_1}{f_1} = c_1 \\ W_2 &= \frac{f'_2}{f_2} = c_2 \\ W_3 &= \frac{f'_3}{f_3} = \frac{1 - c_2 f_2 - c_1 f_1}{1 - f_2 - f_1} \end{aligned} \tag{8.8}$$

1738 The nominal values for the fraction of events with no track or failing the beam cuts in
1739 MC are 0.164 and 0.2305 respectively. The central value of the no-track parameter was set

1740 to 1.62 taken from comparisons to data and MC, and its uncertainty was naively set to 20%.
1741 The central value and uncertainty for the beam cut parameter were naively set to 1.00 and
1742 10% respectively.

1743 **8.5 Hadronic Interaction Modelling**

1744 In addition to uncertainties in the modeling of the detector systems described in the
1745 previous few sections, there are uncertainties in the hadronic interaction model. While
1746 the π^+ absorption and charge exchange interactions are measured by this analysis, the
1747 rate of background interactions (quasielastic, double charge exchange, production) can differ
1748 between data and MC as well. This can lead to wrongly estimated rates of categorization
1749 errors within the fit, and cause biased results of the signal interactions. The same is true
1750 of the rate of proton interactions as well. Protons are often emitted into the detector as
1751 a result of the primary π^+ -Ar interactions, and can go on to interact in the nearby argon,
1752 producing their own interaction products. These products can influence the event selection
1753 and produce categorization errors. Thus, differing rates of proton-argon interactions within
1754 data and MC can also bias the cross section results.

1755 To facilitate the propagation of hadronic modeling uncertainties related to the Geant4
1756 stage of the MC simulation (as discussed in Section 4.6), the Geant4Reweight [59] framework
1757 was used. This framework is able to create weights for events based on some variation applied
1758 to a cross section model in GEANT4. The weights created from this framework work by
1759 determining how likely the event was to occur given the nominal cross sections and the set
1760 of steps taken by a particle, and then comparing this to how likely the same event was to
1761 occur under some variation. The weights are generated under some flat scale factor applied
1762 over a user-defined region of momentum. The momentum regions and prior uncertainties
1763 for each variation were determined by a crude examination of the spread of models studied
1764 within Reference [27]. The description of the systematic parameters are given in Table 8.1.

1765 Geant4Reweight creates a weight for each parameter by running over each π^+ and proton

Channel	Momentum Range	Prior Uncertainty
π^+ Quasielastic	0–500 MeV/c	$\pm 36\%$
π^+ Quasielastic	500–2000 MeV/c	$\pm 33\%$
π^+ Pion Production	0–2000 MeV/c	$\pm 33\%$
π^+ Double Charge Exchange	0–2000 MeV/c	$\pm 33\%$
Proton Reaction	0–2000 MeV/c	$\pm 33\%$

TABLE 8.1: Description of the Geant4Reweight parameters used within the fit.

1766 created within the event and calculating a weight for that particle. These are all multiplied
1767 together to create full event weights. For each parameter, a weight is created at intervals
1768 of 10% from -90% to +100%. In order to create a smoothly varying effect within the fit,
1769 the variations must be interpolated between. Prior to the fit, sets of MC are produced
1770 at each variation for each parameter (note: only one parameter is varied at a time). For
1771 each truth category and reconstructed bin, the ratio between the varied and nominal MC
1772 are calculated and interpolated between using a spline. Then, when running the fit, each
1773 parameter contributes a weight to the event corresponding to the value of the spline at the
1774 parameter's value. All weights from all Geant4Reweight parameters are multiplied together
1775 when creating the predicted distributions for each step of the fit.

1776 8.6 Systematic Covariance Matrix

1777 Table 8.2 summarizes the size of the prior uncertainties that comprise the systematic
1778 covariance matrix. Note that all uncertainties described in this section are treated as uncor-
1779 related before the fit.

Parameter	Nominal Value	Prior Uncertainty
$dE/dX C_{cal}$	1.011×10^{-3}	$\pm 10\%$
Beam Momentum	1.00	$\pm 1.2\%$
Electron Diverter Fraction	0.5	± 0.20
No Track Fraction	1.62	± 0.20
Failed Beam Cuts Fraction	1.00	± 0.10
π^+ Quasielastic Low	1.00	$\pm 36\%$
π^+ Quasielastic High	1.00	$\pm 33\%$
π^+ Pion Production	1.00	$\pm 33\%$
π^+ Double Charge Exchange	1.00	$\pm 33\%$
Proton Reaction	1.00	$\pm 33\%$

TABLE 8.2: Description of the Geant4Reweight parameters used within the fit.

CHAPTER 9

FIT VALIDATION

1780

1781 This chapter demonstrates validation of the fit framework described in Section 7.4. It in-
1782 cludes the systematic uncertainties detailed in Chapter 8. In all tests, a set of MC simulation
1783 produced according to the 1 GeV/c beam setting is fit to various fake data inputs also pro-
1784 duced from MC simulation. These inputs could be the nominal MC or a set of varied MC.
1785 The specifics of the fake data will be described in each section.

1786 To evaluate how the fit performed, several quantities will be examined including the
1787 post-fit values of the parameters, the extracted cross sections, and a goodness of fit metric.
1788 Particular attention will be paid toward the post-fit values of the systematic parameters
1789 as they compare to their prior uncertainties. The goodness of fit will be investigated by
1790 comparing the minimum $-2 \ln \lambda$ (defined in Section 7.4) found by the fit in question to the
1791 distribution of minimum $-2 \ln \lambda$ found in a set of fits to systematically and statistically varied
1792 fake data. This comparison will take the form of a p-value, defined to be the probability
1793 of a fit resulting in a $-2 \ln \lambda_{\text{Min}}$ at least as extreme as the one in question. This is defined
1794 in Equation 9.1 where t_{Fit} represents the $-2 \ln \lambda_{\text{Min}}$ of the fit in question, and $f(t)$ is the
1795 distribution of $-2 \ln \lambda_{\text{Min}}$ found from the set of systematically and statistically varied fake
1796 data.

$$p = \int_{t_{\text{Fit}}}^{\infty} f(t) dt \quad (9.1)$$

1797 Figure 9.1 shows the distribution of $-2 \ln \lambda_{\text{Min}}$ from 1000 toy fits to systematically and
1798 statistically varied fake data. The systematic variations were created with the systematic
1799 parameters chosen according to the input covariance matrix (in a manner similar to the
1800 post-fit throws described in Section 7.5). Then, each set of systematically-varied fake data
1801 was statistically fluctuated. This distribution will be used throughout the following sections
1802 to determine p-values for each fit.

1803 Finally, the cross sections extracted from the post-fit MC will be compared to the cross

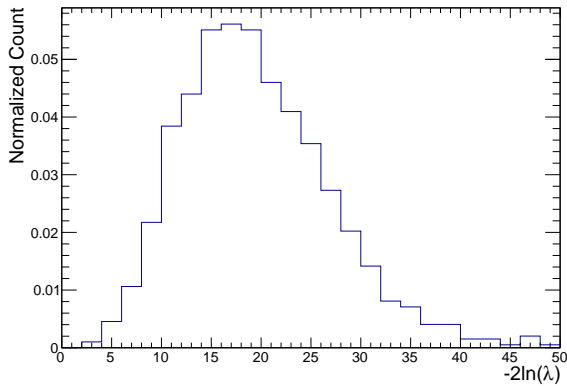


Figure 9.1: Post-fit $-2 \ln \lambda$ distribution of the toy experiments described above.

1804 sections as produced by the fake data input using the χ^2 defined in Equation 9.2.

$$\chi_{\sigma}^2 = \sum_{i,j} (\sigma_i - \bar{\sigma}_i)(V^{\sigma})_{i,j}^{-1}(\sigma_j - \bar{\sigma}_j) \quad (9.2)$$

1805 Here, σ_i represents the measured cross section in bin i , $\bar{\sigma}_i$ represents the cross section from
 1806 either the nominal MC or fake data input (this will be specified), and $(V^{\sigma})_{i,j}^{-1}$ is the value
 1807 of bin i, j of the inverted cross section covariance matrix as computed in the error prop-
 1808 agation procedure described in Section 7.5. This will be used similar to the minimum fit
 1809 statistic distribution discussed above to determine a p-value for the cross section results.
 1810 The distribution of χ_{σ}^2 from the set of 1000 toy fits is shown in Figure 9.2.

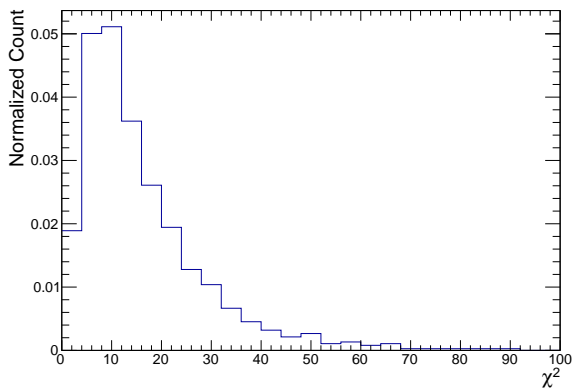


Figure 9.2: χ_{σ}^2 distribution of the toy experiments described above.

1811 **9.1 Asimov Fit**

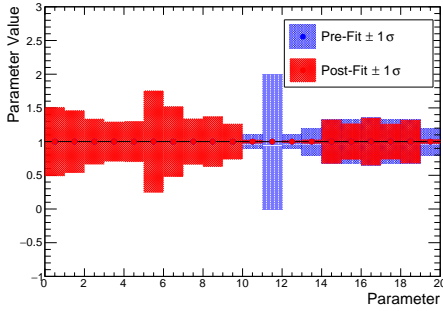
1812 The first validation test is a simple “Asimov” fit. In this fit, the input fake data is the
 1813 same as the nominal MC within the fit. This tests the base functionality of the fit and
 1814 whether or not the fit can correctly identify the minimum (the starting point of the fit). It
 1815 also shows the level of sensitivity the fit has for the signal and nuisance parameters. The
 1816 results are shown in Figures 9.3, 9.4, and 9.5. The first shows that the best-fit parameters
 1817 are at the starting point, as expected. The parameters in these plots are enumerated as in
 Table 9.1. This will be the same for the rest of the chapter.

0	Absorption factor 400–500 MeV/ <i>c</i>	10	Beam cut efficiency
1	Absorption factor 500–600 MeV/ <i>c</i>	11	Beam momentum resolution
2	Absorption factor 600–700 MeV/ <i>c</i>	12	<i>dE/dX</i> calibration constant
3	Absorption factor 700–800 MeV/ <i>c</i>	13	Electron diverter effect strength
4	Absorption factor 800–1000 MeV/ <i>c</i>	14	Geant4Reweight Double Charge Exchange
5	Charge Exchange factor 500–600 MeV/ <i>c</i>	15	Geant4Reweight Pion Production
6	Charge Exchange factor 600–700 MeV/ <i>c</i>	16	Geant4Reweight Quasielastic Low
7	Charge Exchange factor 700–800 MeV/ <i>c</i>	17	Geant4Reweight Quasielastic High
8	Charge Exchange factor 800–900 MeV/ <i>c</i>	18	Geant4Reweight Proton
9	Muon factor	19	No-track efficiency

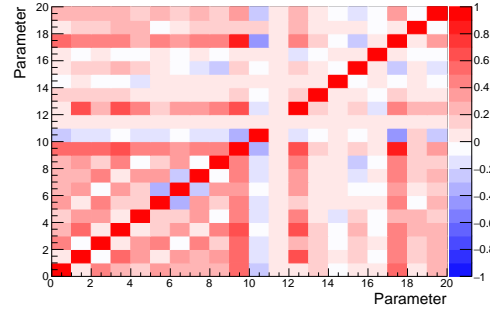
TABLE 9.1: The parameters used within the fit. The numbers correspond to the bins shown in the figures throughout the chapter.

1818

1819 The post-fit and nominal MC reconstructed distributions in Figure 9.4 are identical to
 1820 the Asimov fake data, and both distributions have a $-2 \ln \lambda$ of 0 with respect to the fake
 data as expected from this closure test.



(a) Pre-fit and post-fit parameters.



(b) Post-fit correlation matrix of the fit parameters.

Figure 9.3: Asimov fit results.

1821

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	0.00
Post-fit $-2 \ln \lambda_{\text{Stat}}$	0.00
Post-fit $-2 \ln \lambda_{\text{Syst}}$	0.00
Fit p-value	1.00
Nominal χ_{σ}^2	0.00
Fake Data χ_{σ}^2	0.00
Nominal σ p-value	1.00
Fake Data σ p-value	1.00

TABLE 9.2: Numerical results of the fit to Asimov fake data.

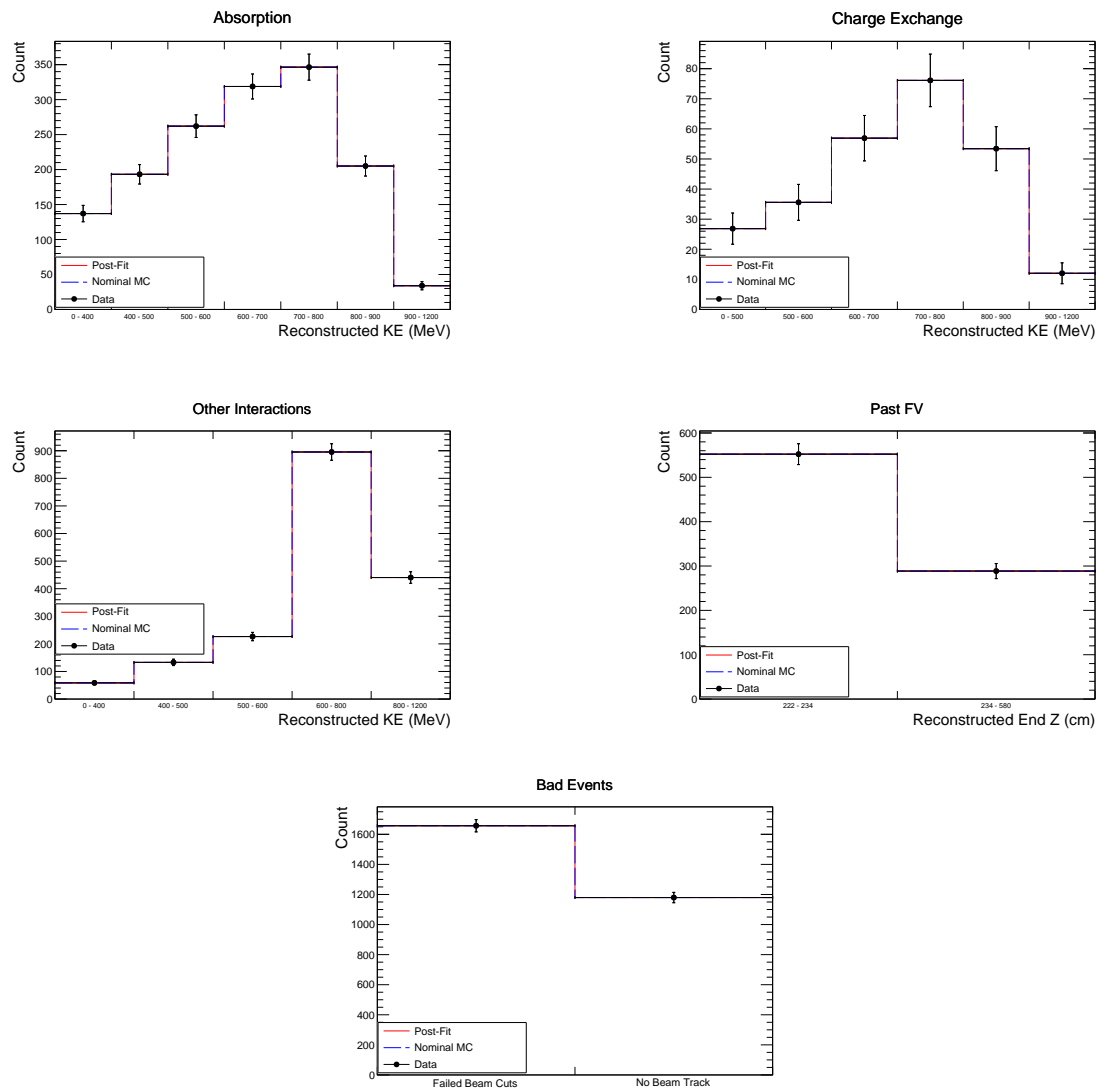
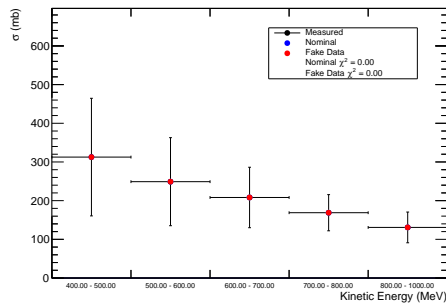
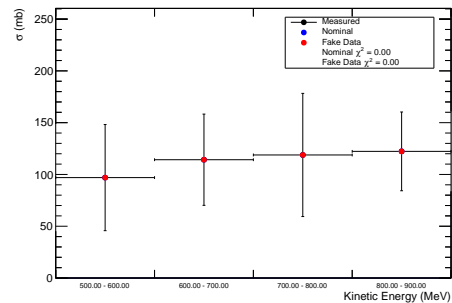


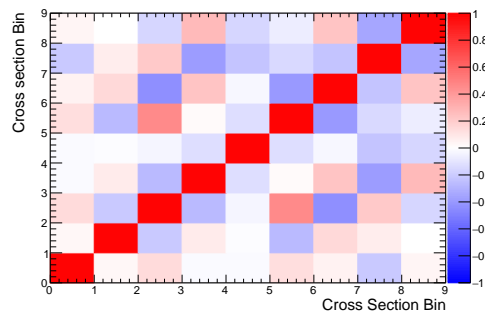
Figure 9.4: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the Asimov fit.



(a) Absorption



(b) Charge Exchange

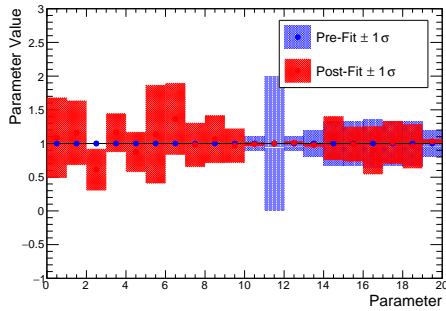


(c) Cross Section Correlations

Figure 9.5: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and Asimov Fake Data (red points). 9.5c is the correlation between the cross sections. The first five rows are the absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

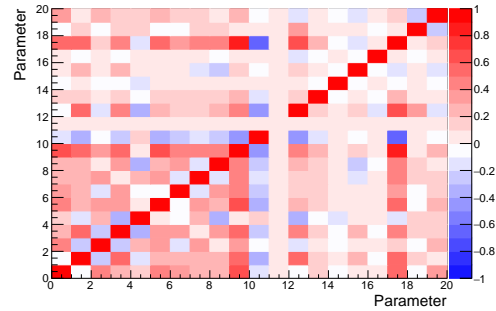
1822 9.2 Statistically Independent Nominal MC

1823 This test is similar to the previous Asimov fit where the input fake data is the nominal
 1824 MC. However, half of the nominal MC was used as the input fake data, and the other half was
 1825 used as the input MC. This is to test the performance of the fit to a statistically-independent
 1826 set of nominal MC. The input fake data is expected to deviate from the input MC by a normal
 1827 statistical fluctuation. This can be seen in Figure 9.7, where the fake data points no longer
 1828 lay directly on top of the input MC. As can also be seen in Table 9.3, the post-fit $-2 \ln \lambda$
 1829 between the post-fit and fake data reconstructed distributions is less than that between the
 1830 pre-fit and fake data distributions, as expected. In Figure 9.6b, the systematic parameters
 1831 can be seen to vary from nominal, but within the set of prior uncertainties presented within
 1832 the plot (as the blue bands). Finally, the χ^2_σ between the measured and fake data cross
 1833 sections as shown in Figure 9.8 shows the measured cross section is statistically consistent
 1834 with the cross section extracted from the fake data set.



(a) Pre-fit and post-fit parameters.

(b)



(c) Post-fit correlation matrix of the fit parameters.

(d)

Figure 9.6: Fit results for the statistically independent nominal MC fit.

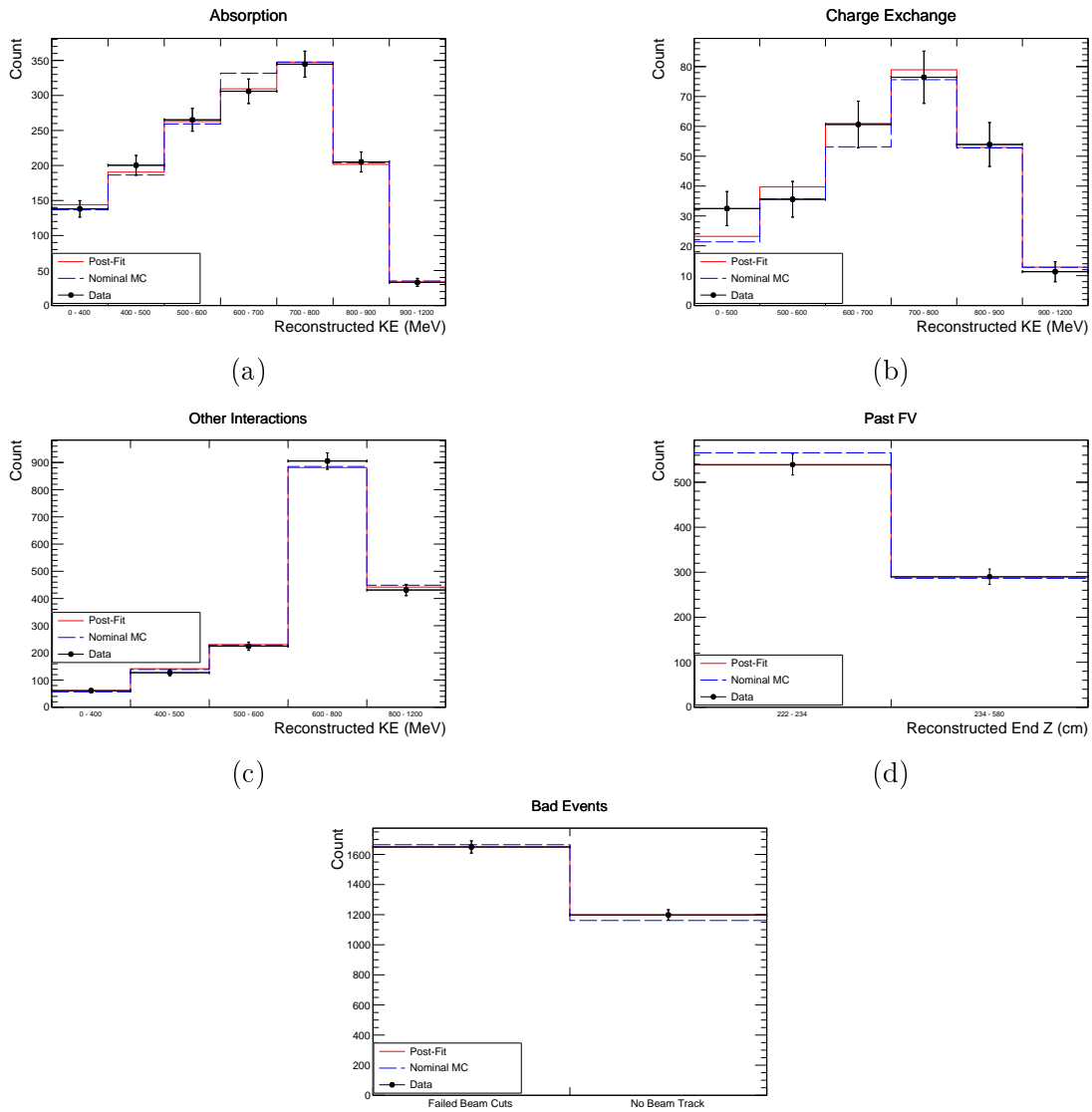
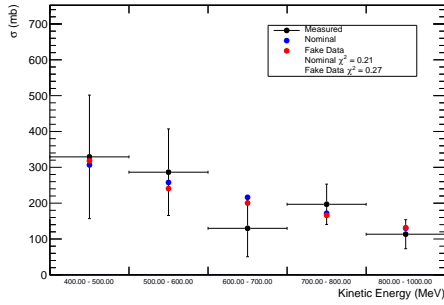
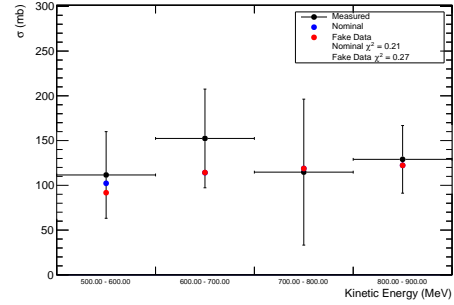


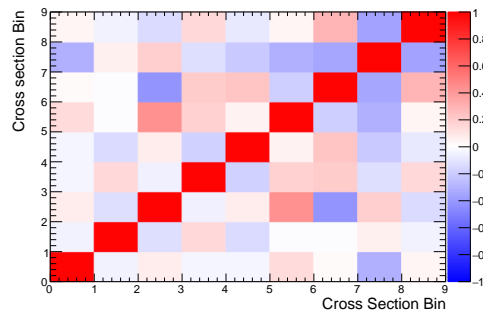
Figure 9.7: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the statistically independent nominal MC fit.



(a) Absorption



(b) Charge Exchange



(c) Cross Section Correlations

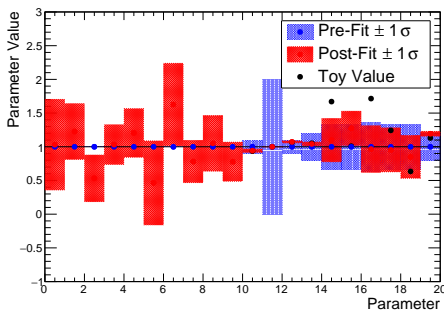
Figure 9.8: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and Fake Data produced from statistically independent nominal MC (red points). 9.8c is the correlation between the cross sections. The first five rows are absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	14.72
Post-fit $-2 \ln \lambda_{\text{Stat}}$	7.54
Post-fit $-2 \ln \lambda_{\text{Syst}}$	0.21
Fit p-value	0.97
Nominal χ_{σ}^2	0.21
Fake Data χ_{σ}^2	0.27
Nominal σ p-value	1.00
Fake Data σ p-value	1.00

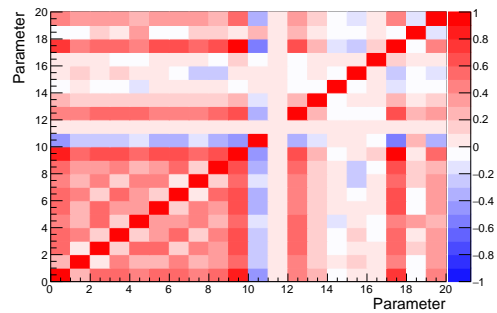
TABLE 9.3: Numerical results of the fit to statistically independent fake data.

1835 **9.3 Systematic Variation**

1836 In this test, the fake data has been generated by using a statistically independent set
 1837 of MC produced with varied systematic parameters. This is generated by first creating a
 1838 random set of systematic parameter values. To create these values, a vector of random, unit-
 1839 Gaussian distributed values is produced and then multiplied by the lower triangle of the
 1840 Cholesky decomposition of the prior covariance matrix of the systematic parameters. This
 1841 produces a set of values for the parameters with all correlations encoded. These systematic
 1842 parameter values are then applied to half of the MC sample. Both the fake data reconstructed
 1843 distributions and the cross sections are extracted from this varied MC sample. The other
 1844 half of the MC is used as the input MC to be varied within the fit. The results are shown in
 1845 Figures 9.9, 9.10, and 9.11. Figure 9.9a now includes the input systematic parameters used
 1846 to create the variation (labeled "Toy Values"). As can be seen in this figure, the post-fit
 1847 systematic parameters approach the input values. Shown in Table 9.4, the $-2 \ln \lambda$ between
 1848 MC and fake data show a large reduction as a result of the fit. Finally, in Figure 9.11 one
 1849 can see that the χ^2_{σ} between the measured and fake data cross sections shows a consistent
 fit result.



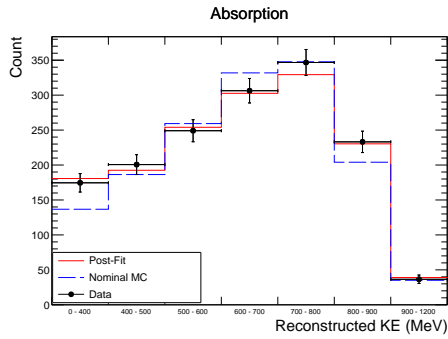
(a) Pre-fit and post-fit parameters.



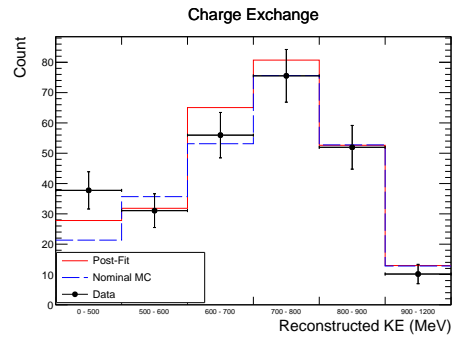
(b) Post-fit correlation matrix of the fit parameters.

Figure 9.9: Fit results for the systematically varied fit.

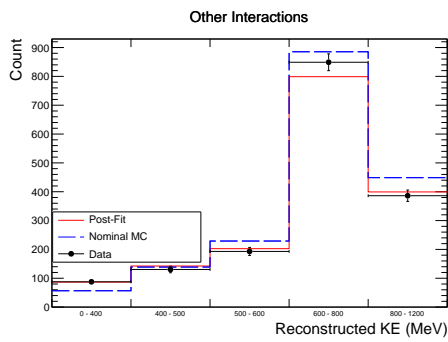
1850



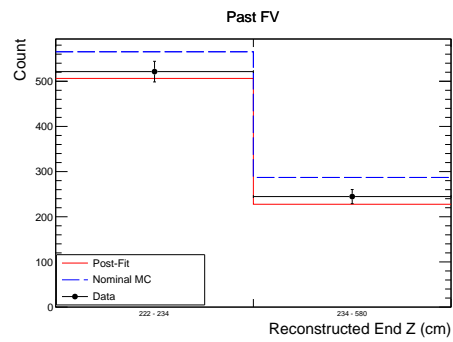
(a) Reconstructed distribution of events selected as Absorption.



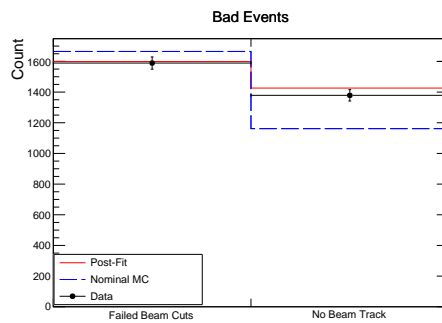
(b) Reconstructed distribution of events selected as Charge Exchange.



(c) Reconstructed distribution of events selected as neither Absorption or Charge Exchange.

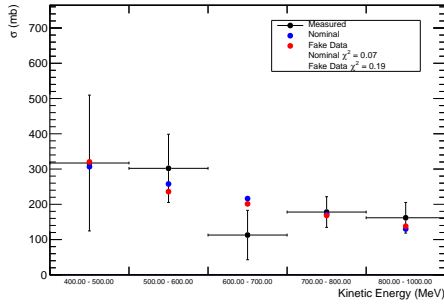


(d) Reconstructed distribution of events which extend past the Fiducial Volume

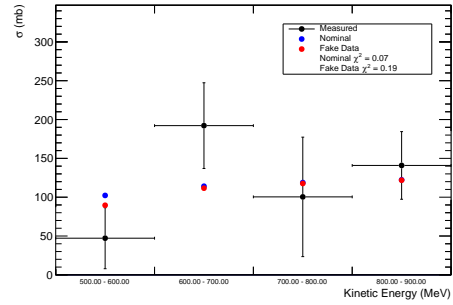


(e) The number of events which fail beam cuts (left bin) or lack a reconstructed beam track (right bin).

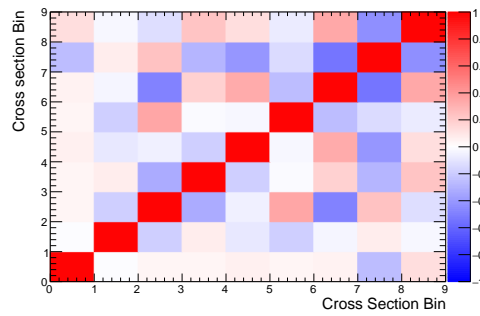
Figure 9.10: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the systematically varied fit.



(a) Absorption



(b) Charge Exchange



(c) Cross Section Correlations

Figure 9.11: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and systematically varied Fake Data (red points). 9.11c is the correlation between the cross sections. The first five rows are absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	155.87
Post-fit $-2 \ln \lambda_{\text{Stat}}$	12.69
Post-fit $-2 \ln \lambda_{\text{Syst}}$	2.84
Fit p-value	0.69
Nominal χ_{σ}^2	0.07
Fake Data χ_{σ}^2	0.19
Nominal σ p-value	1.00
Fake Data σ p-value	1.00

TABLE 9.4: Numerical results of the fit to systematically and statistically varied fake data.

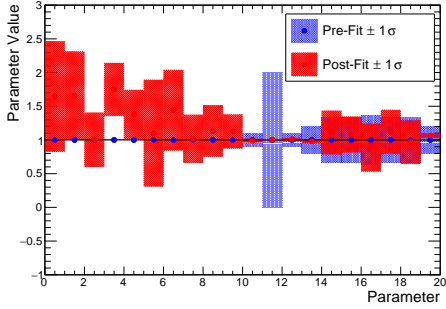
1851 9.4 Geant4Reweight Fake Data

1852 For this test, fake data is produced by reweighting¹ half of the nominal MC according
1853 to some set of π^+ -Ar and p -Ar cross section variations using Geant4Reweight. Three sets of
1854 fake data were created. The first set was created varying the signal cross sections by some
1855 “reasonable” amount (i.e. similar to the level of the prior uncertainties of the Geant4Reweight
1856 parameters). The second set was created by varying the signal cross sections by an amount
1857 larger than the prior uncertainties on the Geant4Reweight parameters. The final set was
1858 created by varying both the signal and background cross sections. The background cross
1859 sections were varied in a different parameterization than those used in the fit: the bins of the
1860 Geant4Reweight variations in the fake data did not align with the bins in the fit parameters.

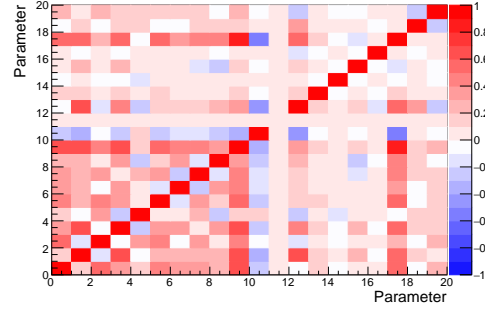
1861 9.4.1 Reasonable Variations

1862 The first set of fake data was created with the absorption cross section increased by 30%
1863 and the charge exchange cross section reduced by 10% across the full MC momentum range.
1864 Shown in Figure 9.12a, the systematic parameters are kept within their prior uncertainties.
1865 The reconstructed distributions in Figure 9.13 shows the fit ends in good agreement with
1866 the fake data distributions as can be seen in the post-fit $-2 \ln \lambda$. Finally, the cross section
1867 extracted from the fit agree quite well with the cross sections extracted from the fake data
1868 set as can be seen in the "Fake Data χ^2 " in Figure 9.14.

¹A process to produce varied Monte Carlo samples assuming alternate cross section models, described in Section 8.5.



(a) Pre-fit and post-fit parameters.



(b) Post-fit correlation matrix of the fit parameters.

Figure 9.12: Fit results for the reasonable-variation Geant4Reweight fake data fit.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	30.34
Post-fit $-2 \ln \lambda_{\text{Stat}}$	7.65
Post-fit $-2 \ln \lambda_{\text{Syst}}$	0.54
Fit p-value	0.97
Nominal χ_{σ}^2	3.58
Fake Data χ_{σ}^2	0.45
Nominal σ p-value	0.94
Fake Data σ p-value	1.00

TABLE 9.5: Numerical results of the fit to reasonable-variation Geant4Reweight fake data.

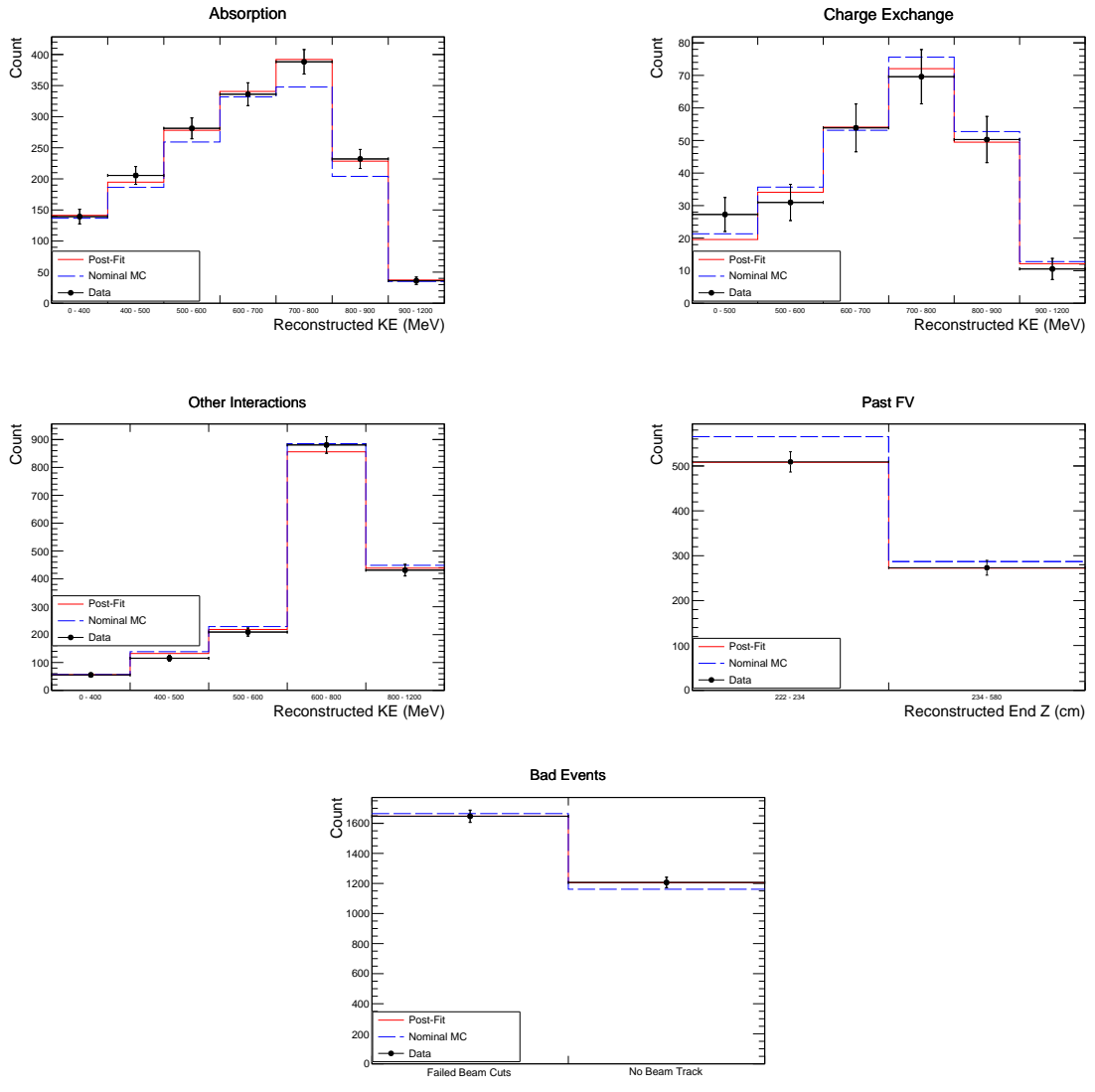
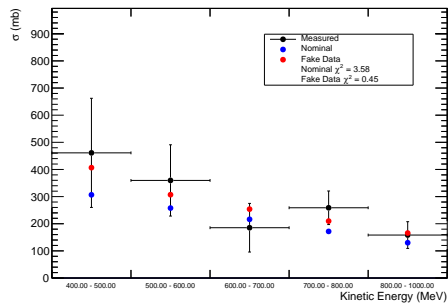
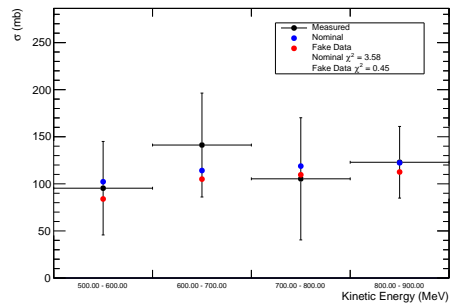


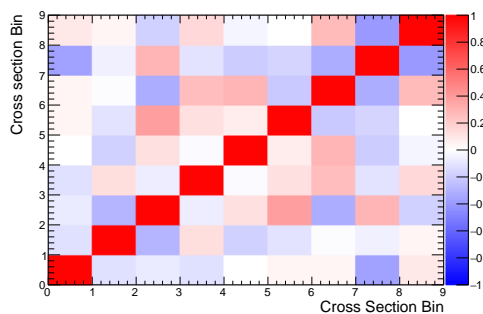
Figure 9.13: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the reasonable Geant4Reweight fake data fit.



(a) Absorption



(b) Charge Exchange



(c) Cross Section Correlations

Figure 9.14: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and less extreme Geant4Reweight Fake Data (red points). 9.14c is the correlation between the cross sections. The first five rows are absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

1869 **9.4.2 Plausible Variations**

1870 The second set of fake data created with Geant4Reweight contained an increase to the
 1871 absorption cross section by 80% and a reduction of the charge exchange cross section across
 1872 the full momentum range by 60%. Shown in Table 9.6, the fit ends with a consistent $-2 \ln \lambda$
 1873 the post-fit nuisance parameters are within their prior uncertainties showing a successful fit
 1874 to the fake data. In Figure 9.17, a drastic reduction in χ^2_σ is shown, indicating the fit can
 1875 successfully pick out a variation to the signal cross sections at this level.

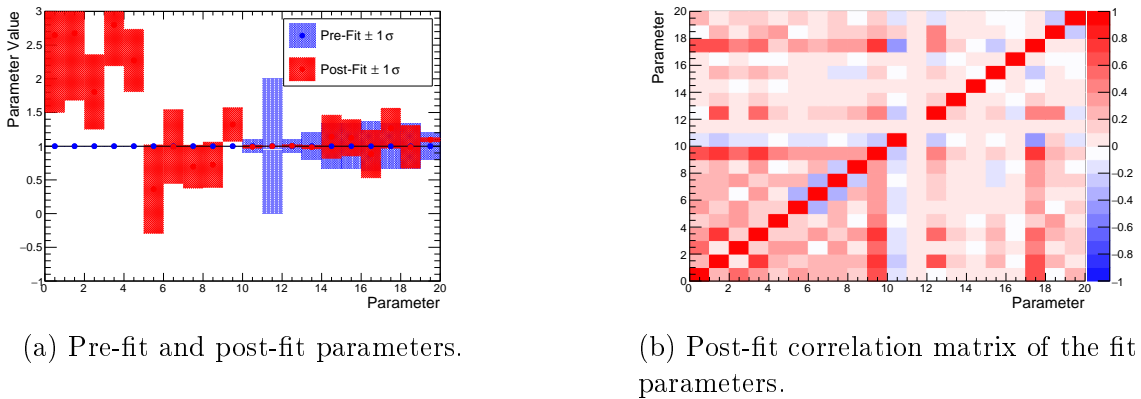


Figure 9.15: Fit results for the plausible-variation Geant4Reweight fake data fit.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	151.75
Post-fit $-2 \ln \lambda_{\text{Stat}}$	8.01
Post-fit $-2 \ln \lambda_{\text{Syst}}$	1.33
Fit p-value	0.94
Nominal χ^2_σ	20.60
Fake Data χ^2_σ	0.92
Nominal σ p-value	0.29
Fake Data σ p-value	1.00

TABLE 9.6: Numerical results of the fit to plausible-variation Geant4Reweight fake data.

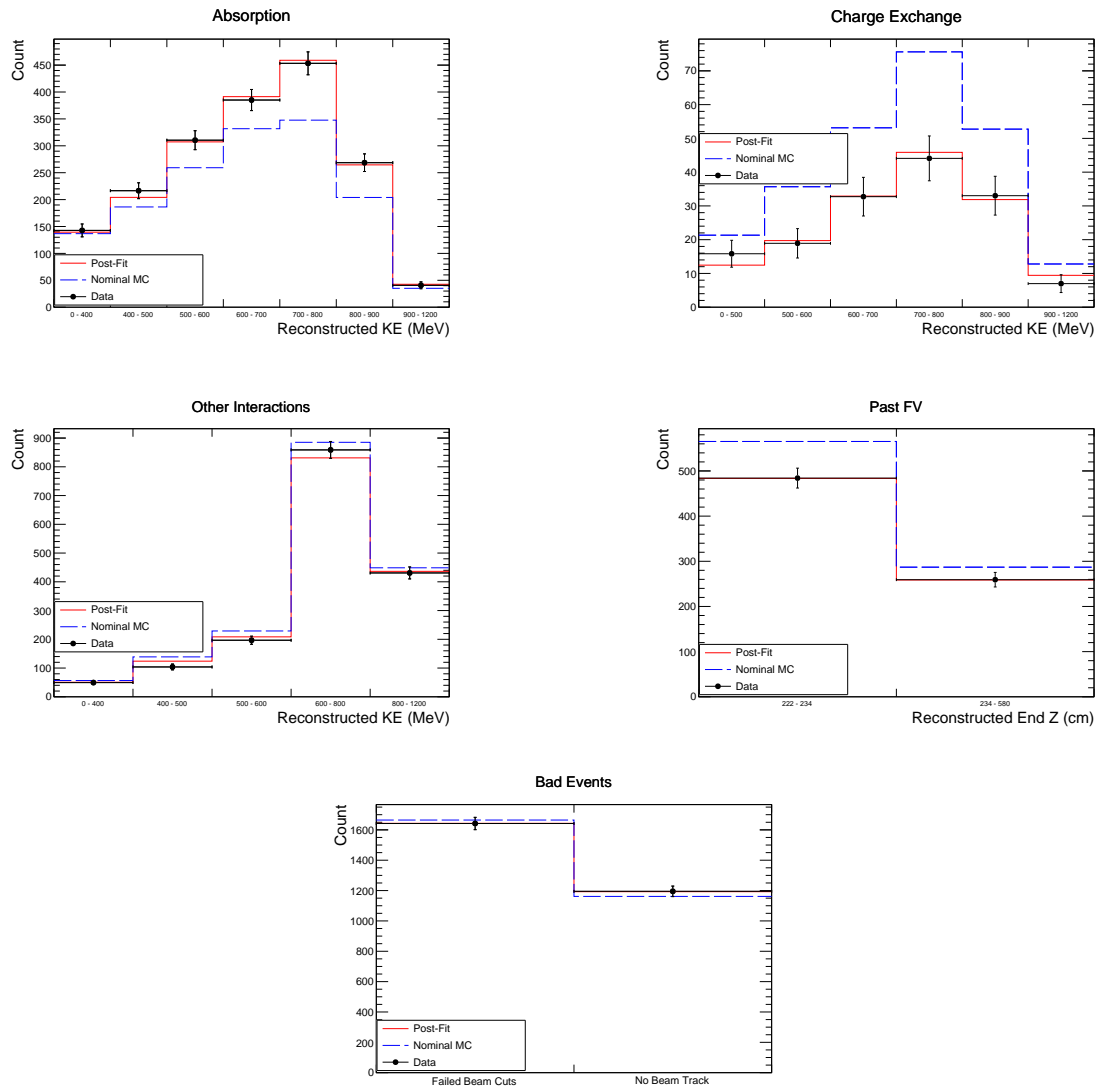
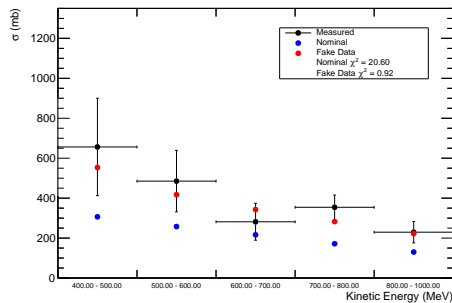
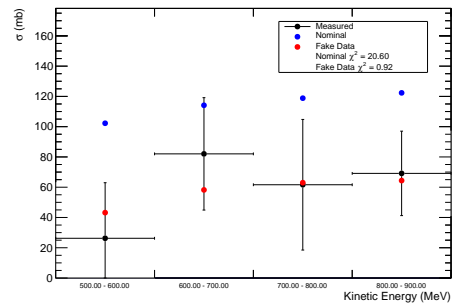


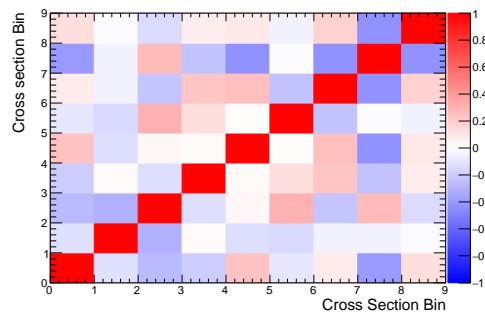
Figure 9.16: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the plausible-variation Geant4Reweight fake data fit.



(a) Absorption



(b) Charge Exchange

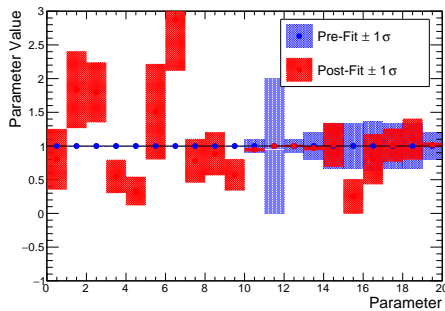


(c) Cross Section Correlations

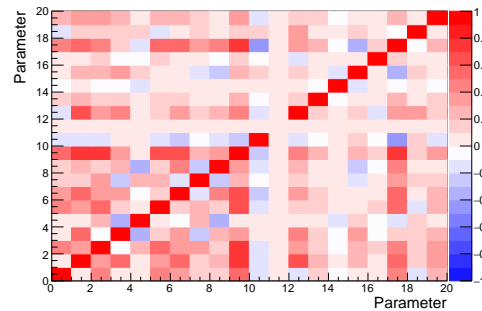
Figure 9.17: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and plausible-variation Geant4Reweight Fake Data (red points). 9.17c is the correlation between the cross sections. The first five rows are absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

1876 **9.4.3 Extreme Variations**

1877 The final set of fake data created with Geant4Reweight weights was intended to represent
 1878 “extreme” variations. The variations applied to the MC are to increase the total inelastic
 1879 cross section by 80% up to 800 MeV/c momentum and to reduce the total inelastic cross
 1880 section by 60% above 800 MeV/c. As can be seen in the Figures 9.18 - 9.20, the fit finds a
 1881 minimum, but the results indicate a poor result. A few of the systematic parameters shown
 1882 in Figure 9.18a are pulled outside of their prior uncertainties. The post-fit $-2\ln\lambda$ in Table
 1883 9.7 is large, indicating a bad goodness-of-fit. Similarly, the χ^2 between the measured cross
 1884 sections and those extracted from fake data is actually *higher* than that to the nominal MC.
 1885 This suggests that the parameterization used within the fit is not suitable for this fake data.
 1886 As described in Section 8.5, the systematic parameter for pion production is a single bin from
 1887 0 to 2 GeV/c, whereas the variation used to create the fake data has a more complicated
 1888 shape to it. This is an example of where the fit on data could fail to find a useful result. If this
 1889 was seen in a fit to data, we would reconsider the parameterization of the Geant4Reweight
 1890 parameters (i.e. the coarseness of the variation bins).



(a) Pre-fit and post-fit parameters.



(b) Post-fit correlation matrix of the fit parameters.

Figure 9.18: Fit results for the extreme Geant4Reweight fake data fit.

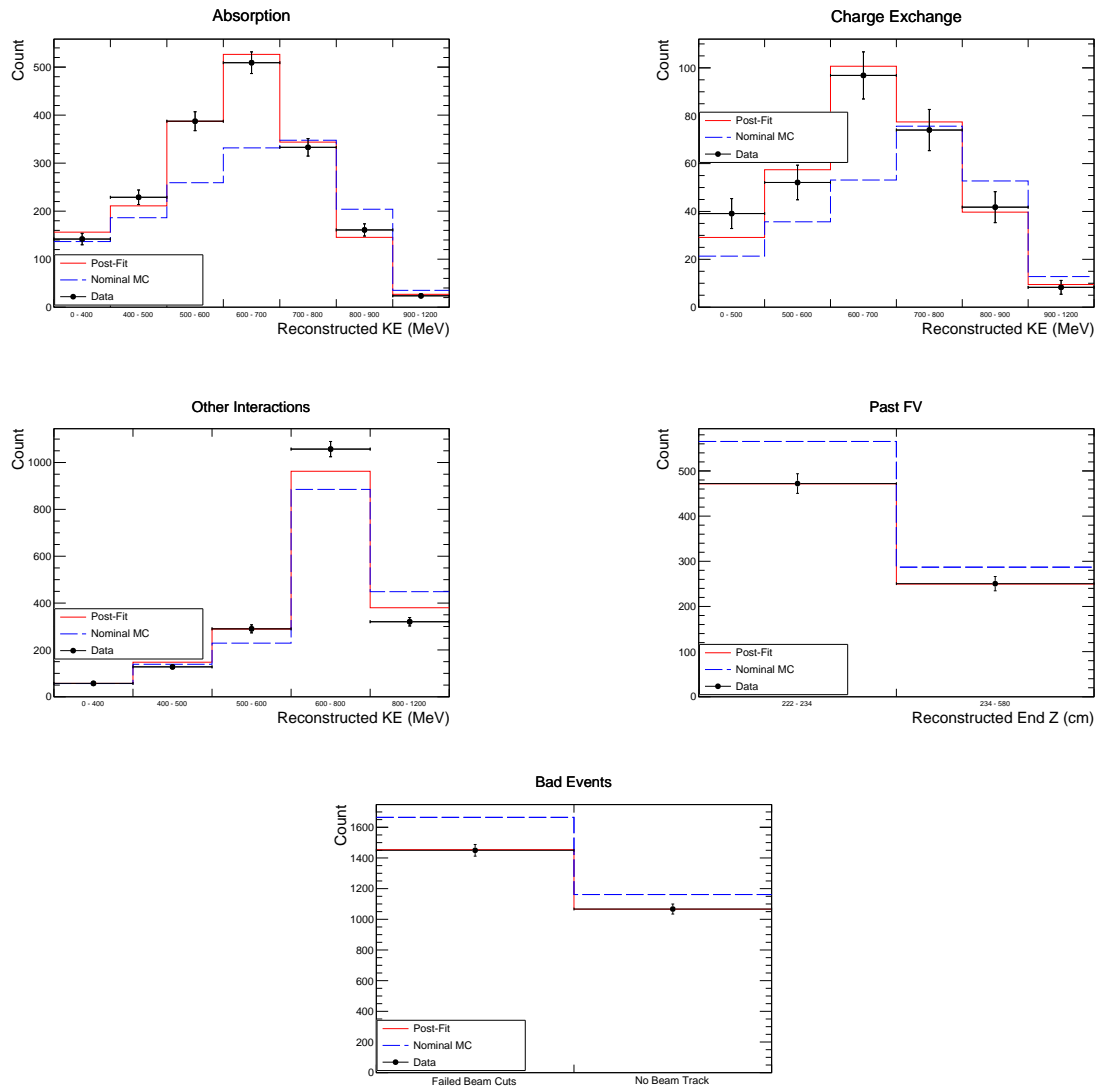
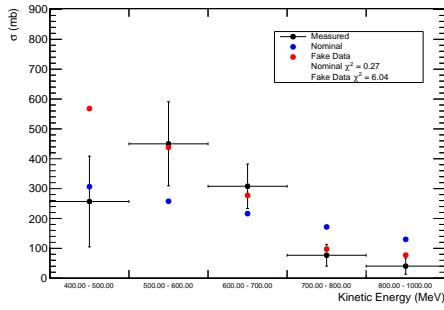
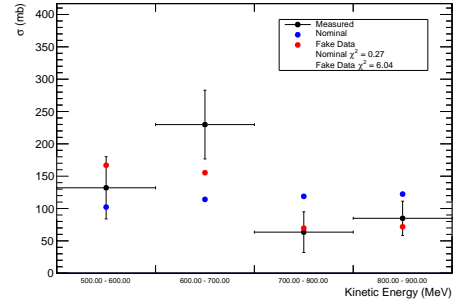


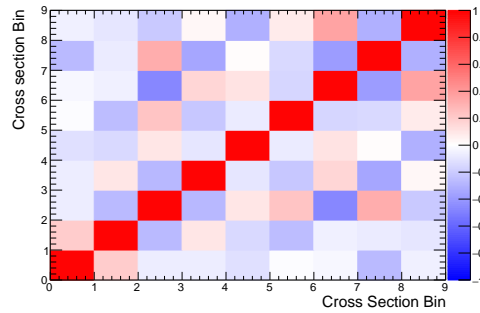
Figure 9.19: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the extreme Geant4Reweight fake data fit.



(a) Absorption



(b) Charge Exchange



(c) Cross Section Correlations

Figure 9.20: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and extreme Geant4Reweight Fake Data (red points). 9.20c is the correlation between the cross sections. The first five rows are absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	358.02
Post-fit $-2 \ln \lambda_{\text{Stat}}$	31.62
Post-fit $-2 \ln \lambda_{\text{Syst}}$	5.87
Fit p-value	0.03
Nominal χ_{σ}^2	0.27
Fake Data χ_{σ}^2	6.04
Nominal σ p-value	1.00
Fake Data σ p-value	0.84

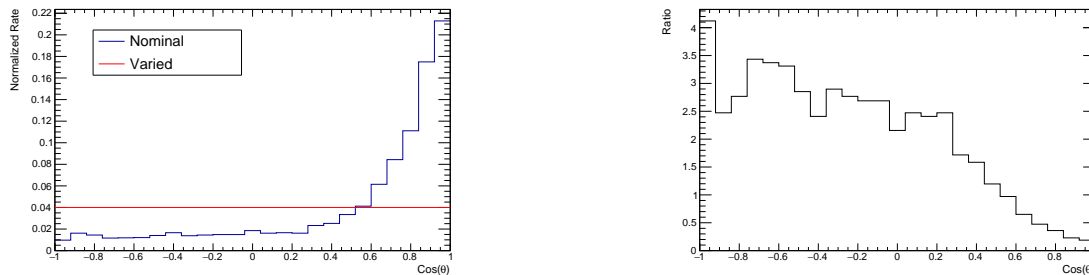
TABLE 9.7: Numerical results of the fit to unreasonable-variation Geant4Reweight fake data.

1891 9.5 Pion Angle Variation

1892 The next set of fake data used to validate the fit considers a variation to the outgoing
1893 direction of pions produced in quasielastic (QE) events. Note that this is defined according
1894 to the signal definition, and the outgoing pion is required to have above 150 MeV/c for the
1895 event to be considered QE. This is done to test whether the fit is resilient to mismodeling of
1896 outgoing pion kinematics in the Geant4 model used. The angular distribution of outgoing
1897 pions in QE events was modified by-hand to create a varied MC sample. This is done in bins
1898 of true momentum at interaction in order to prevent some variation to be imparted on the
1899 distribution of pion momentum at interaction. The bin edges for these are (0, 400, 600, 800,
1900 1000, 2000) MeV/c. In each final momentum bin, the ratio of the varied distribution to the
1901 nominal distribution is used as an event weight to create a set of fake data. Cross sections
1902 from the set of varied MC used to create the fake data are extracted and compared to the
1903 post-fit MC.

1904 9.5.1 Flat Distribution

1905 In this set of fake data, the angular distribution of final state pions is flattened. The nominal
1906 distribution, varied distribution, and ratio used for weighting from the 600–800 MeV/c bin
1907 are shown in Figure 9.21.



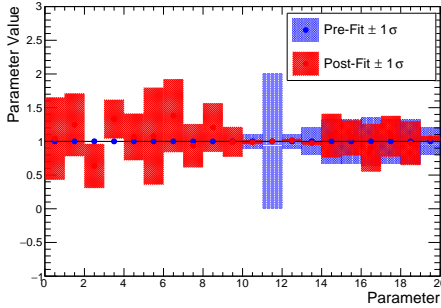
(a) Outgoing pion angular distributions.

(b) Ratio used for weighting QE events.

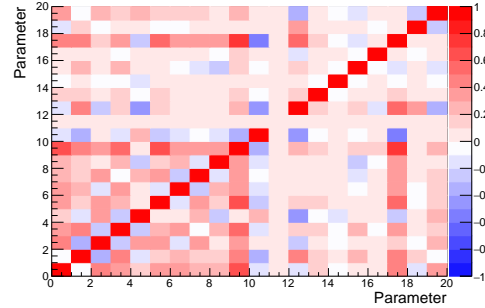
Figure 9.21: Inputs for this fake data test.

1908 As can be seen in Figure 9.23, small variations occur throughout the distributions. The

1909 most interesting bin is the lowest bin in Figure 9.23d, as it can be explained from the
 1910 variation applied. This bin contains QE events with a forward-going pion. Since these
 1911 forward-going events have been suppressed in the fake data, this bin has been lowered. Still,
 1912 the results of this fit are promising. The post-fit $-2 \ln \lambda$ in Table 9.8 shows that the fit is
 1913 insensitive to variation, indicating a robustness against data-MC disagreements of this type.
 1914 The small variations in the bins mentioned previously show low sensitivity to this type of
 1915 physical variation. Furthermore, the χ^2 between the Measured and Nominal & Fake Data
 1916 cross sections in Figure 9.24 show a consistent measurement.



(a) Pre-fit and post-fit parameters.



(b) Post-fit correlation matrix of the fit parameters.

Figure 9.22: Fit results for the flat pion fit.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	18.56
Post-fit $-2 \ln \lambda_{\text{Stat}}$	6.61
Post-fit $-2 \ln \lambda_{\text{Syst}}$	0.33
Fit p-value	0.98
Nominal χ^2_{σ}	0.31
Fake Data χ^2_{σ}	0.55
Nominal σ p-value	1.00
Fake Data σ p-value	1.00

TABLE 9.8: Numerical results of the fit to varied pion angular distribution fake data.

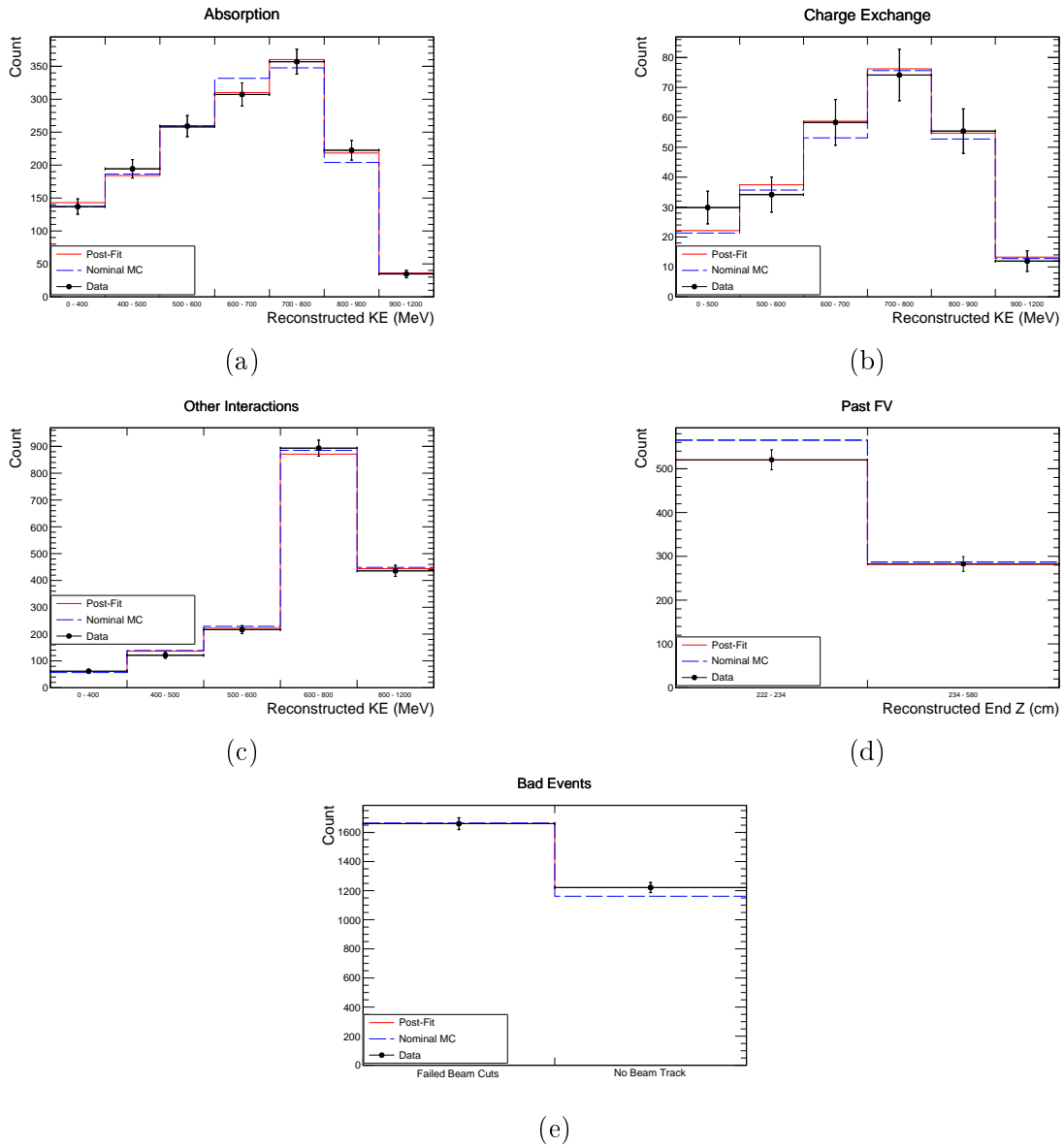
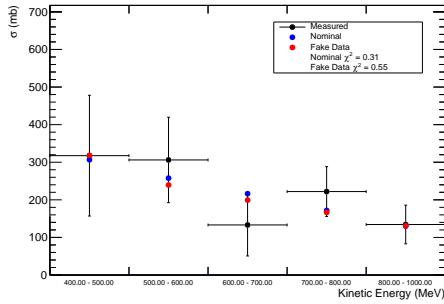
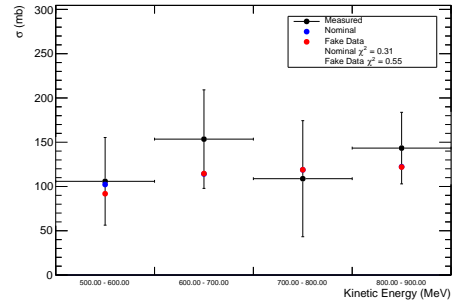


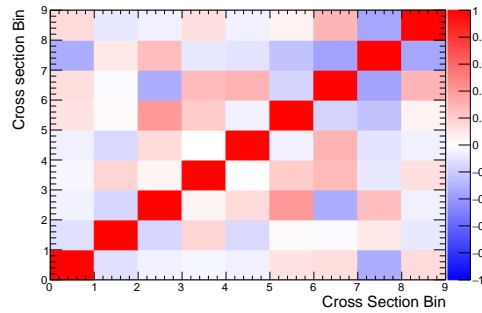
Figure 9.23: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the flat pion fit.



(a) Absorption



(b) Charge Exchange



(c) Cross Section Correlations

Figure 9.24: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points), Nominal (blue points), and flat pion Fake Data (red points). 9.24c is the correlation between the cross sections. The first five rows are absorption, and the last four rows are charge exchange. Note that the correlations between the two cross section types are included.

CHAPTER 10

RESULTS

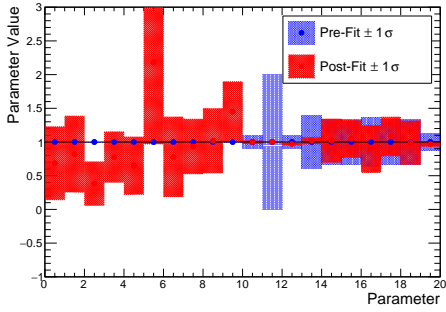
1917

1918 This chapter shows the results of the fit to ProtoDUNE-SP beam data from Run 5809 taken
1919 during its initial running period in the Fall of 2018. The data is comprised of beam-triggered
1920 events with a PID corresponding to π/μ (as defined in Section 5.2). The pre-fit and post-fit
1921 parameters and their corresponding uncertainties are shown in Figure 10.1a along with their
1922 post-fit correlation matrix. The pre-fit (referred to as “Nominal”) MC, post-fit MC, and data
1923 event distributions are shown in Figure 10.2. Lastly, the cross sections extracted from the
1924 fit to data are shown in Figure 10.3. This figure contains the correlation matrix and 1-D
1925 error bars (taken from the diagonal of the covariance matrix) as computed from the error
1926 propagation procedure described in Section 7.5.

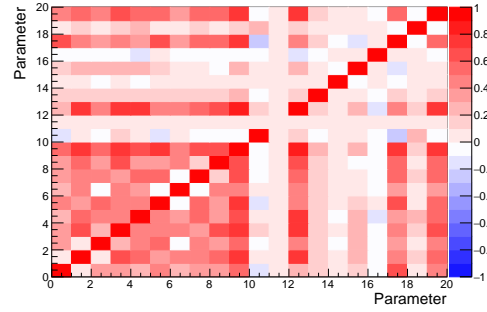
1927 Figure 10.1a shows the effect on the systematic parameters as a result of the fit. The
1928 dE/dX calibration and beam cut efficiency parameters are tightly constrained within the
1929 fit, but remain within their prior uncertainties. The Geant4Reweight parameters remain
1930 unconstrained by the fit.

1931 The event distributions shown in Figure 10.2 show good agreement between the post-fit
1932 MC and data distributions. The resulting p-value of the fit is 0.998, pointing to a successful
1933 parameterization of the fit. However, it suggests that the fit could be “too good”. This could
1934 be due to the choice of parameterization of the fit (perhaps the effects of the efficiency-
1935 like systematic parameters are too strong), or it could be due to the size of the post-fit
1936 uncertainties being too large.

1937 Finally, the extracted cross sections shown in Figure 10.3 remain consistent with the
1938 nominal-MC cross section, as indicated by the χ^2_σ of 6.70 shown in these plots. The σ p-
1939 value of this is 0.81, indicating the post-fit cross sections remain consistent with the nominal
1940 cross sections.



(a) Pre-fit and post-fit parameters.



(b) Post-fit correlation matrix of the fit parameters.

Figure 10.1: Fit results for the fit to ProtoDUNE-SP data.

Pre-fit $-2 \ln \lambda_{\text{Stat}}$	874.37
Post-fit $-2 \ln \lambda_{\text{Stat}}$	4.14
Post-fit $-2 \ln \lambda_{\text{Syst}}$	0.31
Fit p-value	0.998
Nominal χ^2_{σ}	6.70
Nominal σ p-value	0.81

TABLE 10.1: Numerical results of the fit to ProtoDUNE-SP run 5809 data.

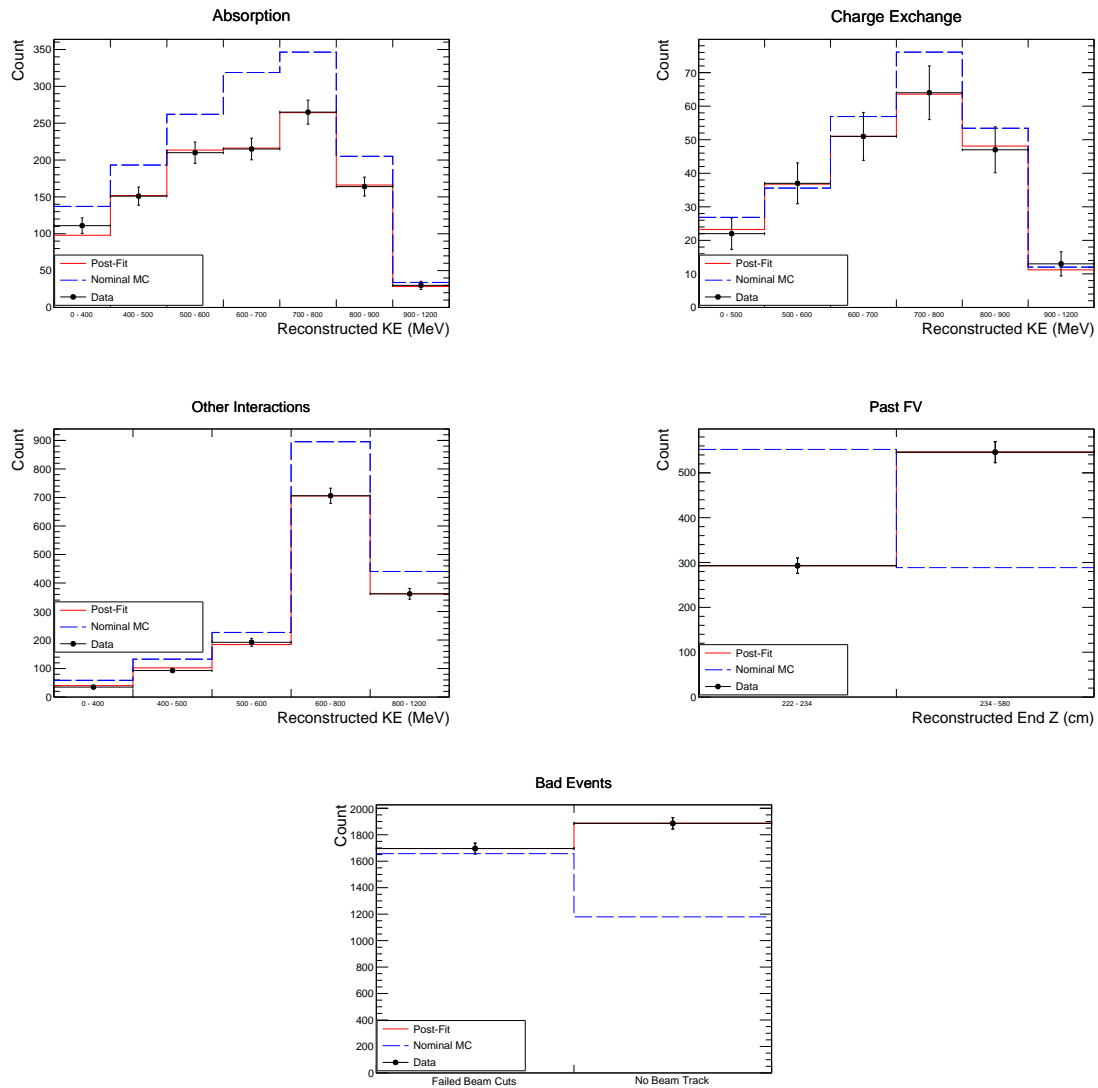
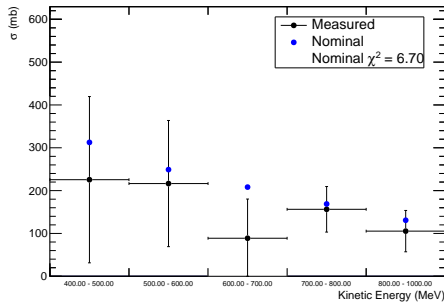
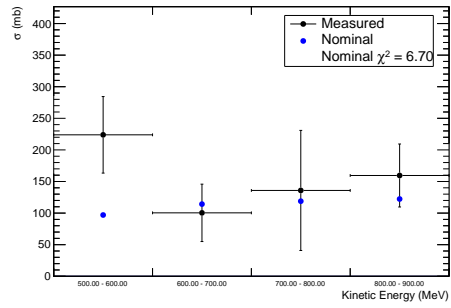


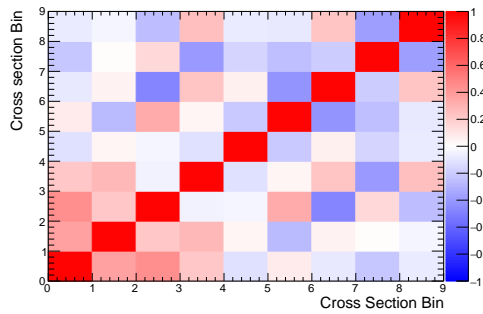
Figure 10.2: Reconstructed distributions of events in data (black points), Nominal MC (blue histogram), and post-fit results (red histogram) for the fit to real data.



(a) Absorption



(b) Charge Exchange



(c) Cross Section Correlations

Figure 10.3: Cross sections extracted from truth information taken from the post-fit MC ("Measured", black points) and Nominal (blue points). 10.3c is the correlation between the cross sections. The first four rows are the absorption, and the last three rows are charge exchange. Note that the correlations between the two cross section types are included.

1941 **10.1 Future Work**

1942 Some shortcomings within the analysis are worth addressing here. The cross sections
1943 shown in this chapter have relatively large error bars, and remain compatible with the nomi-
1944 nal MC at this level of uncertainty. Though this compatibility is not an issue, a reduction in
1945 the measurement's uncertainty would allow us to determine if there is a significant difference
1946 to the cross section models used in DUNE's simulation. A larger data set (such as more
1947 runs taken during Fall 2018) would of course reduce the statistical uncertainties, while a
1948 better understanding of the underlying cause of the Pandora reconstruction efficiency would
1949 improve the systematic uncertainties. The improved understanding of what is causing the
1950 data-MC discrepancies regarding the reconstruction efficiency could provide a more suitable
1951 uncertainty parameterization than the ad-hoc efficiency factors currently used in the fits.
1952 We also chose to neglect SCE uncertainties in this analysis, though these are in development
1953 and will be added in the future. Following the implementation of the SCE uncertainties, the
1954 other runs taken in Fall 2018 will be added. This will increase the size of the data set in
1955 the fit by a factor of 9. These issues will all be iterated upon in future work as this analysis
1956 moves toward publication by the DUNE collaboration. Additionally, a planned second run
1957 of ProtoDUNE-SP will provide even more data for this and future measurements.

1958 **10.2 Conclusion**

1959 Presented in this thesis is one of the first measurements of π^+ interactions on Argon using
1960 ProtoDUNE-SP data. This measurement would have been impossible without the large
1961 amount of work undertaken within the DUNE collaboration to construct and commission
1962 this detector, currently the largest single-phase LArTPC to have operated. The rapid data-
1963 taking in the Fall of 2018 was followed by an immense effort to carefully categorize and
1964 calibrate the data and also to produce an accurate simulation of the detector. An exciting
1965 future awaits ProtoDUNE-SP, as additional configurations of the detector and more beam
1966 data are planned in the coming years.

1967 This analysis shows the first ever measurement of π^+ -Ar charge exchange and the
 1968 first measurement of π^+ -Ar absorption in this energy range. The LADS collaboration
 1969 measured[22] π^+ -Ar absorption at an energy range below that shown in this thesis. The
 1970 LADS measurement is compared to the measurement from this analysis as well as to the
 1971 prediction from Geant4¹ in Figure 10.4a, while the this measurement of charge exchange is
 1972 compared to Geant4 in Figure 10.4b. Of interest is the disagreement between the Geant4
 1973 model and the LADS data in the resonance region. A similar analysis using data from
 1974 ProtoDUNE-SP at lower momentum will provide a chance to explore this region further. At
 1975 higher momentum, the results of this analysis show close agreement with the Geant4 model
 1976 for both channels.

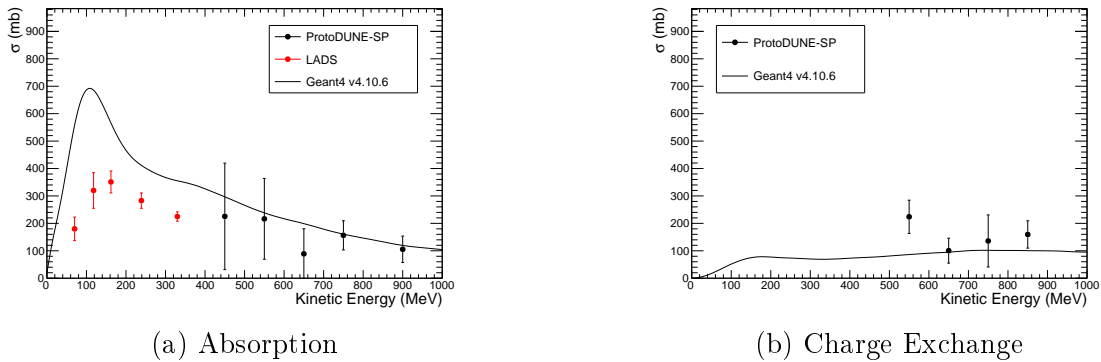


Figure 10.4: The measured π^+ -Ar absorption (left) and charge exchange (right) cross sections compared to the nominal Geant4 model. Additionally, the left plot contains an earlier measurement of absorption from the LADS experiment [22].

1977 The fit used to perform this analysis is based on analyses from the T2K experiment to
 1978 measure ν_μ and $\bar{\nu}_\mu$ interactions in their near detector ND280 [60][61][62][63][64]. It includes
 1979 systematic uncertainties due to the detector and signal model. The benefit of doing such
 1980 an analysis is that it produces robust estimates of the uncertainties and correlations on the
 1981 extracted cross sections. This allows the data to be properly compared to interaction models
 1982 used within detector simulations for upcoming experiments including the Short-Baseline

¹These curves were generated by an application within the Geant4Reweight framework. A 150 MeV/c pion momentum threshold was applied in order to match the signal definition for these measurements. The LADS data does not include this threshold.

1983 Neutrino Program at Fermilab and, ultimately, DUNE. By doing this, important systematic
1984 uncertainties regarding the rate of secondary interactions of neutrino interaction products
1985 can be constrained, and will help allow DUNE to achieve the experimental precision required
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